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y Tecnológicas

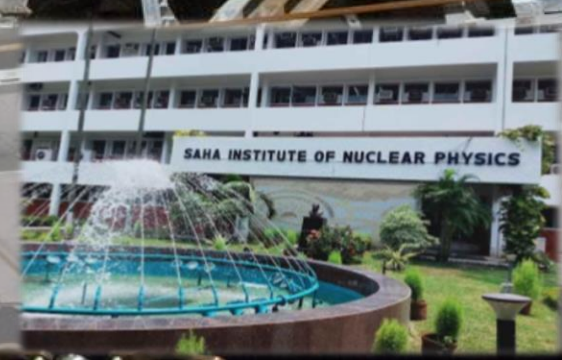


Recent Results from the AMS Experiment

Jorge Casaus
CIEMAT – Madrid
for the AMS Collab.



ICHEPAP2023
**Saha Institute of Nuclear
Physics (SINP, Kolkata)**
11-15 December 2023



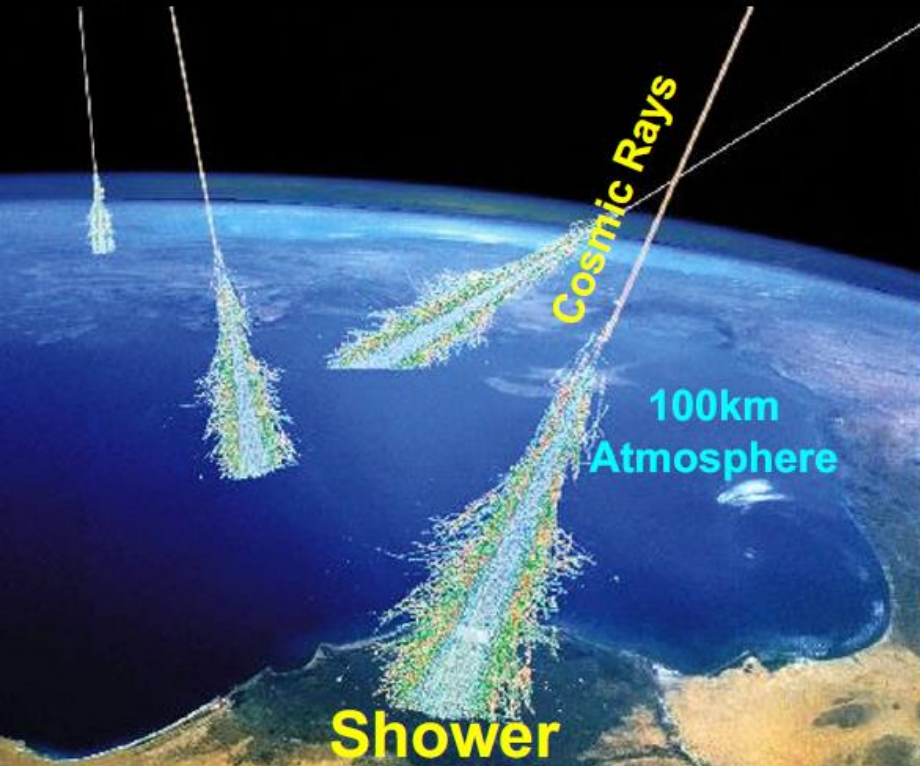
AMS on the Space Station:

Physics of Dark Matter, Antimatter, the Origin of the Cosmos, and new phenomena through the precision, long-duration measurement of charged cosmic rays

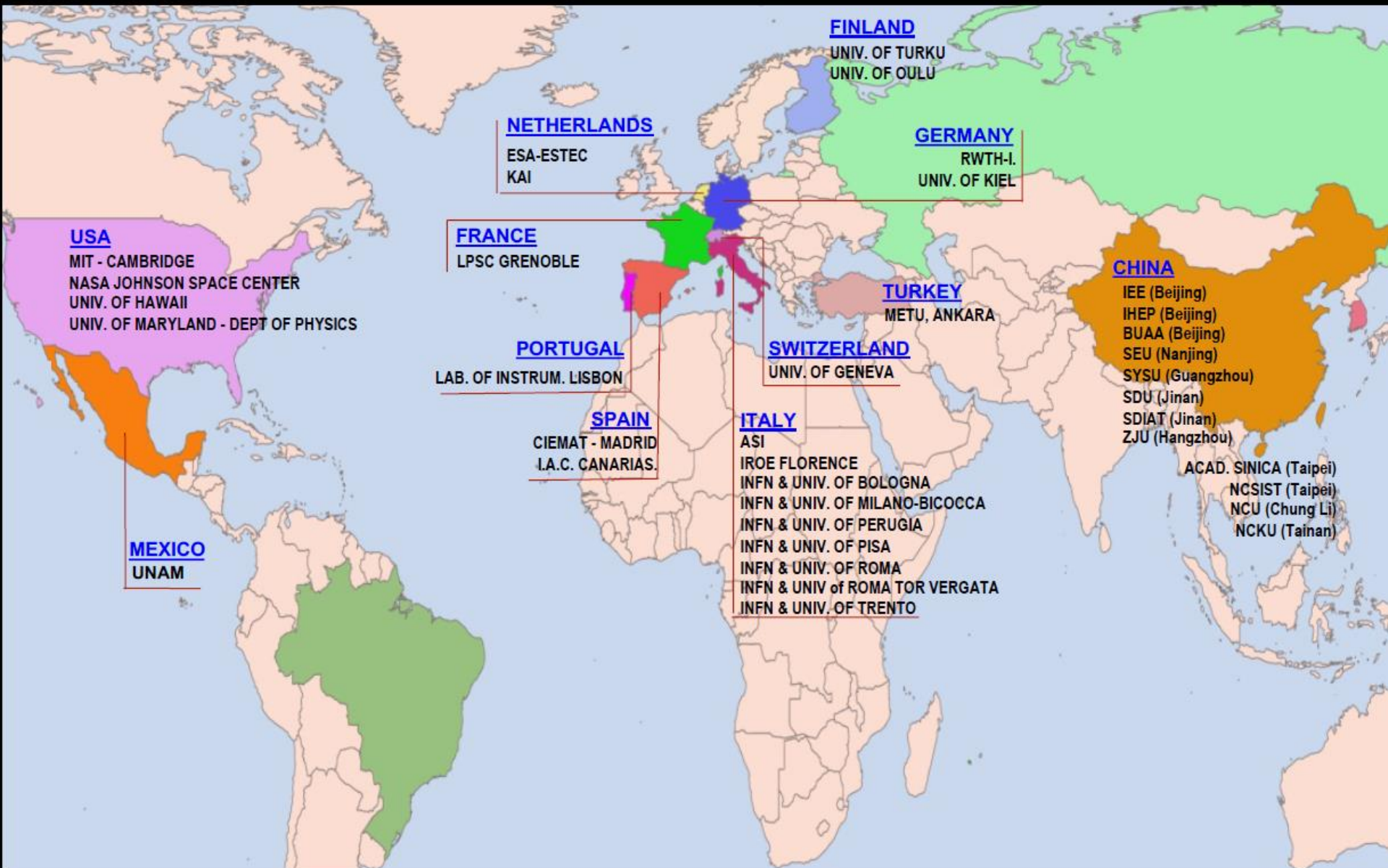
Charged cosmic rays have mass.
They are absorbed by the
100 km of Earth's atmosphere
(10m of water).

To measure cosmic ray
charge and momentum
requires
a magnetic spectrometer
in space

The properties ($\pm Z, P$) of charged cosmic rays cannot be studied on the ground.

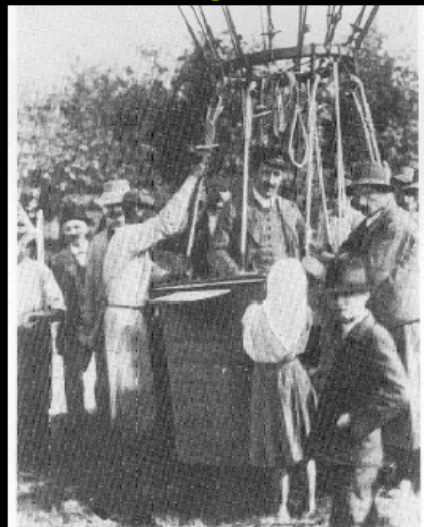


Alpha Magnetic Spectrometer experiment (AMS) on the Space Station



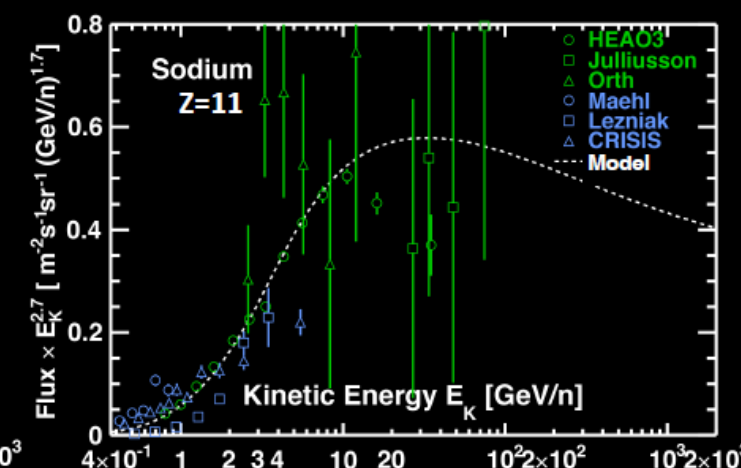
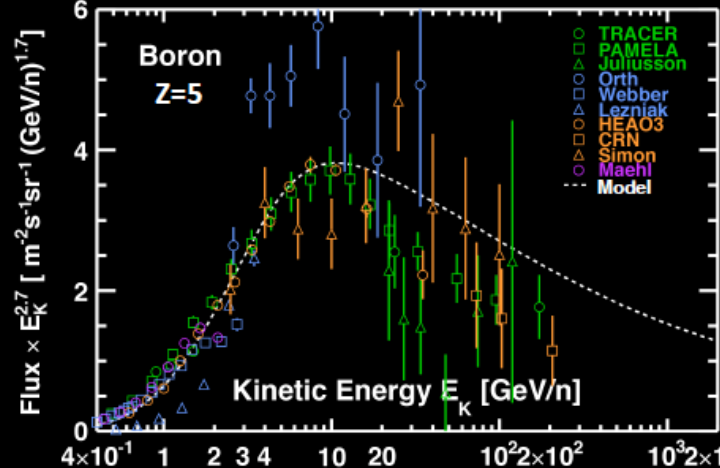
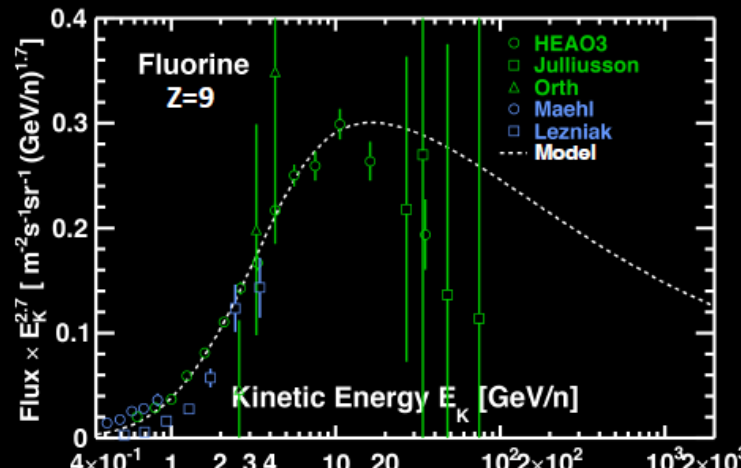
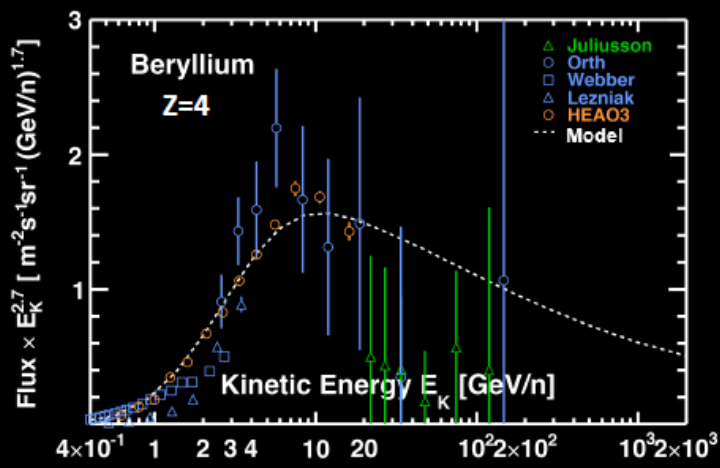
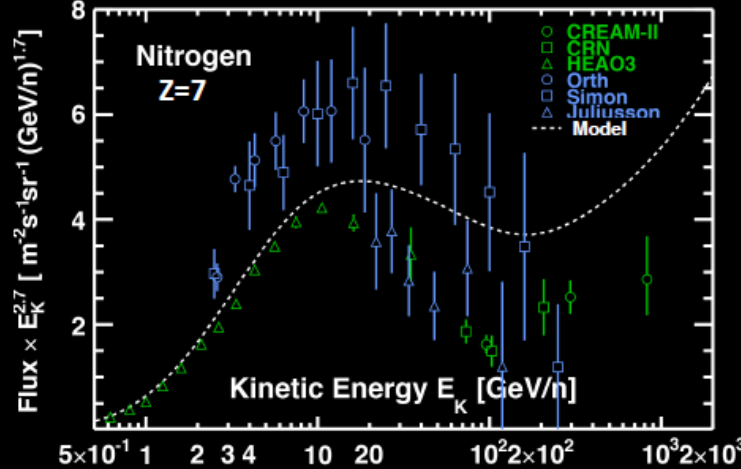
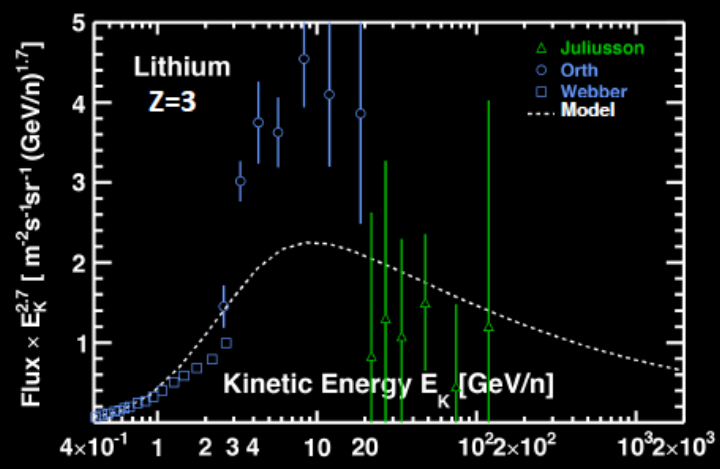
Cosmic Rays in the last 100 years

**1912:
Discovery of Cosmic
Rays**



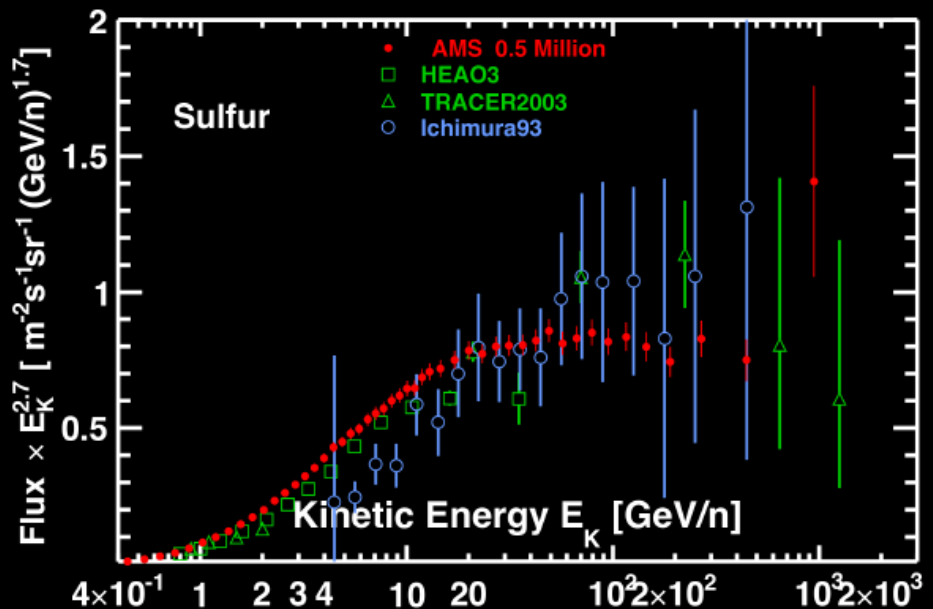
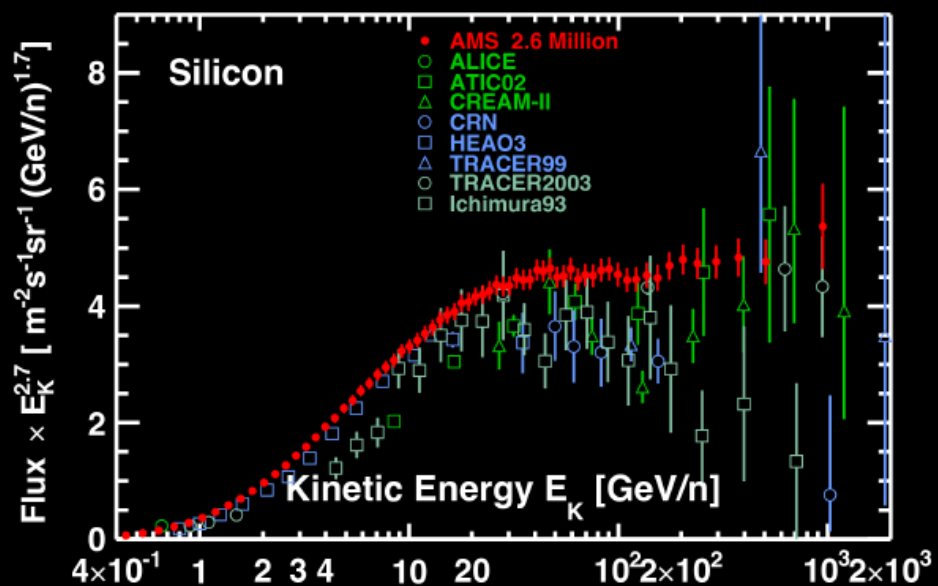
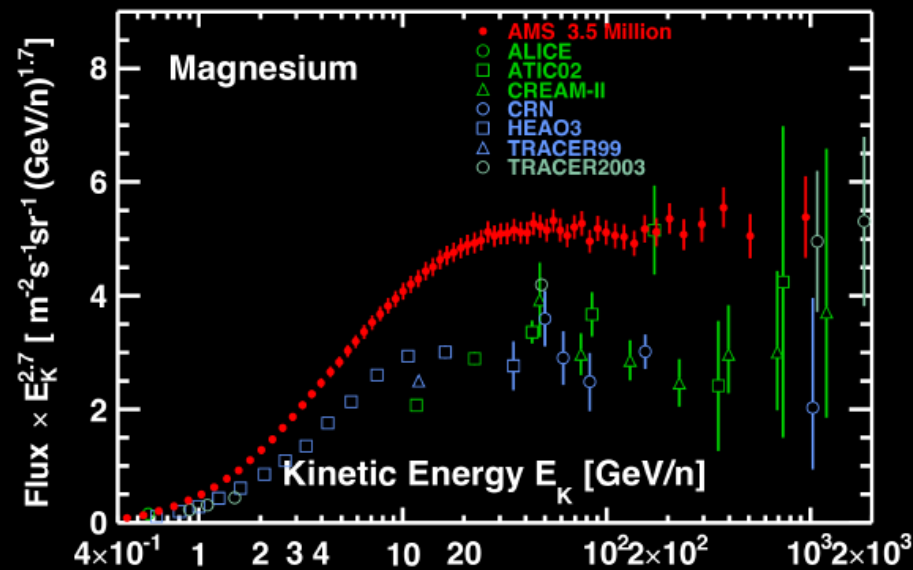
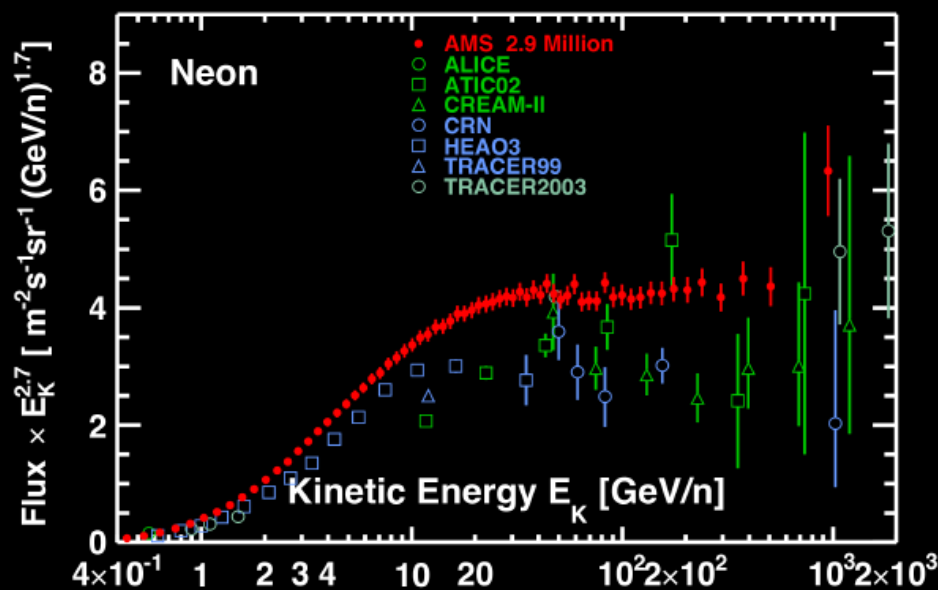
**Victor Hess
Nobel Prize (1936)**

**Before AMS:
Theoretical models
agree with
experimental data
(large errors)**



Examples of AMS Results compared with earlier measurements

The precision AMS results cannot be explained by current models.



AMS is a space version of a precision detector used in accelerators

Transition Radiation Detector (TRD)

identify e^+ , e^-



Upper TOF measure Z, E



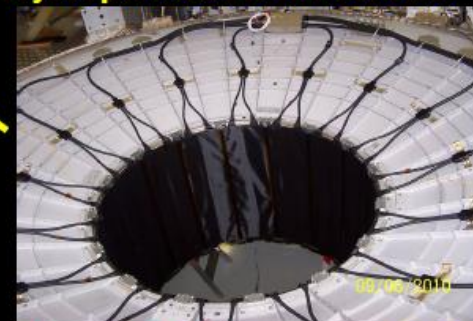
Magnet identify $\pm Z, P$



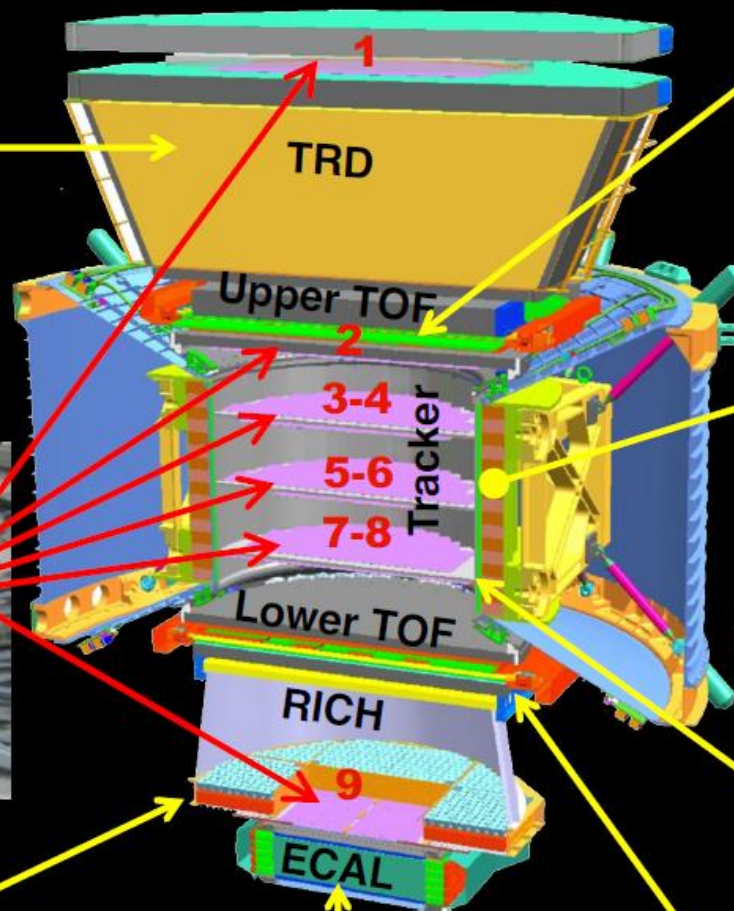
Silicon Tracker
measure Z, P



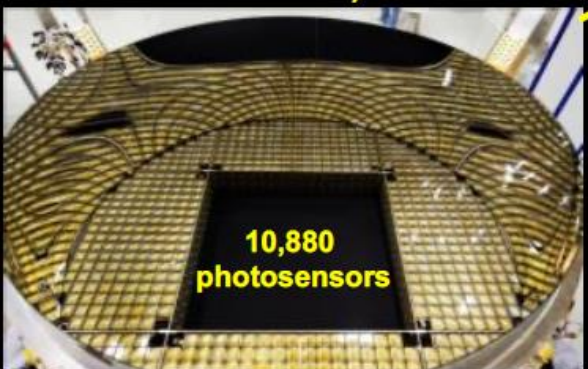
Anticoincidence Counters (ACC)
reject particles from the side



Lower TOF measure Z, E



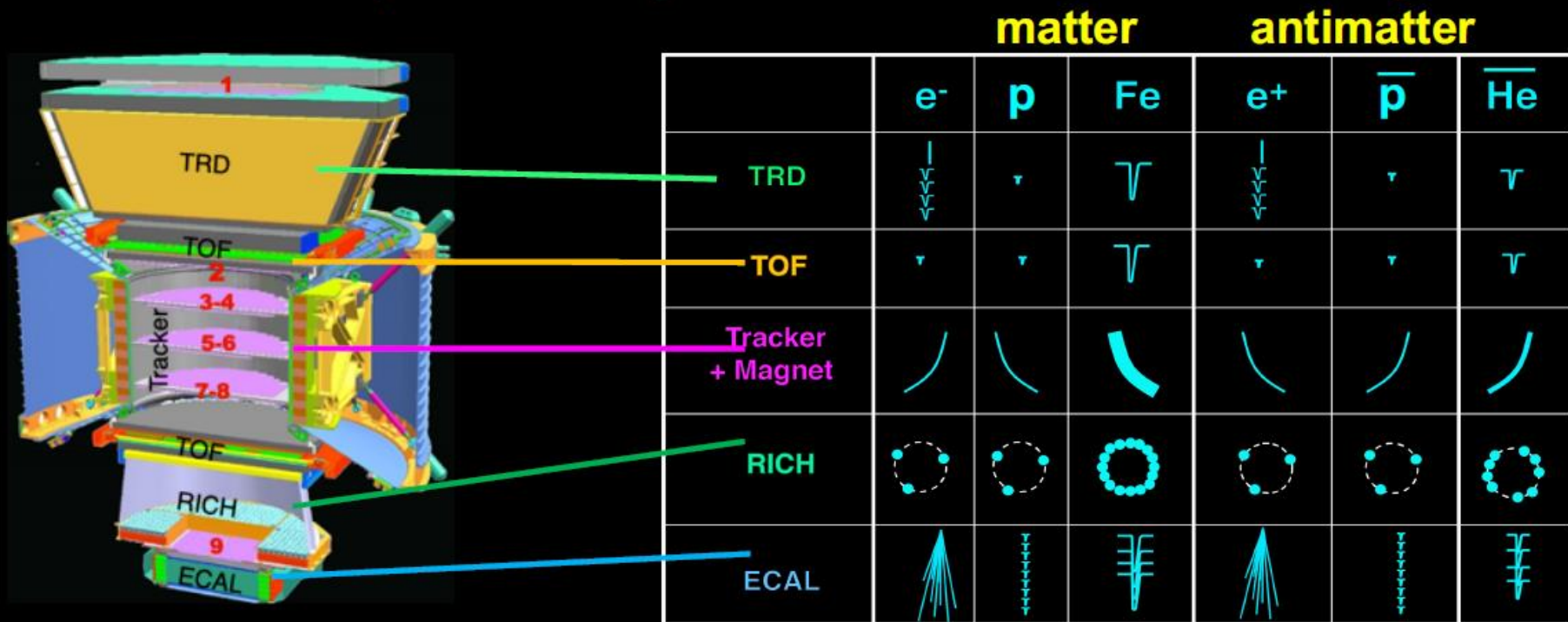
Ring Imaging Cerenkov (RICH)
measure Z, E



Electromagnetic Calorimeter (ECAL)
measure E of e^+ , e^-



The detectors provide independent information of cosmic rays



Redundant measurements of:

- Momentum (P , GeV/c)
- Charge (Z)
- Rigidity ($R=P/Z$, GV)
- Energy (E , GeV/A)
- Flux (particles/(s sr m² GeV))

Periodic Table of the Elements

1 H Hydrogen 1.008	2 He Helium 4.003	3 Li Lithium 6.941	4 Be Beryllium 9.012	5 B Boron 10.811	6 C Carbon 12.011	7 N Nitrogen 14.007	8 O Oxygen 15.999	9 F Fluorine 18.998	10 Ne Neon 20.180	11 Na Sodium 22.990	12 Mg Magnesium 24.305	13 Al Aluminum 26.982	14 Si Silicon 28.086	15 P Phosphorus 30.974	16 S Sulfur 32.065	17 Cl Chlorine 35.453	18 Ar Argon 39.948	19 K Potassium 39.098	20 Ca Calcium 40.078	21 Sc Scandium 44.956	22 Ti Titanium 47.88	23 V Vanadium 50.942	24 Cr Chromium 51.996	25 Mn Manganese 54.938	26 Fe Iron 55.933	27 Co Cobalt 58.933	28 Ni Nickel 58.693	29 Cu Copper 63.546	30 Zn Zinc 65.39	31 Ga Gallium 69.723	32 Ge Germanium 72.61	33 As Arsenic 74.922	34 Se Selenium 78.96	35 Br Bromine 79.904	36 Kr Krypton 84.961	37 Rb Rubidium 85.468	38 Sr Strontium 87.62	39 Y Yttrium 88.906	40 Zr Zirconium 91.224	41 Nb Niobium 92.906	42 Mo Molybdenum 95.94	43 Tc Technetium 98.907	44 Ru Ruthenium 101.07	45 Rh Rhodium 102.906	46 Pd Palladium 106.42	47 Ag Silver 107.868	48 Cd Cadmium 112.411	49 In Indium 114.818	50 Sn Tin 118.71	51 Sb Antimony 121.760	52 Te Tellurium 127.6	53 I Iodine 126.905	54 Xe Xenon 131.29	55 Cs Cesium 132.905	56 Ba Barium 137.327	57-71 Lanthanides	72 Hf Hafnium 178.49	73 Ta Tantalum 180.948	74 W Tungsten 183.85	75 Re Rhenium 186.207	76 Os Osmium 190.23	77 Ir Iridium 192.22	78 Pt Platinum 195.08	79 Au Gold 196.967	80 Hg Mercury 200.59	81 Tl Thallium 204.383	82 Pb Lead 207.2	83 Bi Bismuth 208.980	84 Po Polonium [209]	85 At Astatine [209]	86 Rn Radon [222]	87 Fr Francium [223]	88 Ra Radium [226]	89-103 Actinides	104 Rf Rutherfordium [261]	105 Db Dubnium [262]	106 Sg Seaborgium [266]	107 Bh Bohrium [264]	108 Hs Hassium [269]	109 Mt Meitnerium [268]	110 Ds Darmstadtium [289]	111 Rg Roentgenium [272]	112 Cn Copernicium [277]	113 Nh Nihonium [278]	114 Fl Flerovium [289]	115 Uup Ununpentium [288]	116 Lv Livermorium [293]	117 Uus Ununseptium [294]	118 Uuo Ununoctium [294]
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AMS on the ISS

AMS 2011-2025

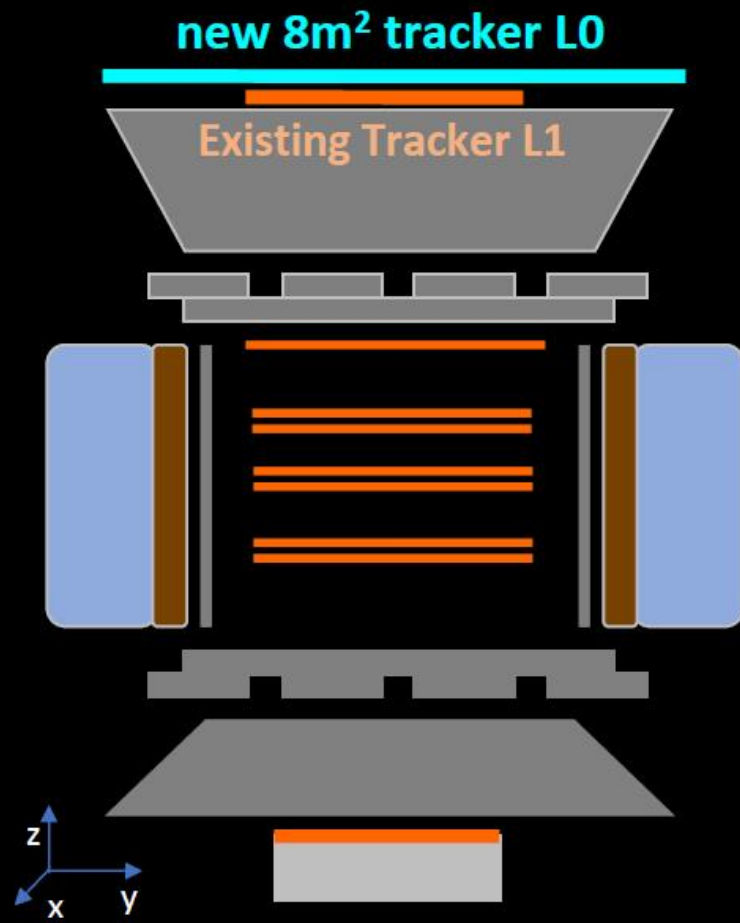
Continuous data taking



Latest results 2011 - 2022: 220 billion cosmic rays

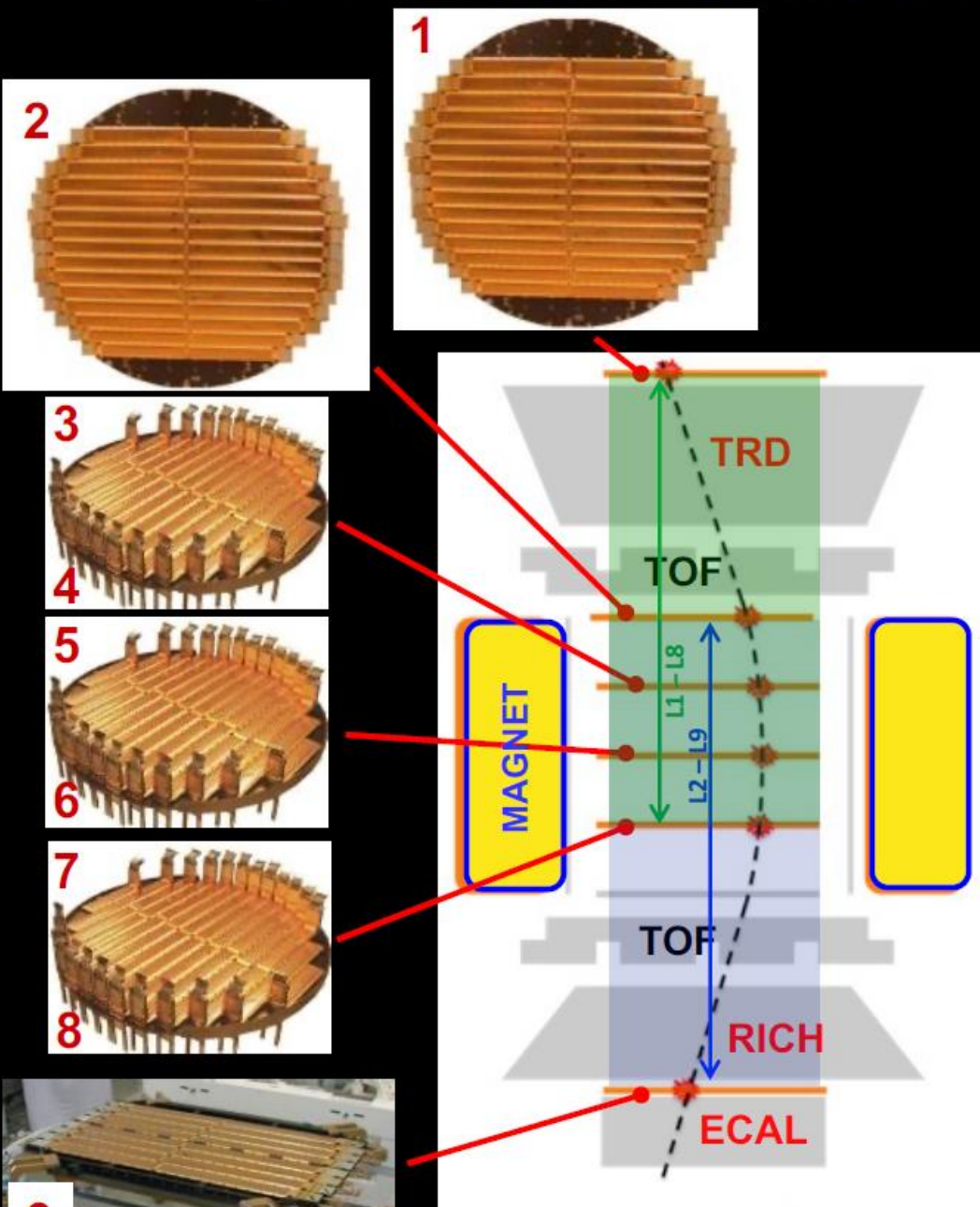
AMS 2025-2030

Acceptance increased to 300%



and projections

Operation of AMS on the ISS: Continuous Calibration

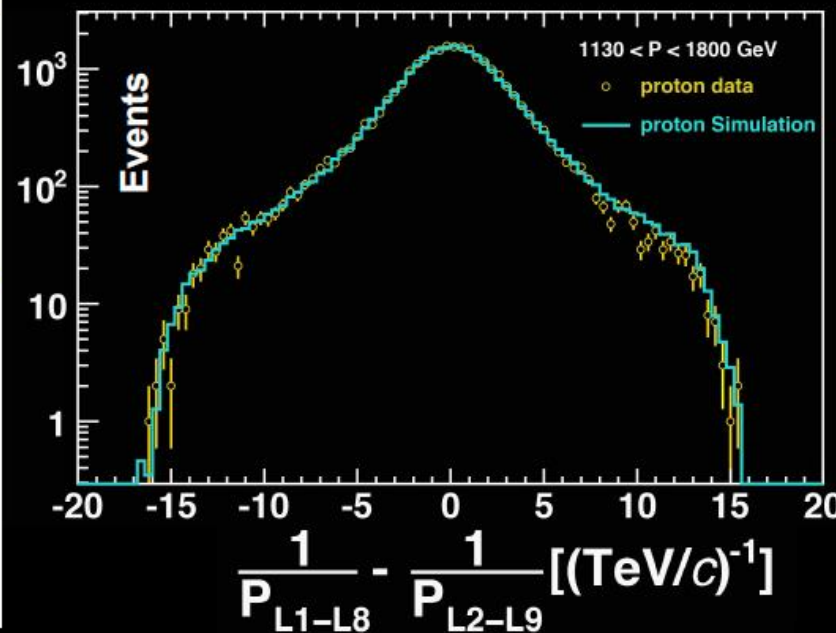


By comparing proton data and simulation from

the upper spectrometer (L1 to L8)

and

the lower spectrometer (L2 to L9)

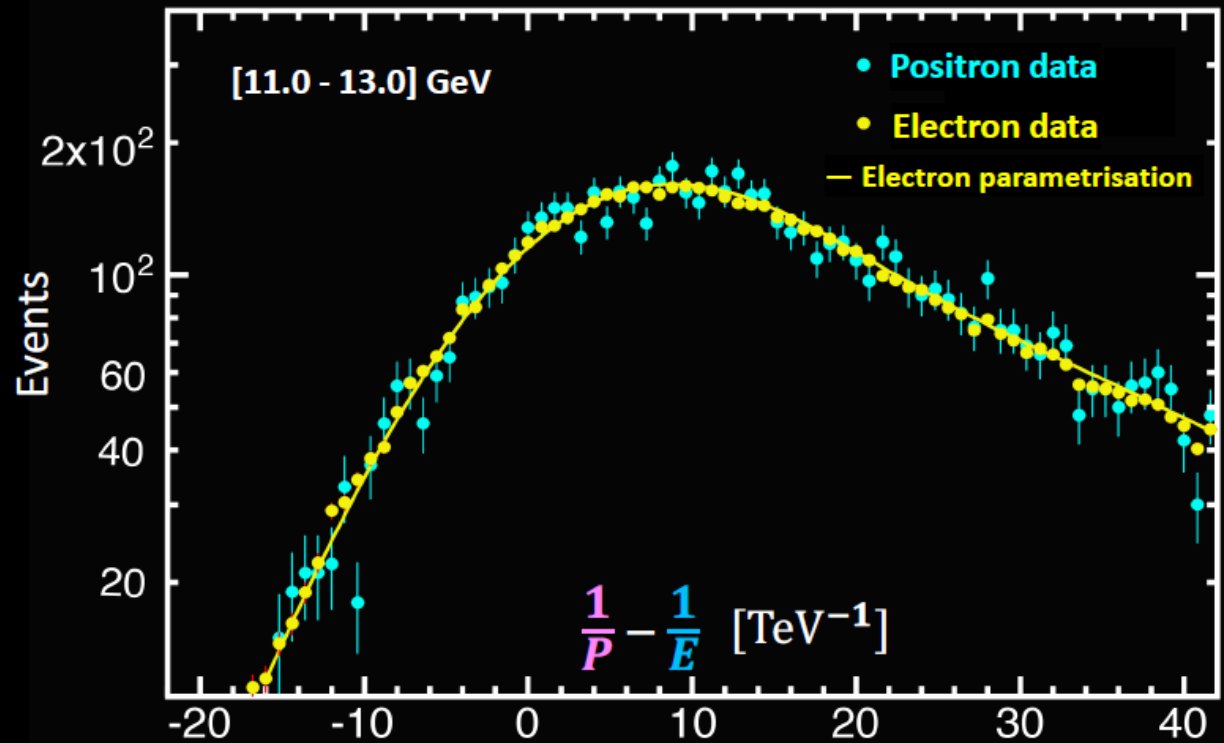
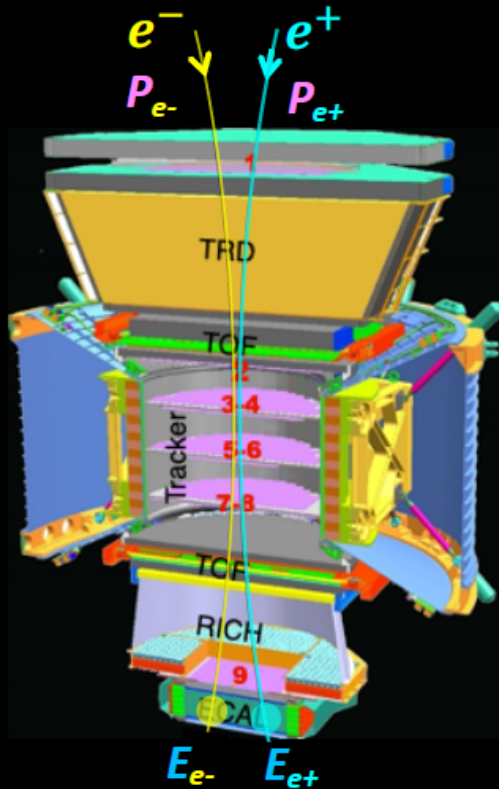


Continuous Momentum Scale Verification

(Unique Advantage of a Magnetic Spectrometer)

In AMS, the largest systematic error in the determination of the fluxes at the highest energies is due to the uncertainty in the absolute momentum scale.

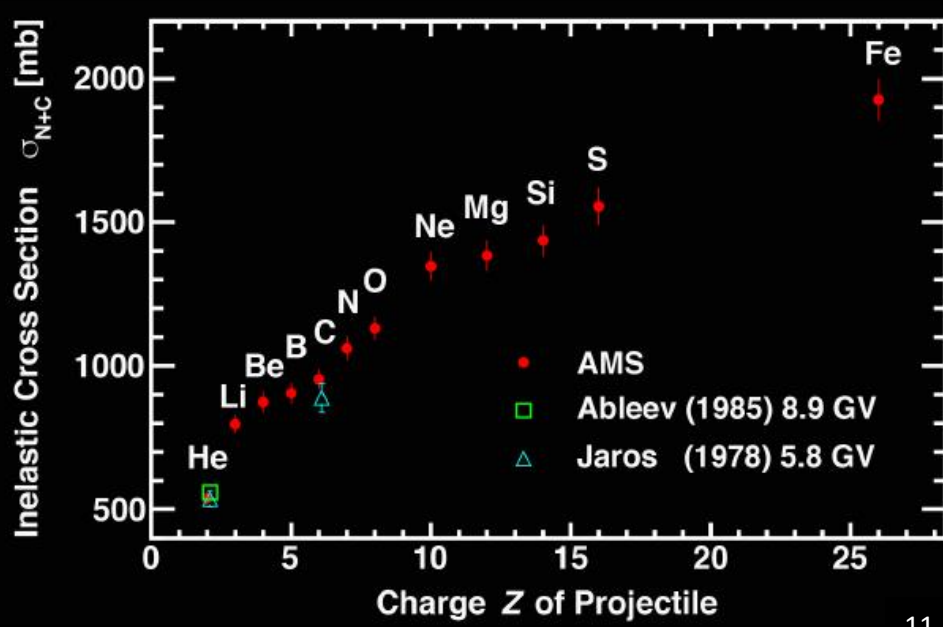
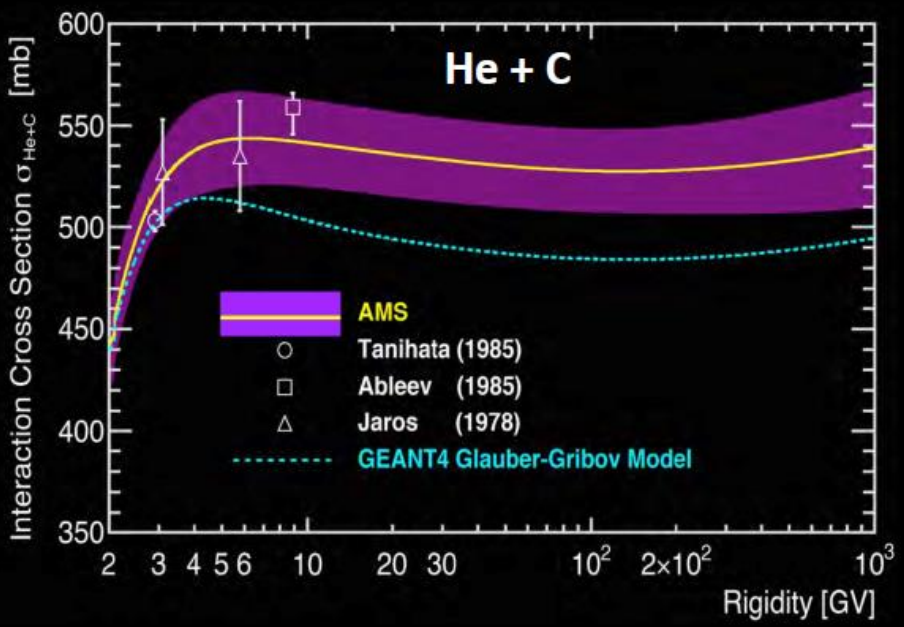
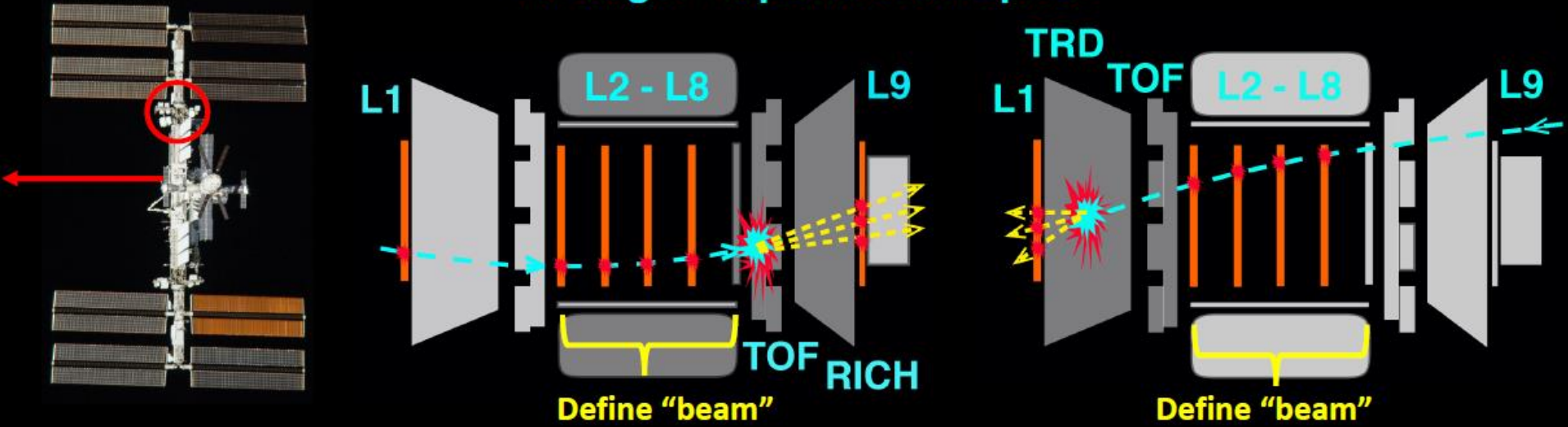
A shift in the central tracker planes of 0.5 microns is sufficient to create a momentum shift of 10% at 1 TeV



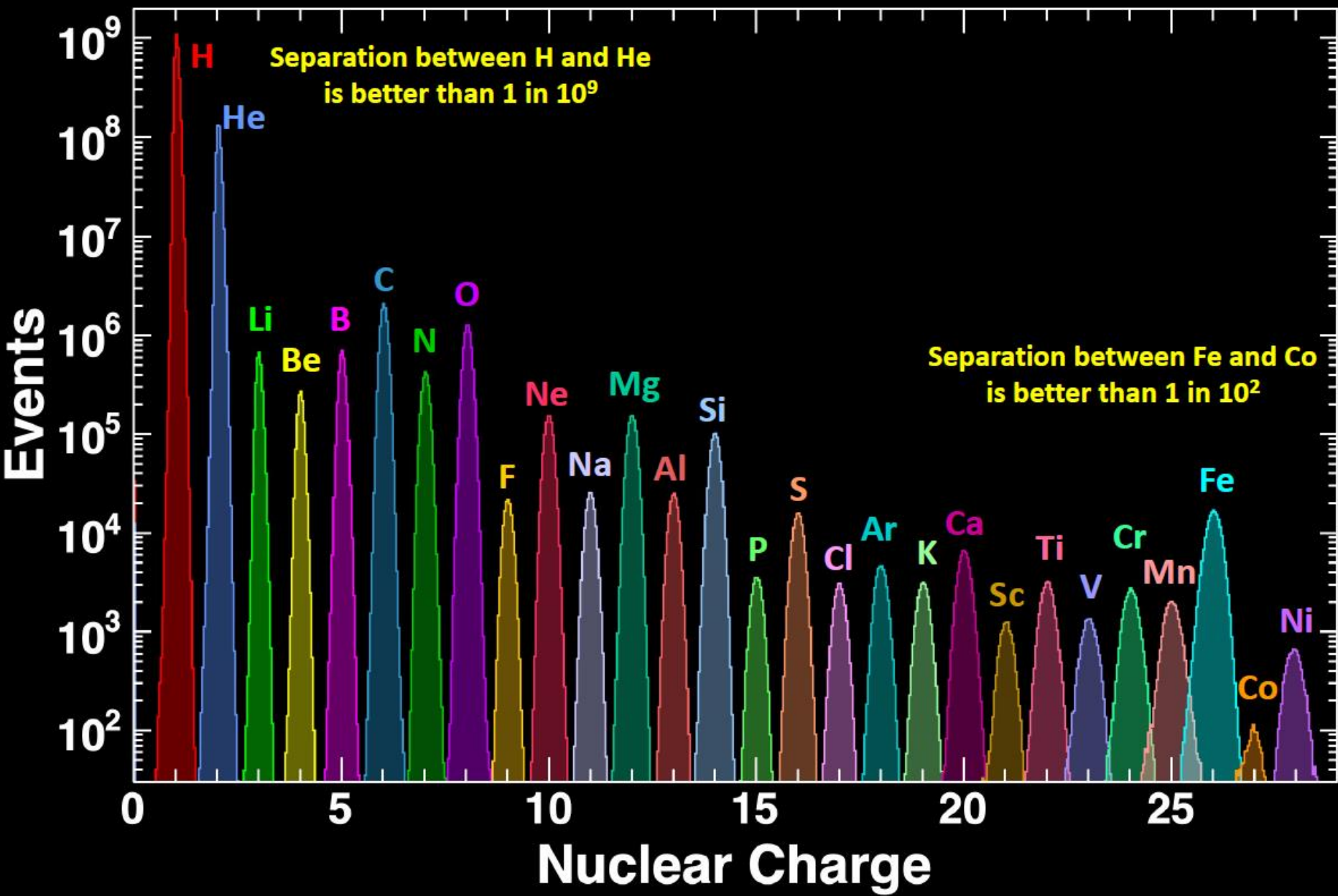
By matching the momentum P and Energy E for both e^+ and e^- ,
The accuracy of the momentum is determined to be $1/(34 \text{ TeV})$;
i.e., at 1 TeV the uncertainty is 3%

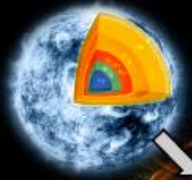
Precision measurements of nuclear inelastic cross sections

“Fix target” experiment in space

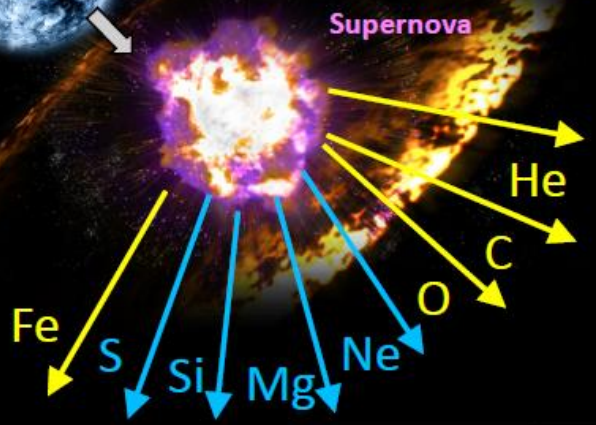


Precision Measurements of Cosmic Nuclei by AMS



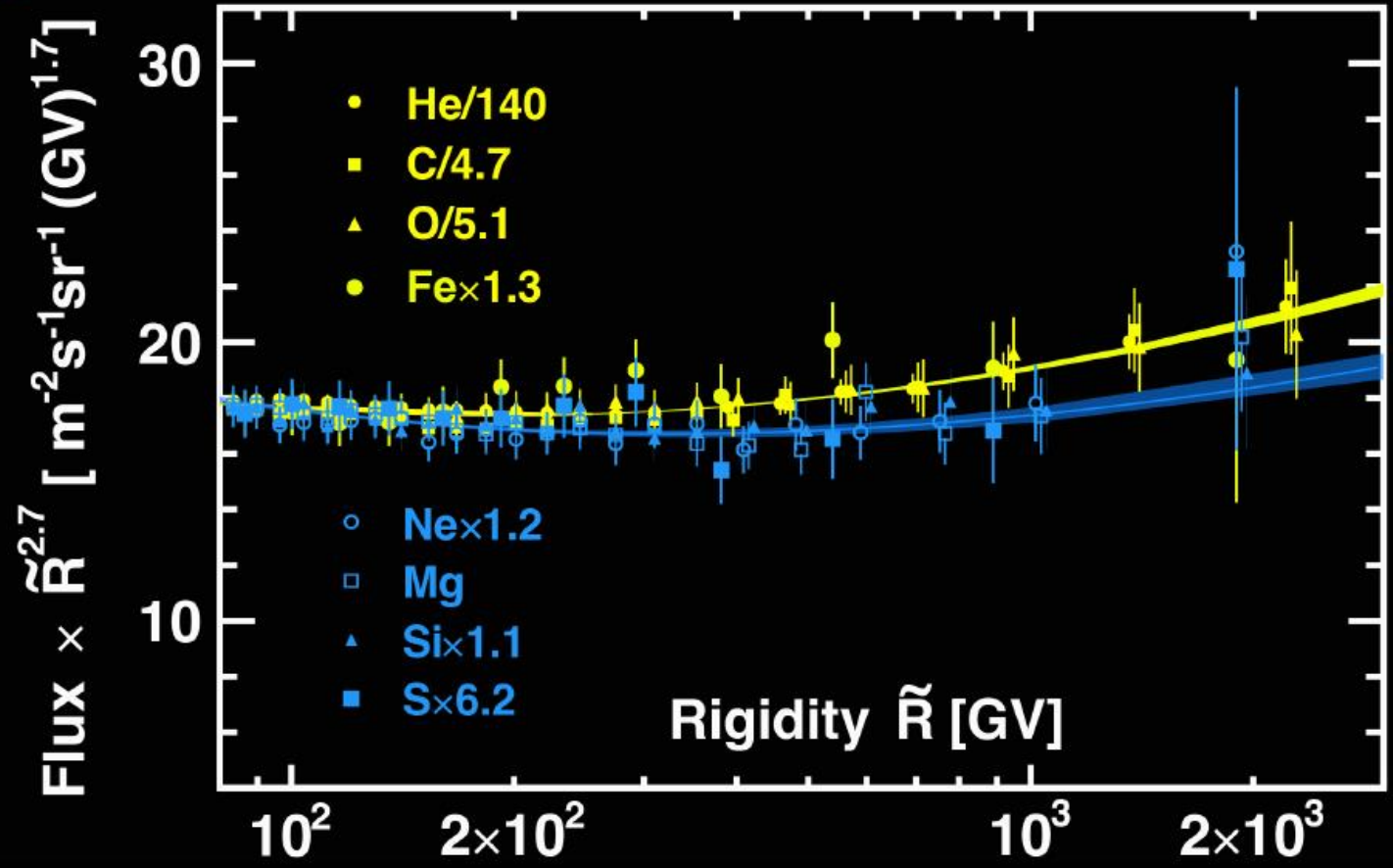


Primary Cosmic Rays

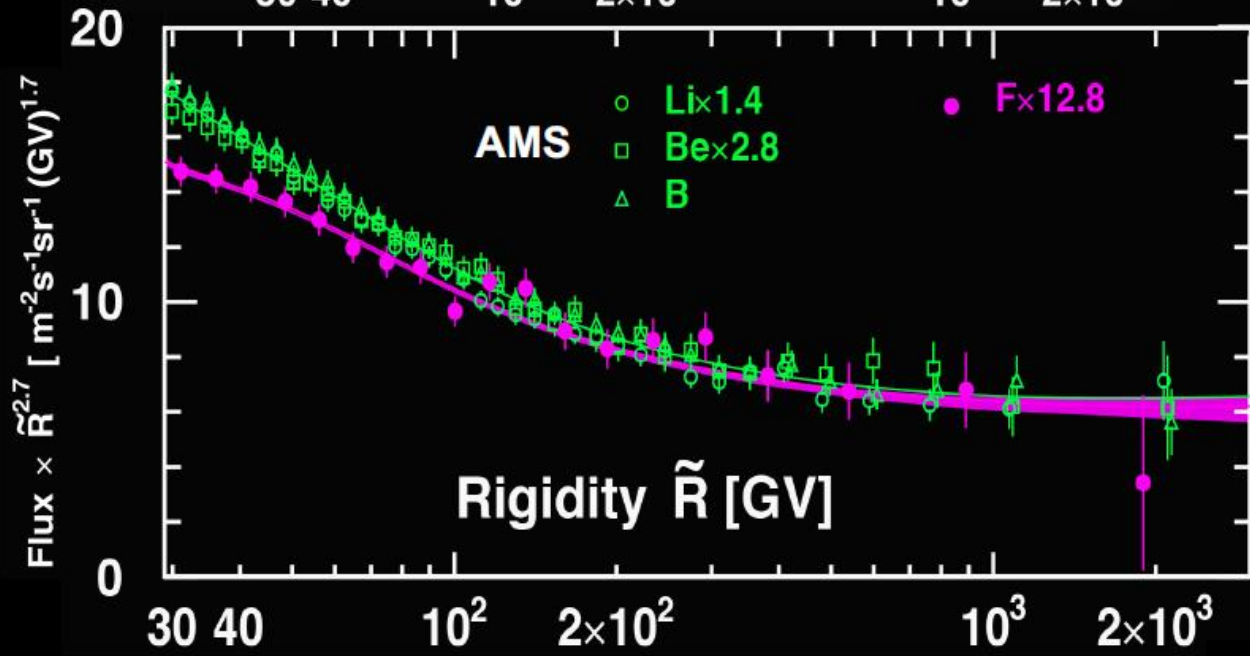
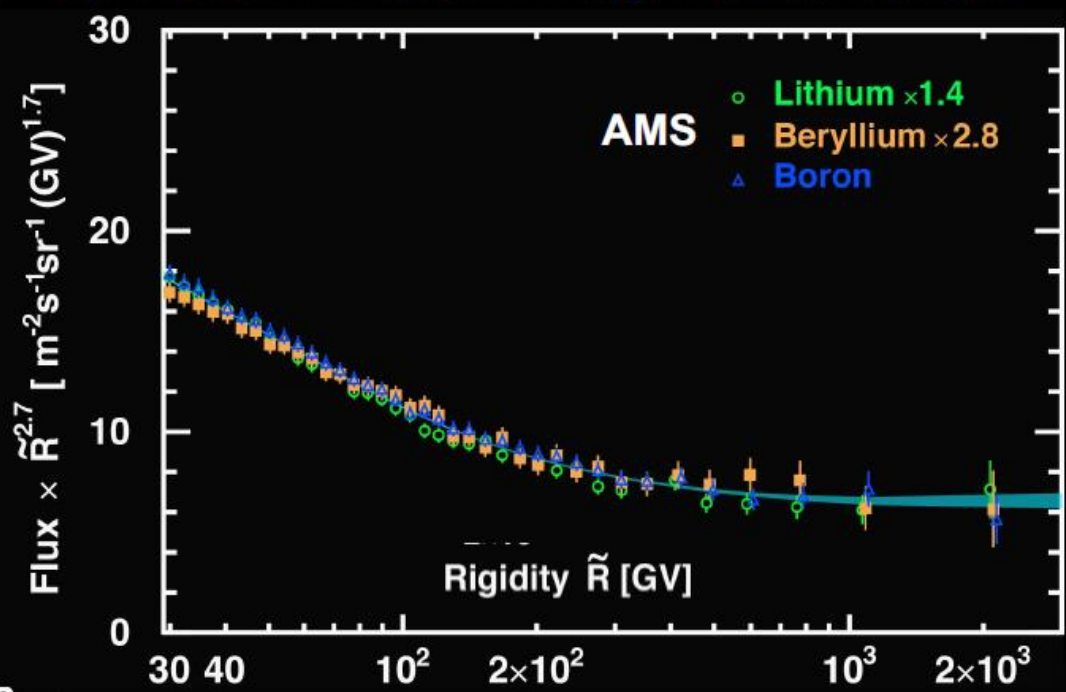
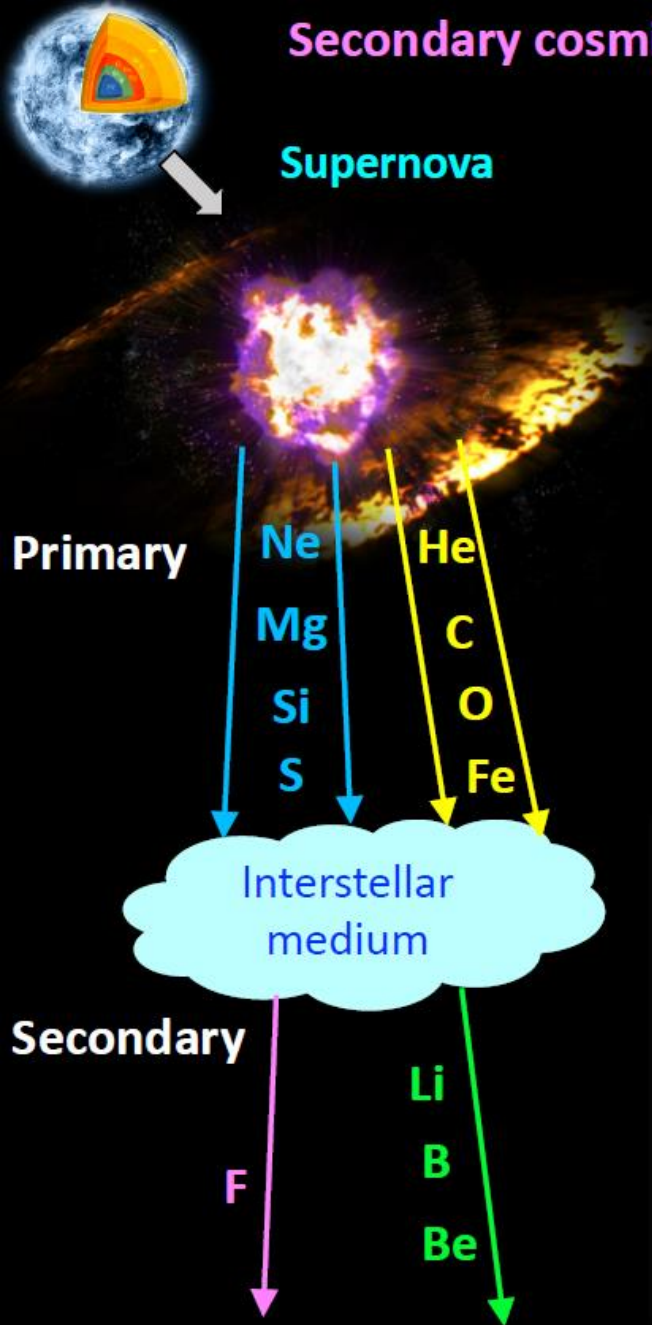


Primary elements (He, C, ..., Fe) are produced during the lifetime of stars.
They are accelerated by the explosion of stars (supernovae).

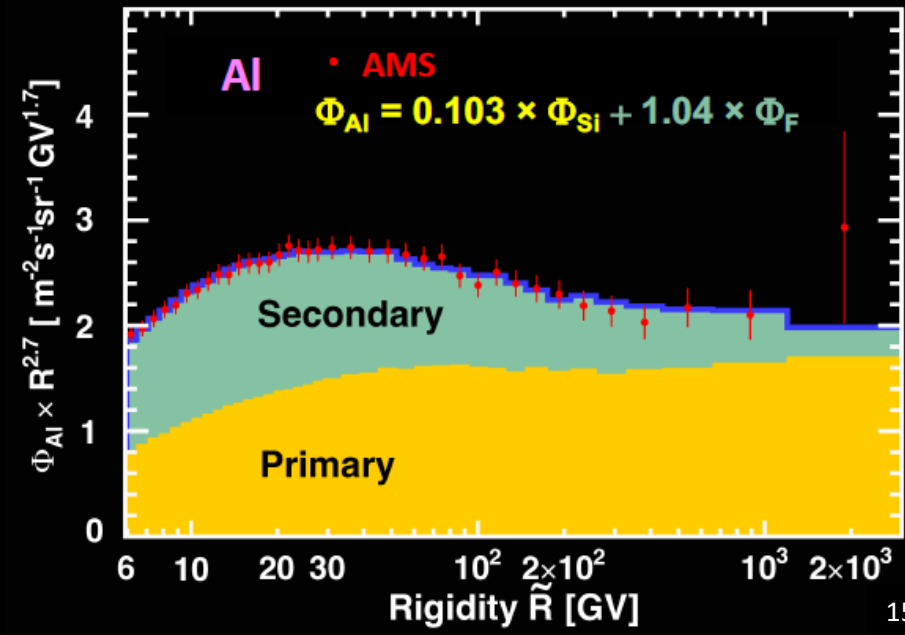
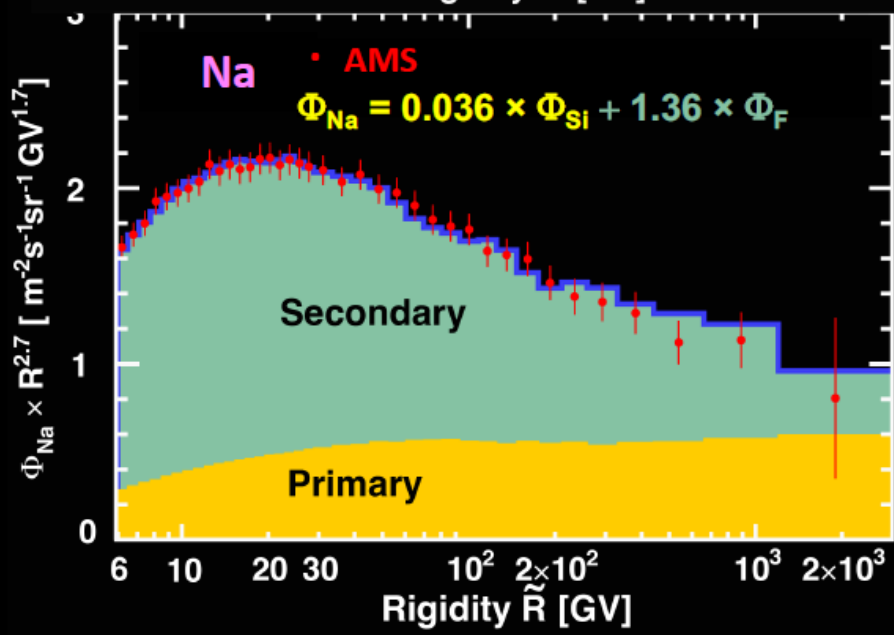
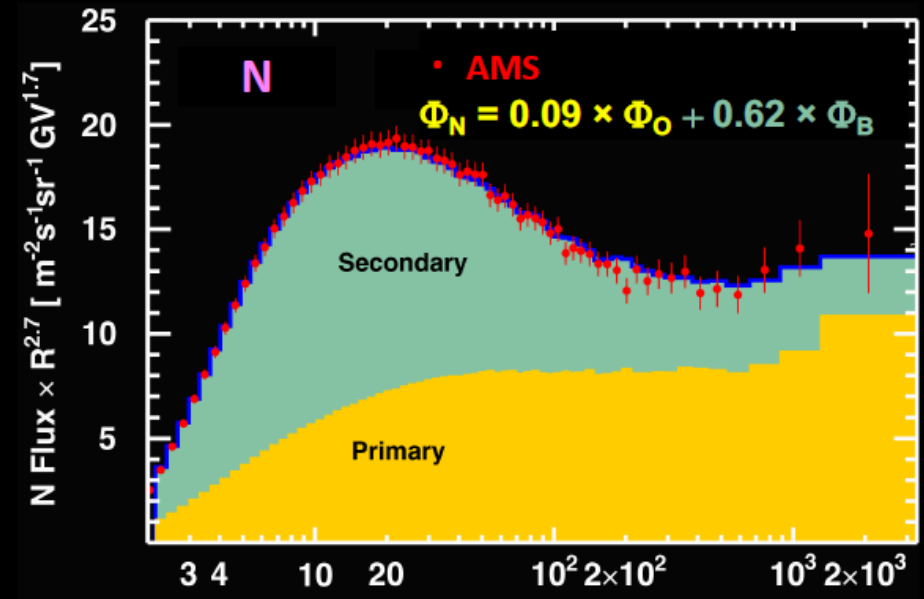
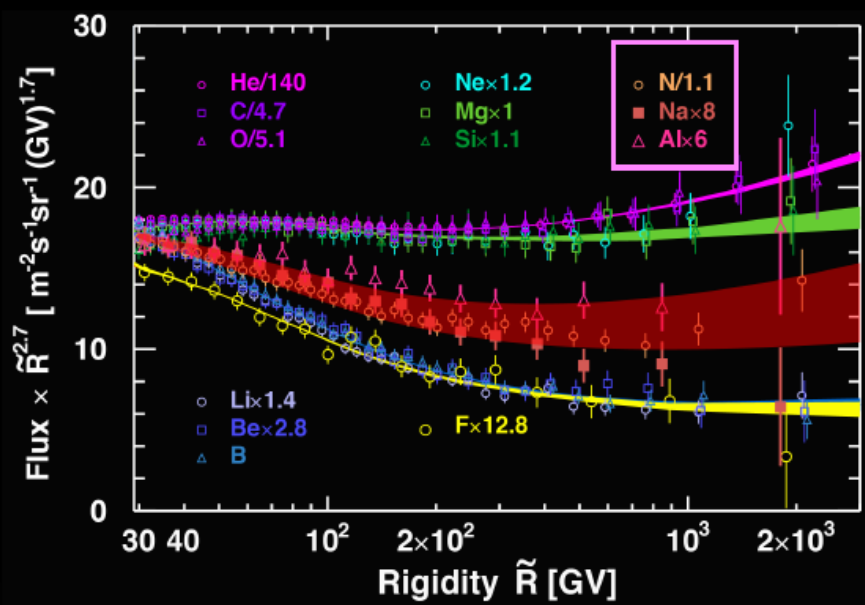
Primary cosmic rays have at least two classes.



Secondary cosmic rays also have two classes of rigidity dependence

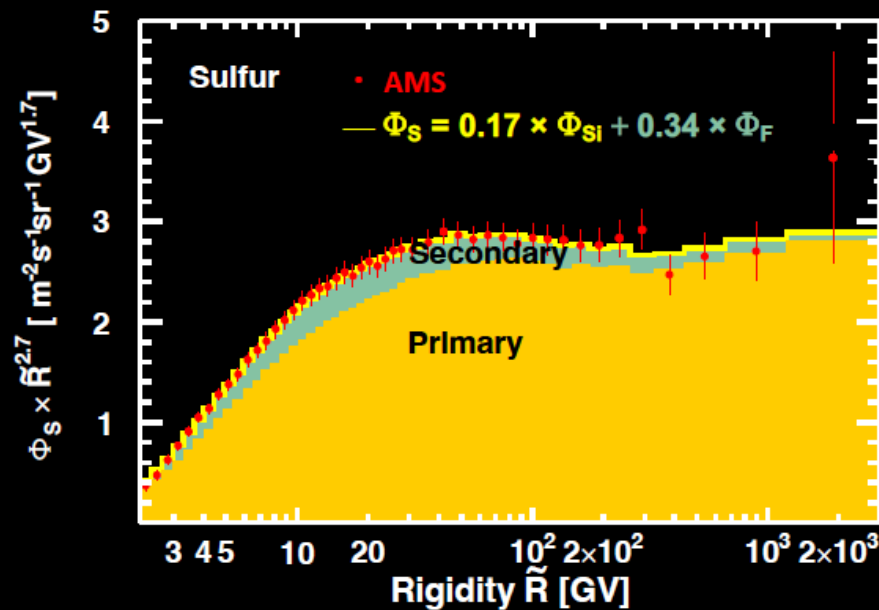
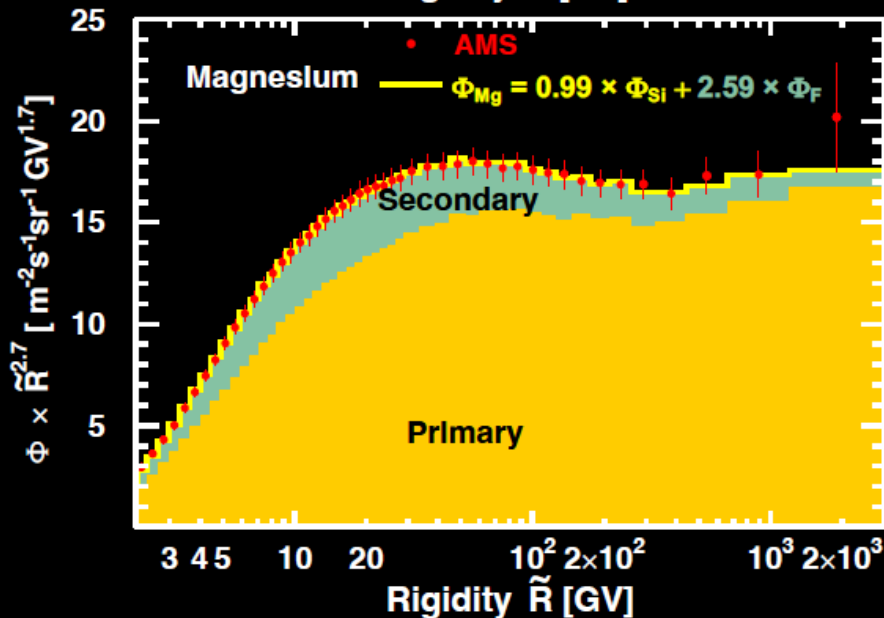
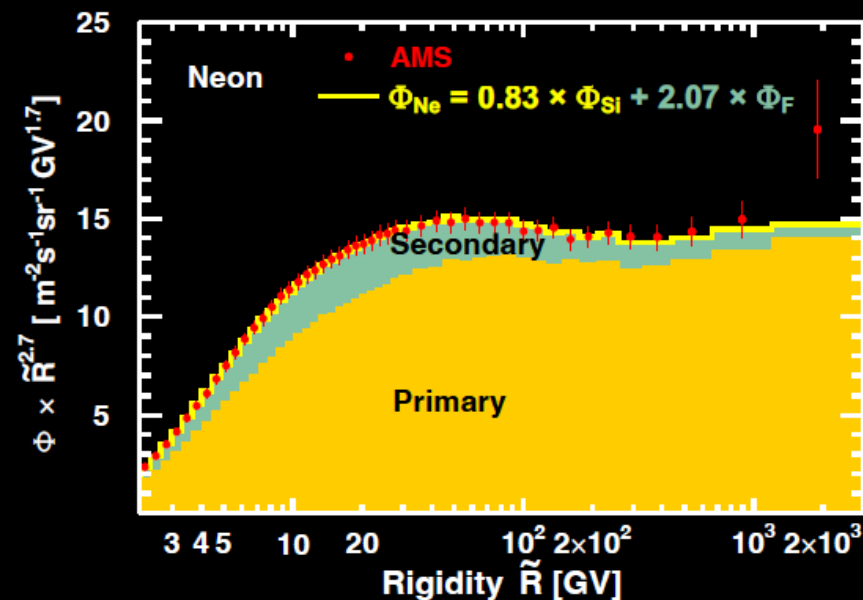
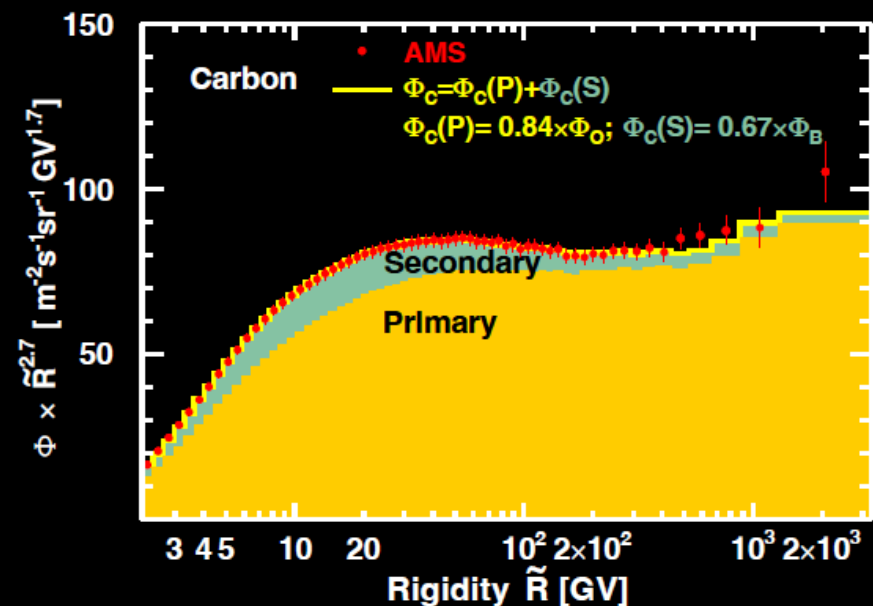


A third group of cosmic rays N-Na-Al partially primary, partially secondary

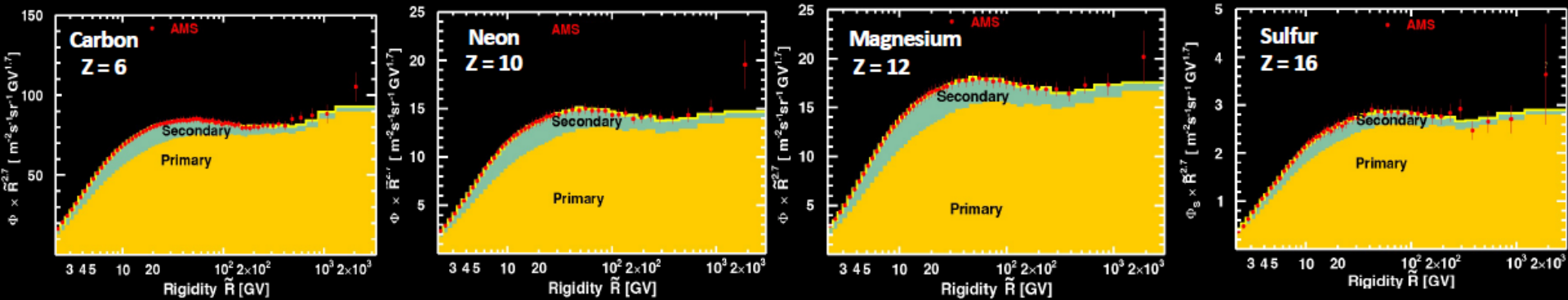


New, unexpected observation :

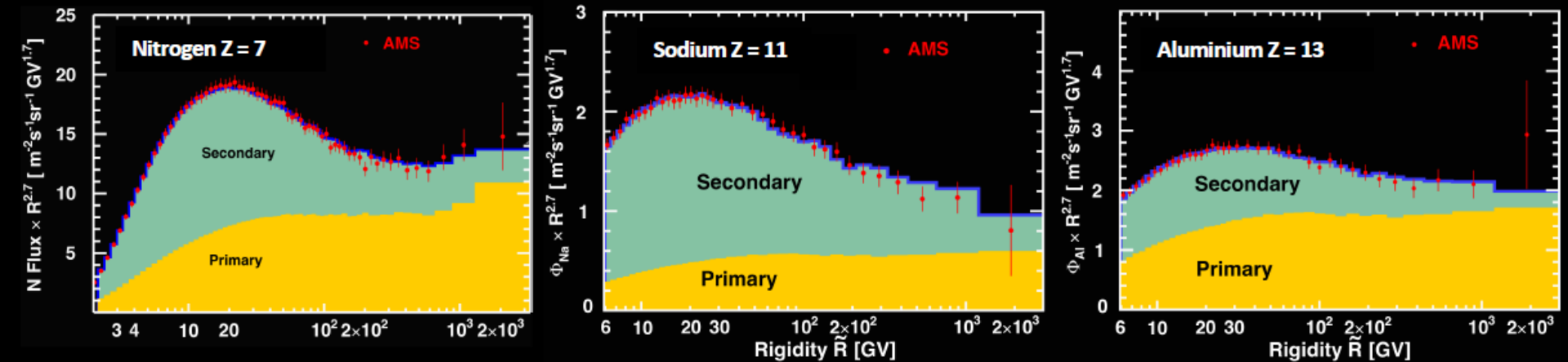
Traditional primary cosmic rays C, Ne, Mg, and S fluxes are not pure primary; they all have a significant secondary component



New Observation: Even-Z nuclei and Odd-Z nuclei have distinctly different primary and secondary composition



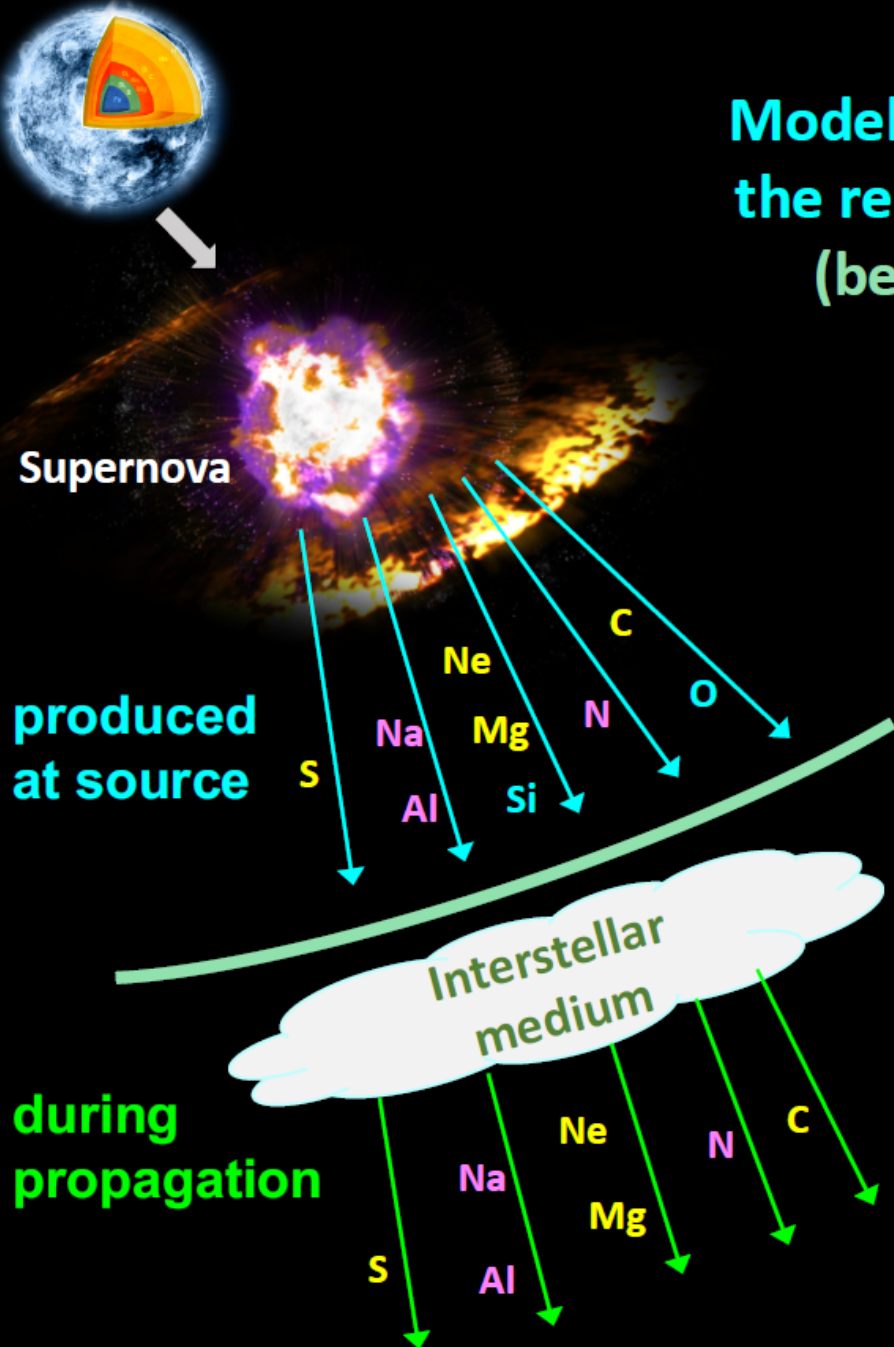
Even-Z nuclei are dominated by primaries



Odd-Z nuclei have more secondaries than even-Z

New AMS Result:

Model-independent measurements of the relative abundances at the source (before cosmic ray propagation)

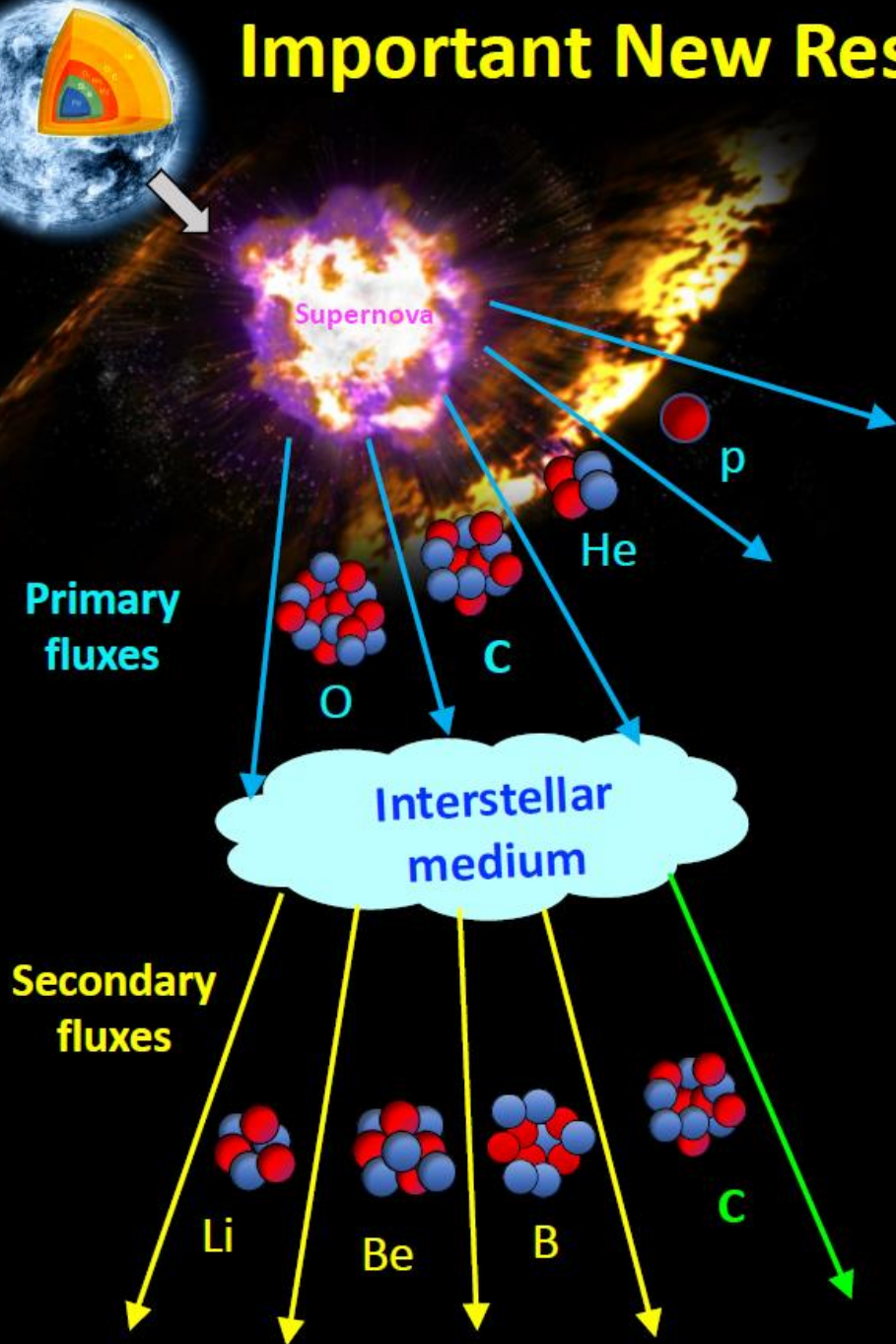


Abundance Ratio

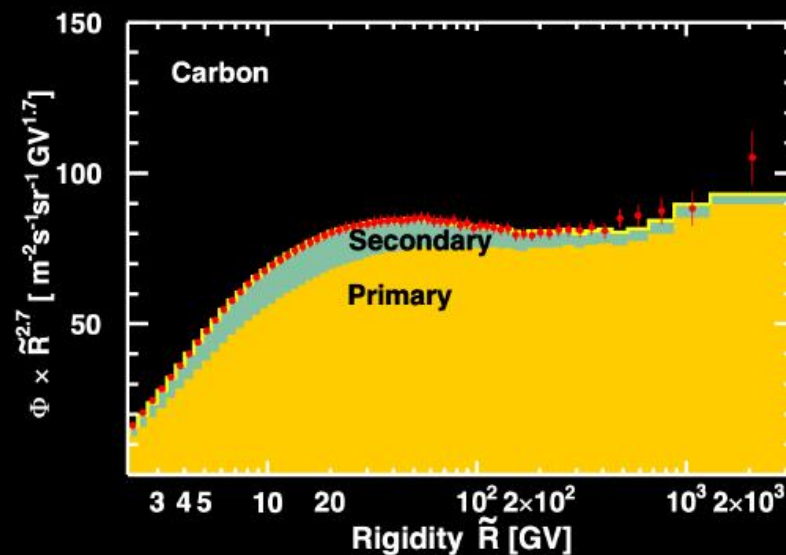
Value at the Source

C/O	0.836 ± 0.025
Ne/Si	0.833 ± 0.025
Mg/Si	0.994 ± 0.029
S/Si	0.167 ± 0.006
N/O	0.092 ± 0.002
Na/Si	0.036 ± 0.003
Al/Si	0.103 ± 0.004

Important New Result and its Implications



Before AMS, taking into account the long-standing idea that **C is pure primary** and **B is pure secondary**, the **(B/C) ratio** has been used in all the theories

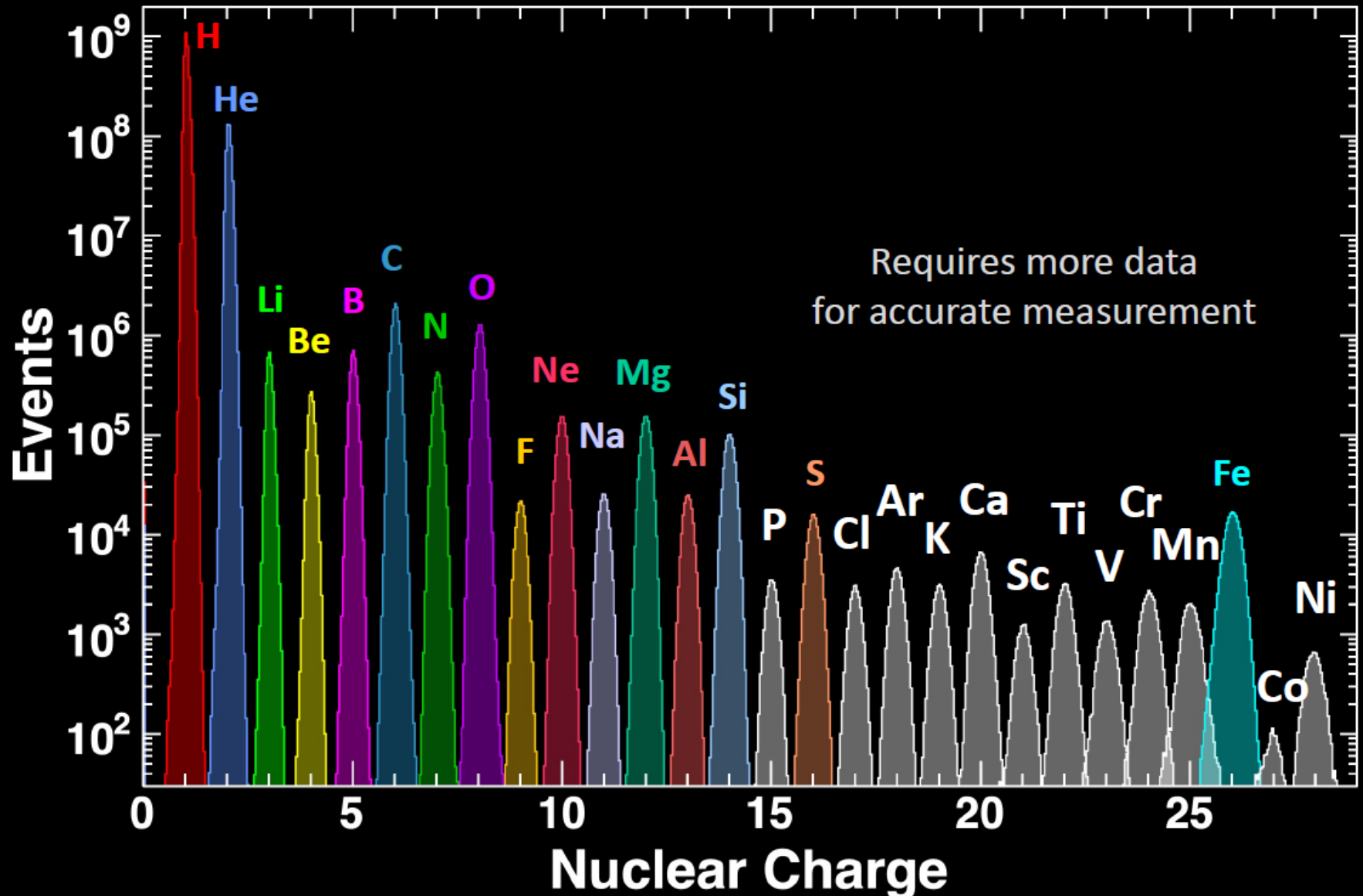


But **C is NOT pure primary**.
Question: how to use **(B/C)** in cosmic ray models?

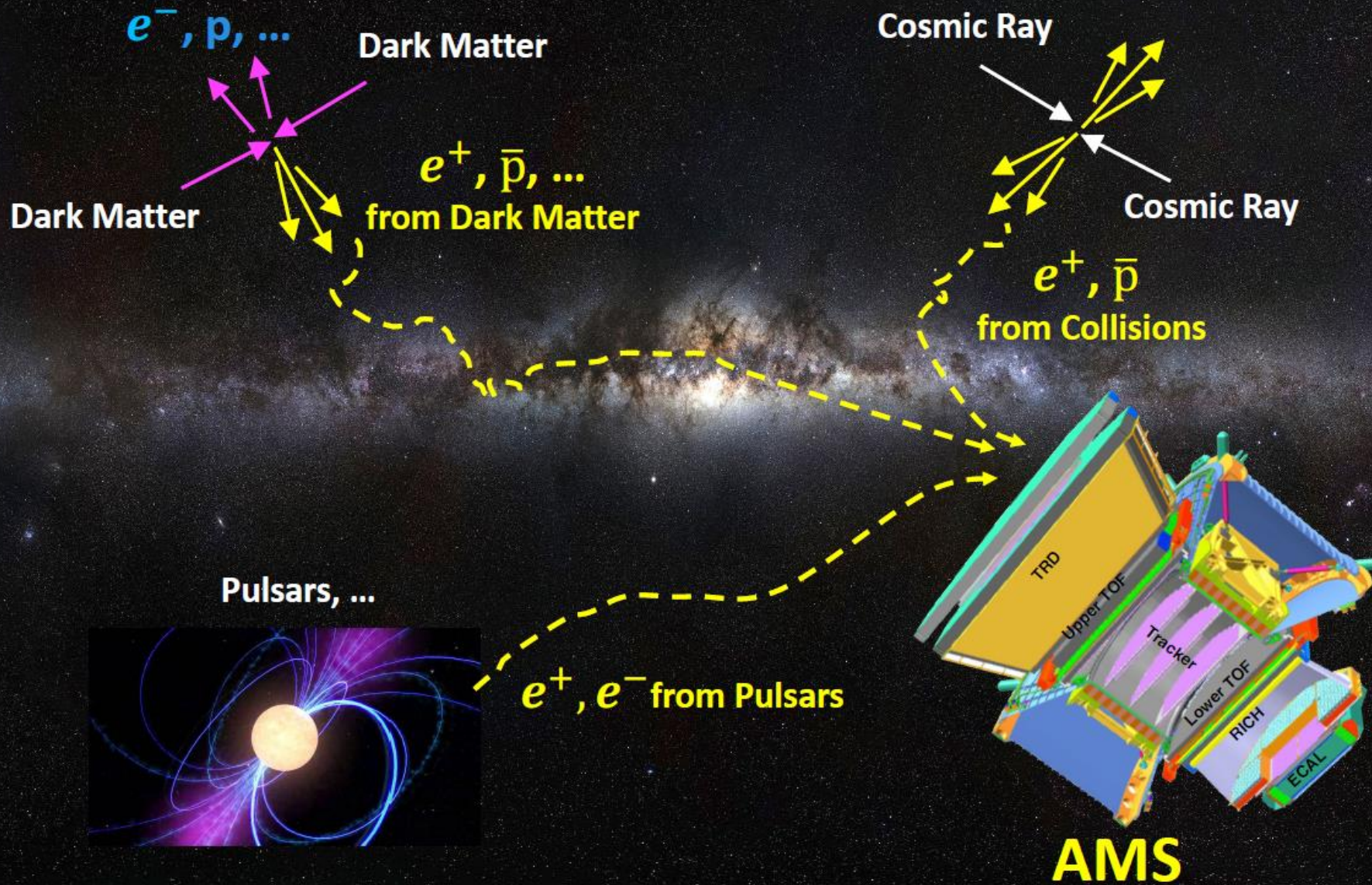
Cosmic Ray Nuclei by 2030

AMS will provide complete and accurate spectra for the

29 elements and provide the foundation for a comprehensive theory of cosmic rays.



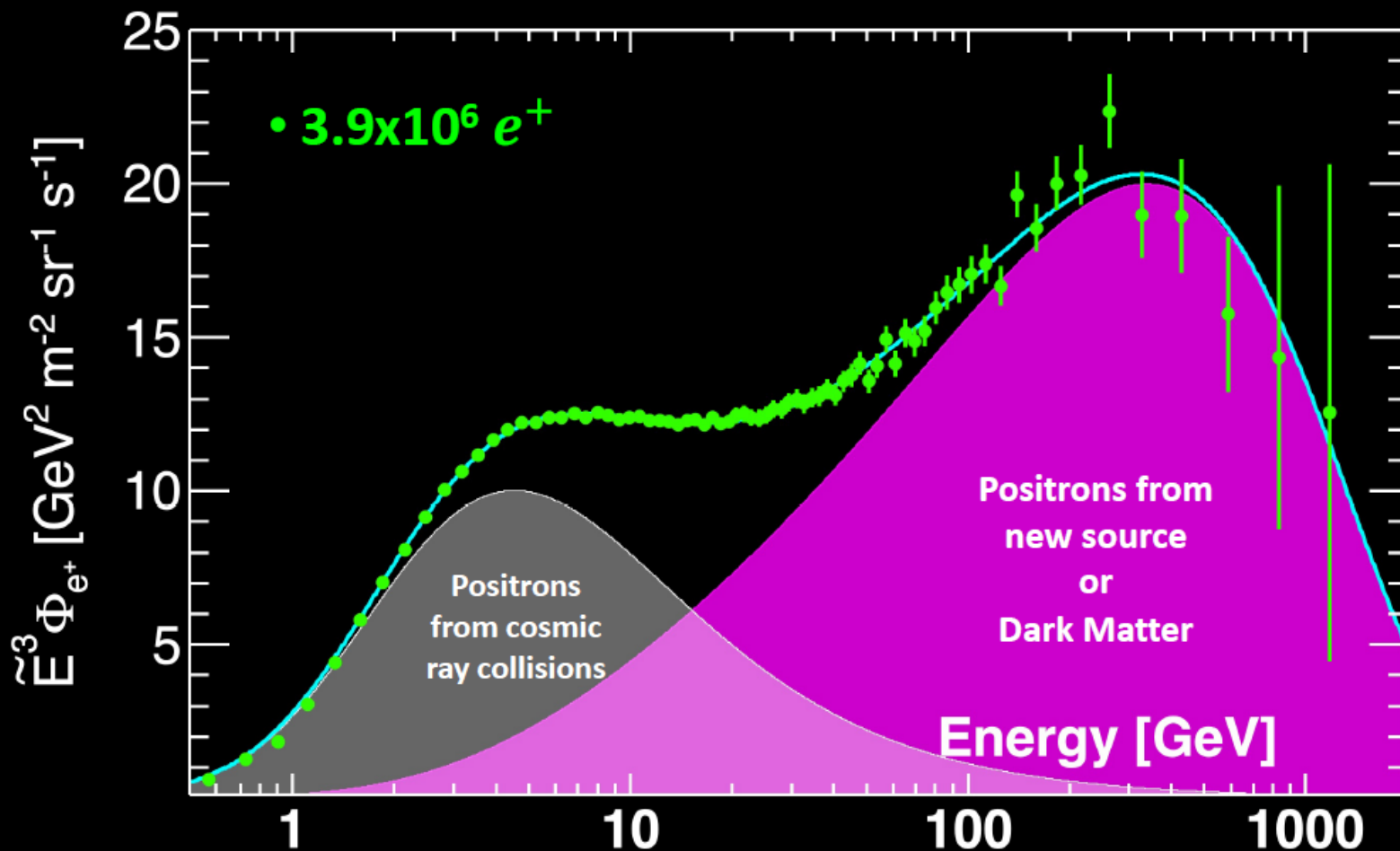
Latest Results on positrons e^+ , electrons e^- , antiprotons \bar{p}



The positron flux is the sum of low-energy part from cosmic ray collisions plus a high-energy part from pulsars or dark matter both with a cutoff energy E_s .

$$\Phi_{e^+}(E) = \frac{E^2}{\hat{E}^2} \left[C_d (\hat{E}/E_1)^{\gamma_d} + C_s (\hat{E}/E_2)^{\gamma_s} \exp(-\hat{E}/E_s) \right]$$

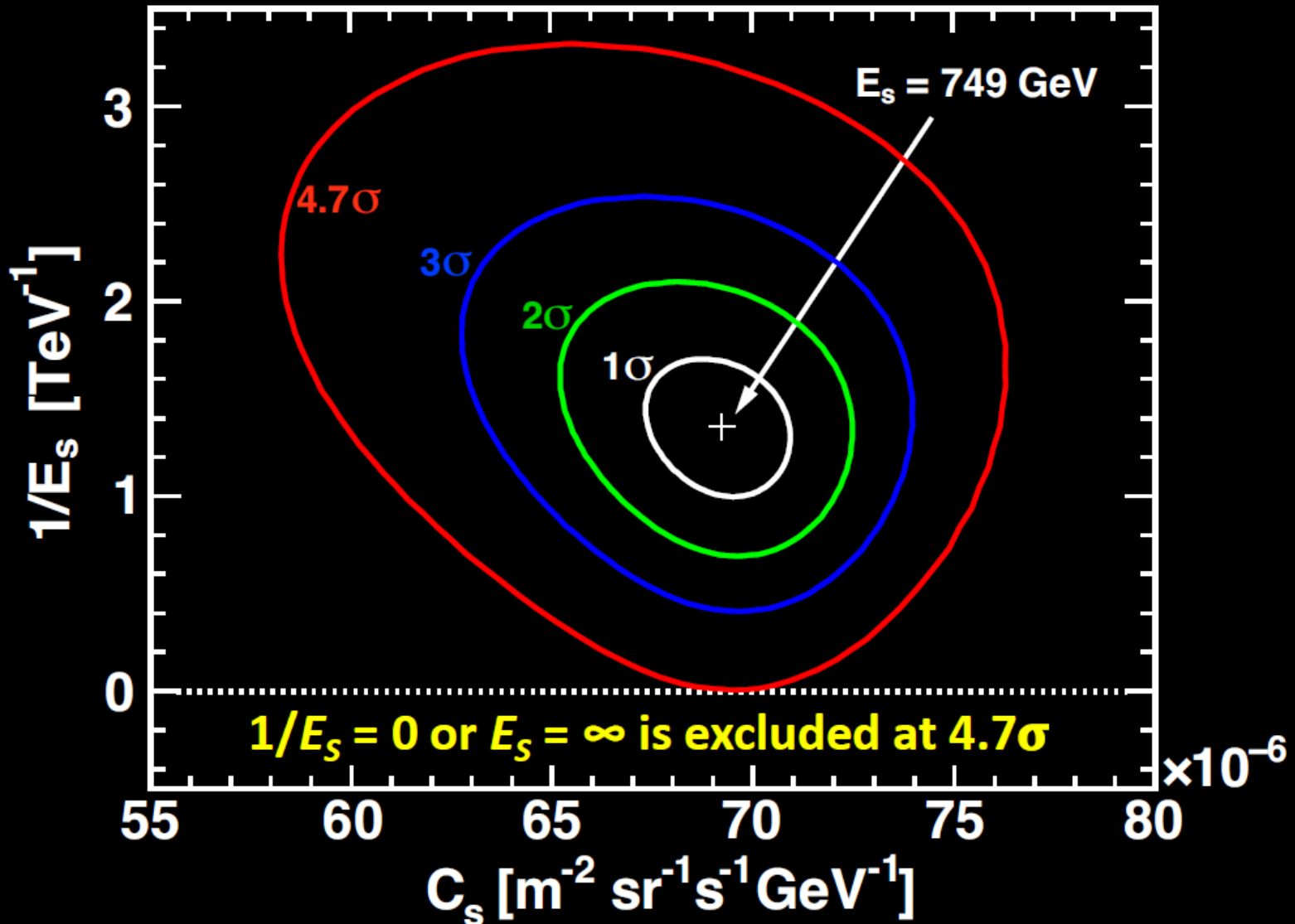
Solar Collisions
Pulsars or Dark Matter



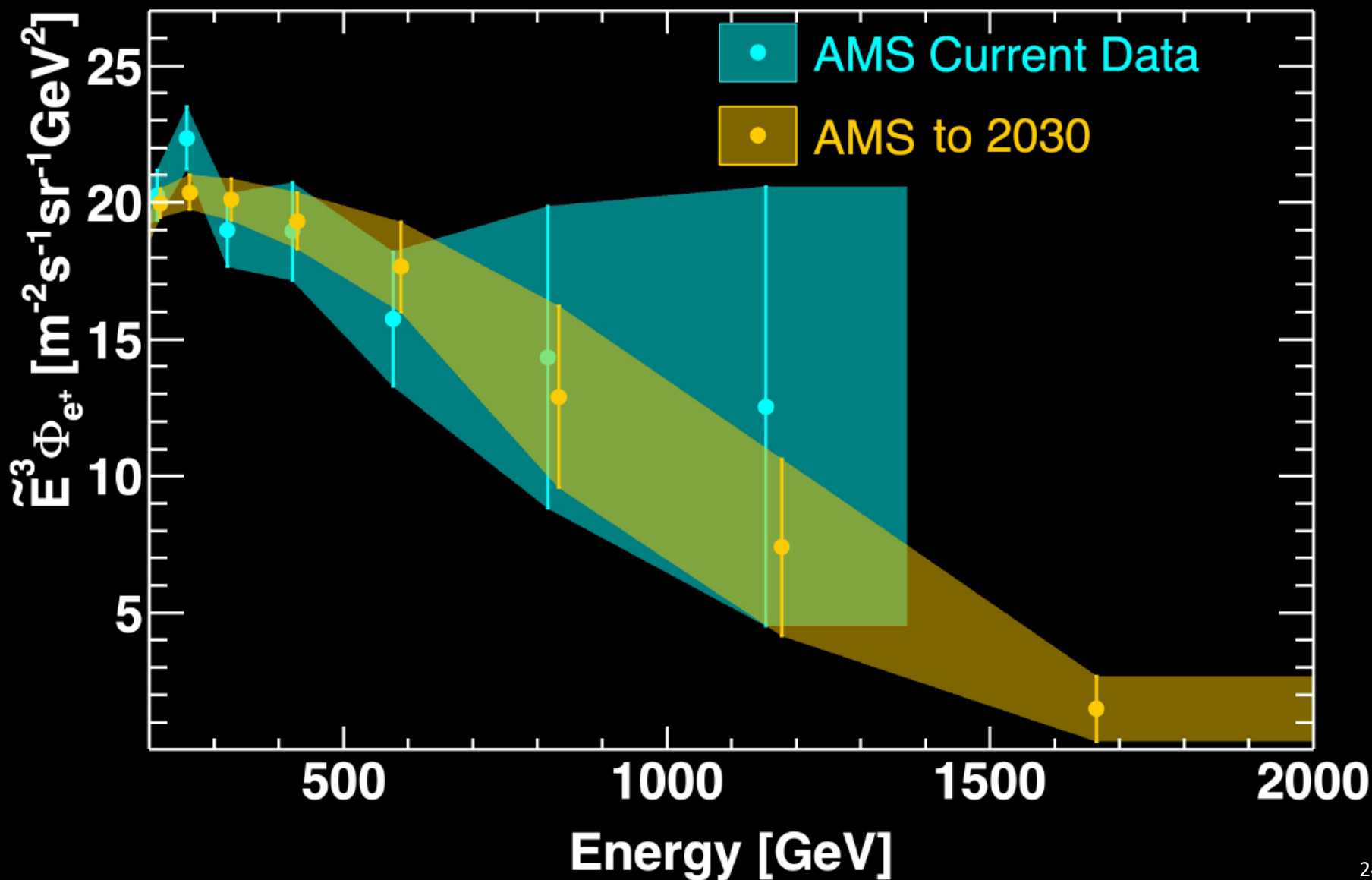
The existence of the finite cutoff energy (4.7σ) is a new and unexpected observation

Determination of the cutoff energy E_s

$$\Phi_{e^+}(E) = \frac{E^2}{\hat{E}^2} \left[\overset{\text{Collisions}}{C_d (\hat{E}/E_1)^{\gamma_d}} + \overset{\text{New Source or Dark Matter}}{C_s (\hat{E}/E_2)^{\gamma_s} \exp(-\hat{E}/E_s)} \right]$$

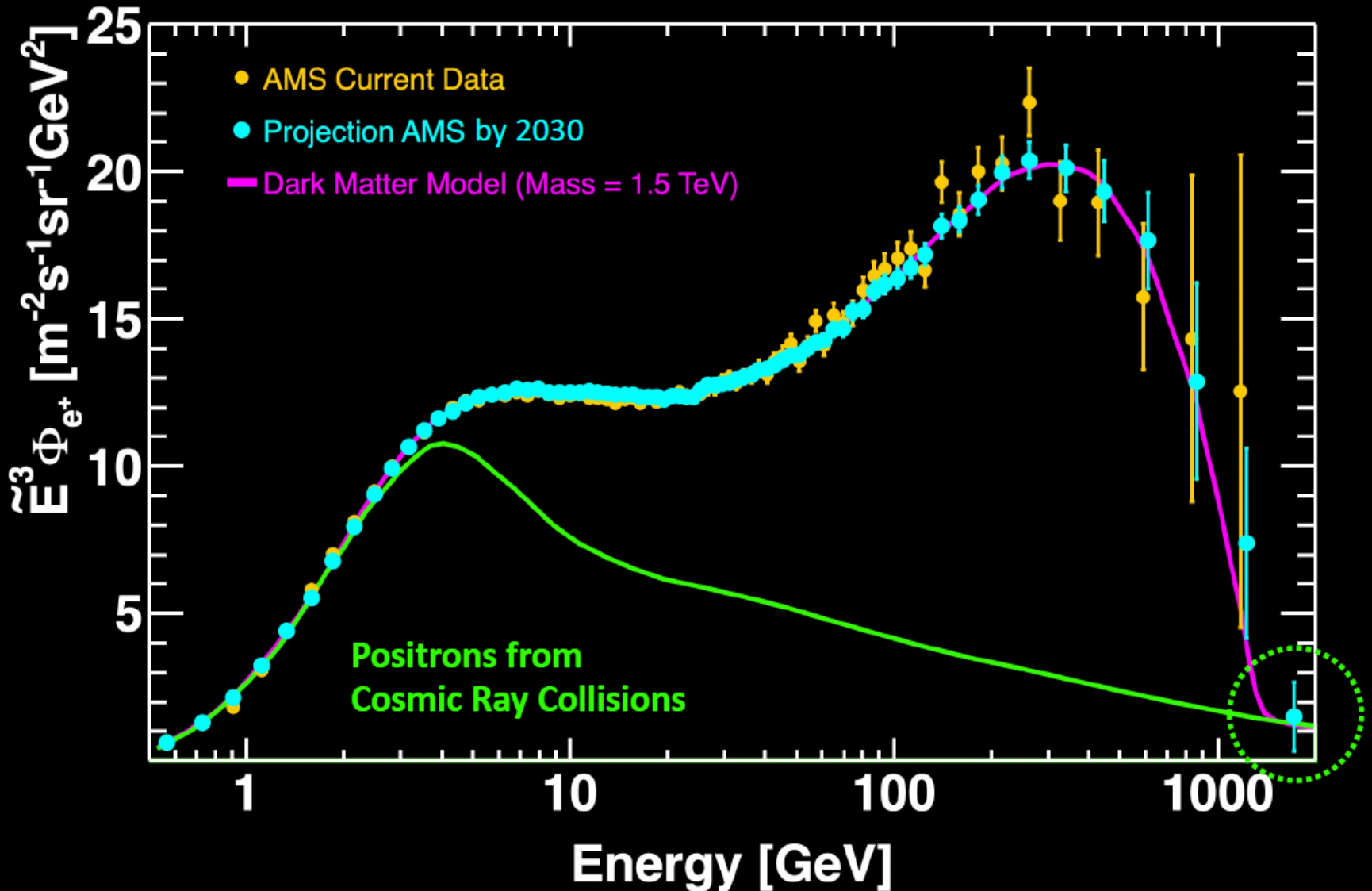


By 2030, AMS will extend the energy range of the positron flux measurement from 1.4 to 2 TeV and reduce the error by a factor of two compared to current data



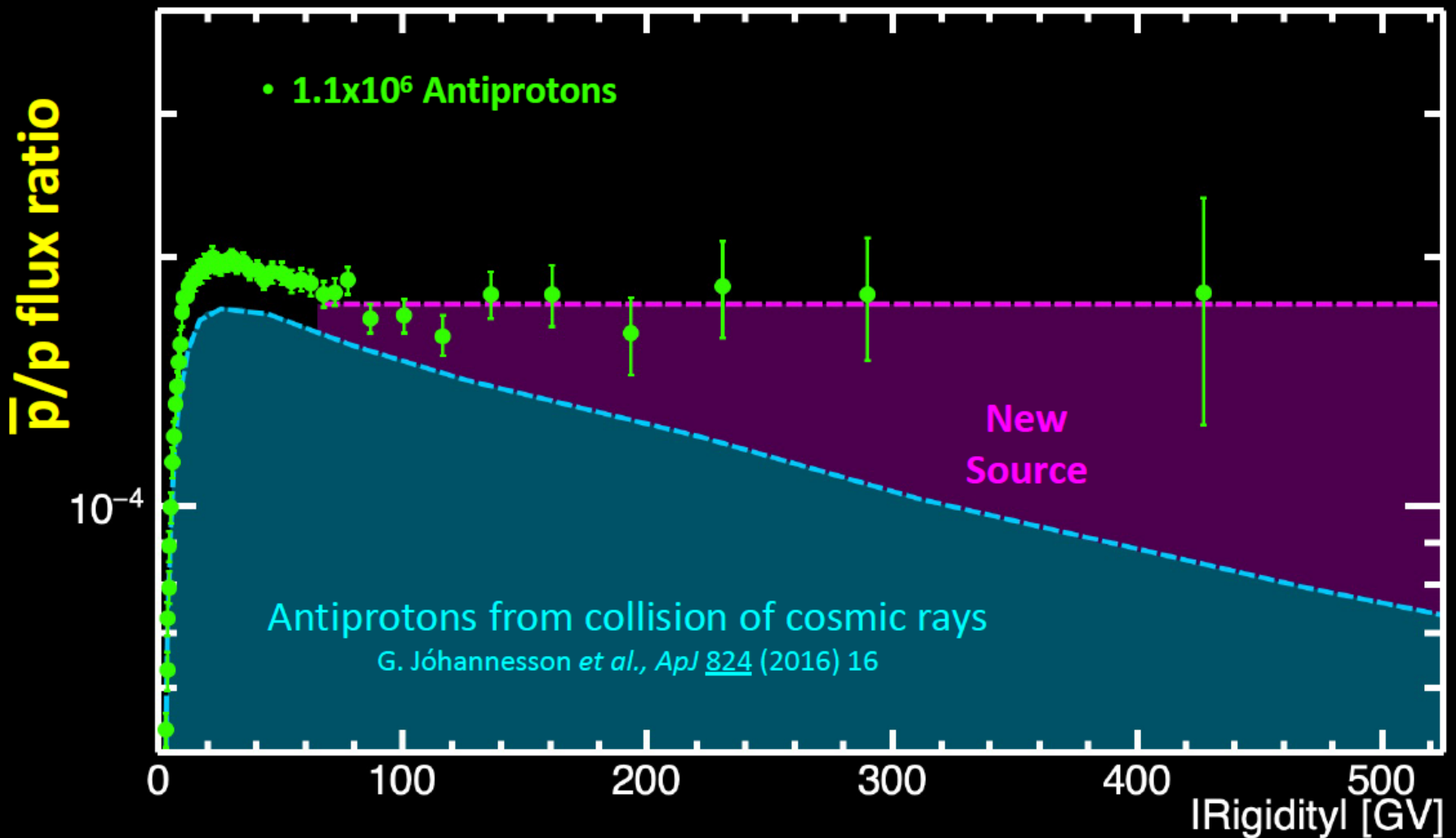
Determination of the Origin of Cosmic Positrons by 2030

AMS will ensure that the measured high energy positron spectrum indeed drops off quickly and, at the highest energies, the positrons only come from cosmic ray collisions as predicted by dark matter models



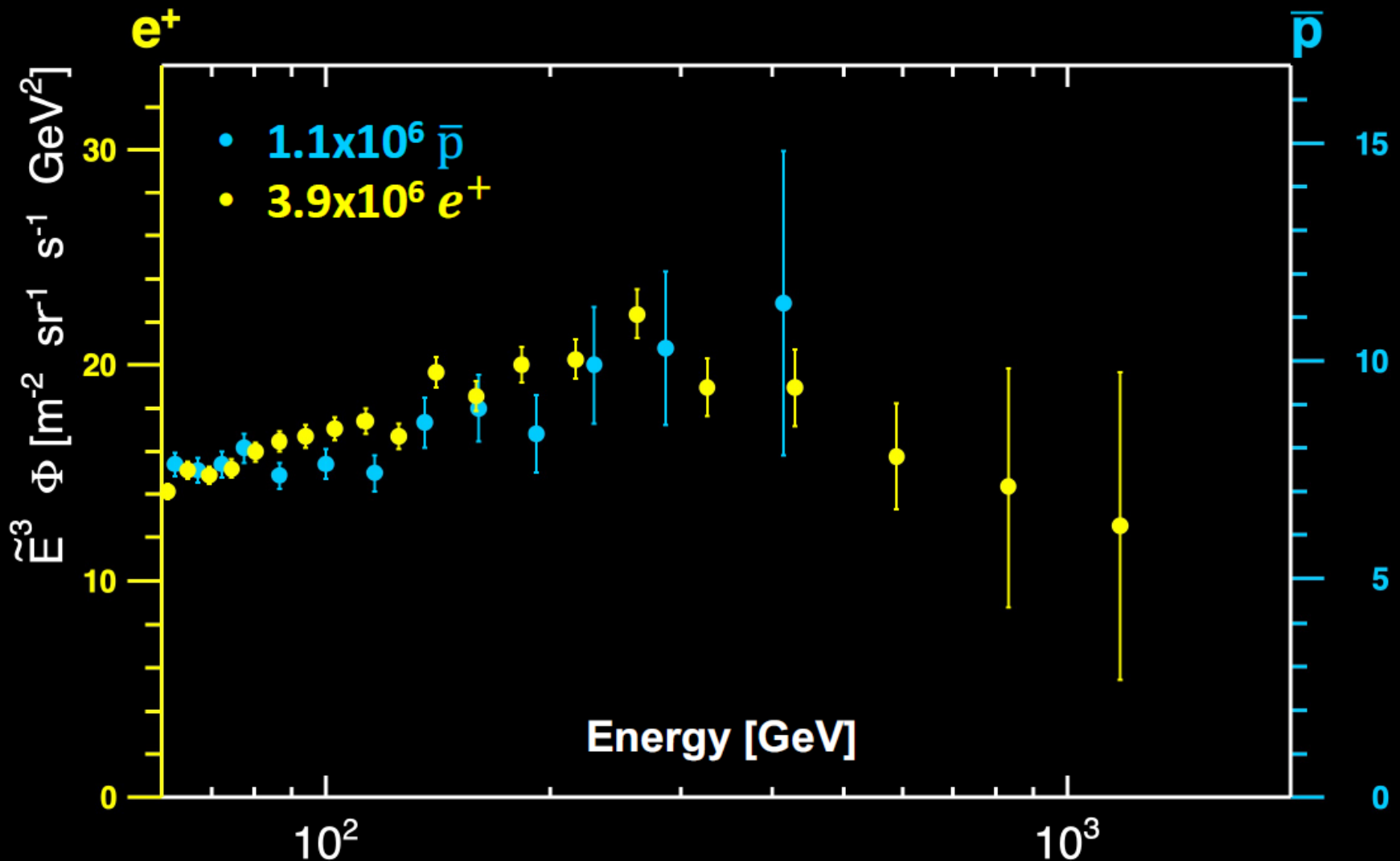
Properties of Cosmic Antiprotons

The antiproton-to-proton flux ratio shows that above 60 GV the ratio is energy independent.



Properties of Cosmic Antiprotons

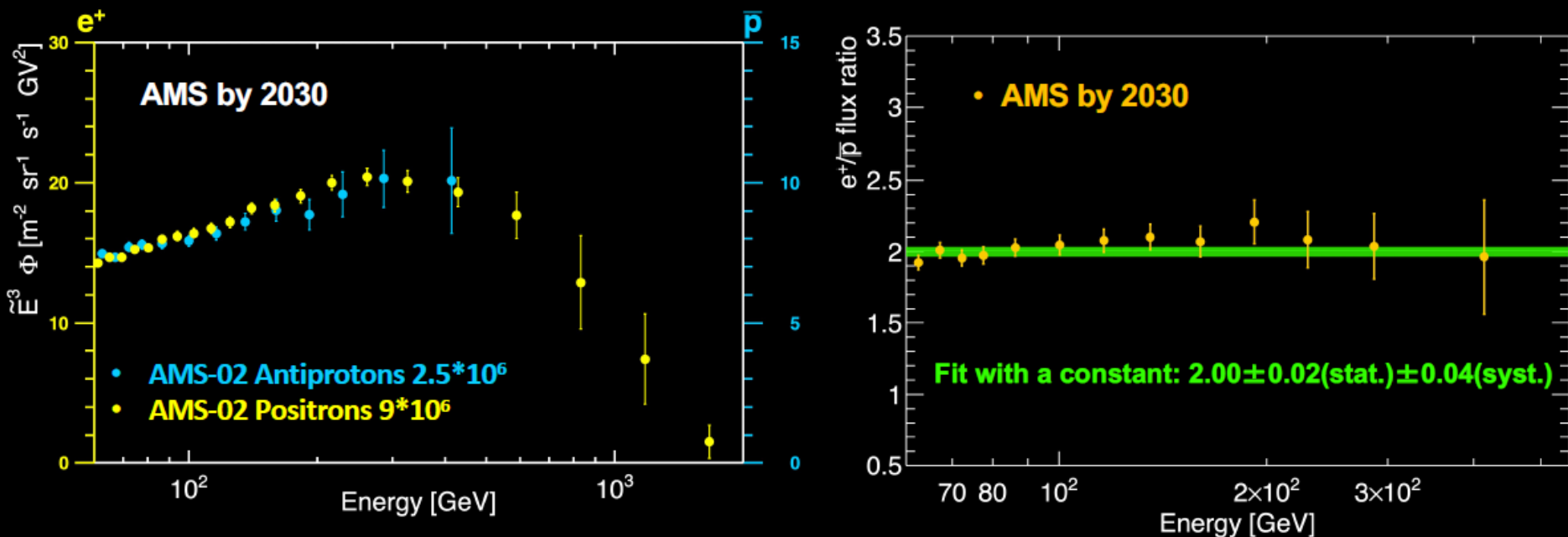
The \bar{p} and e^+ fluxes have identical rigidity dependence.



\bar{p} are not produced by pulsars.

By 2030, AMS will greatly improve the accuracy of the antiproton spectra.

**The identical behaviour of positrons and antiprotons
excludes the pulsar origin of positrons.**

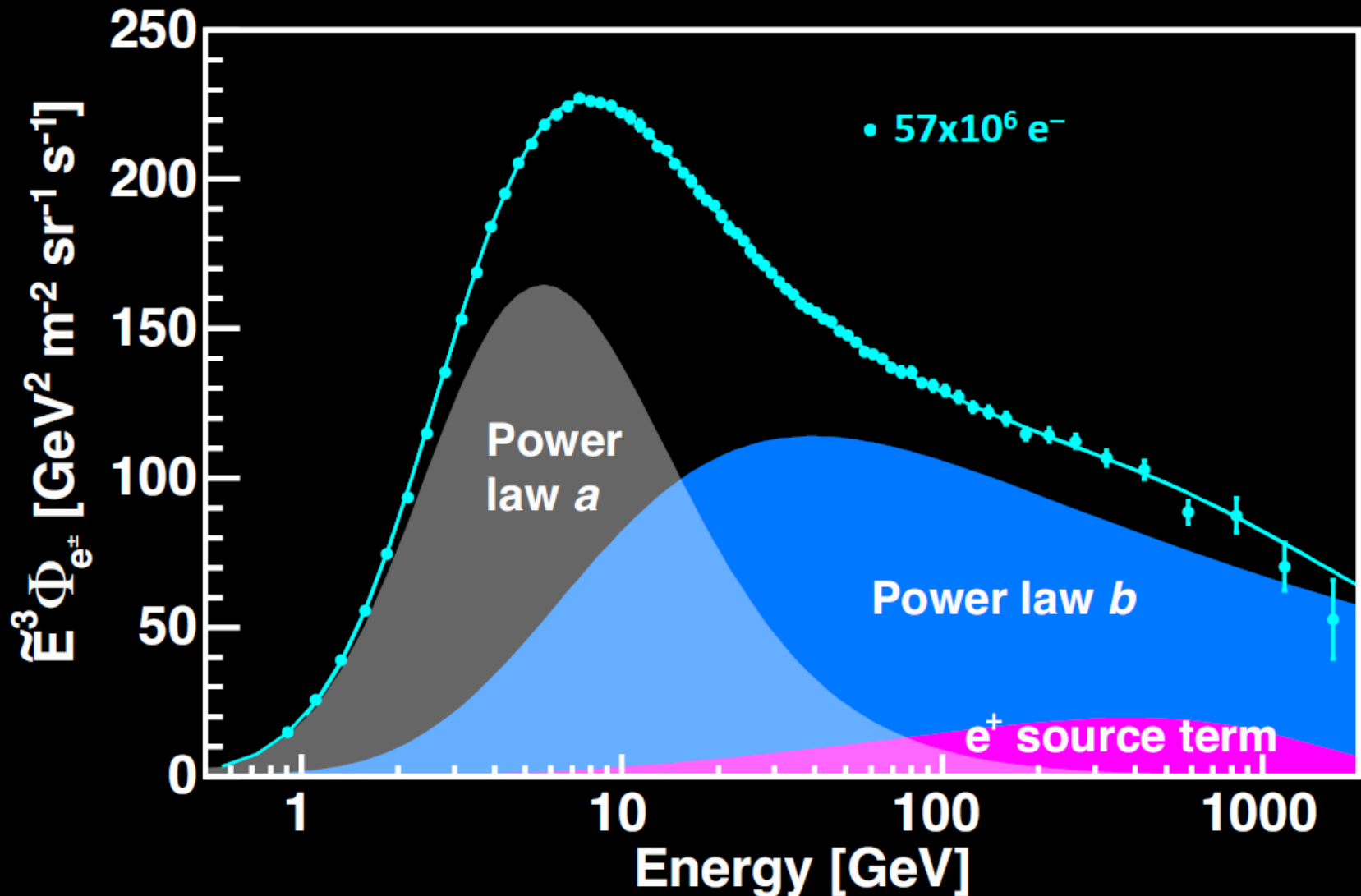


AMS Result on the electron spectrum

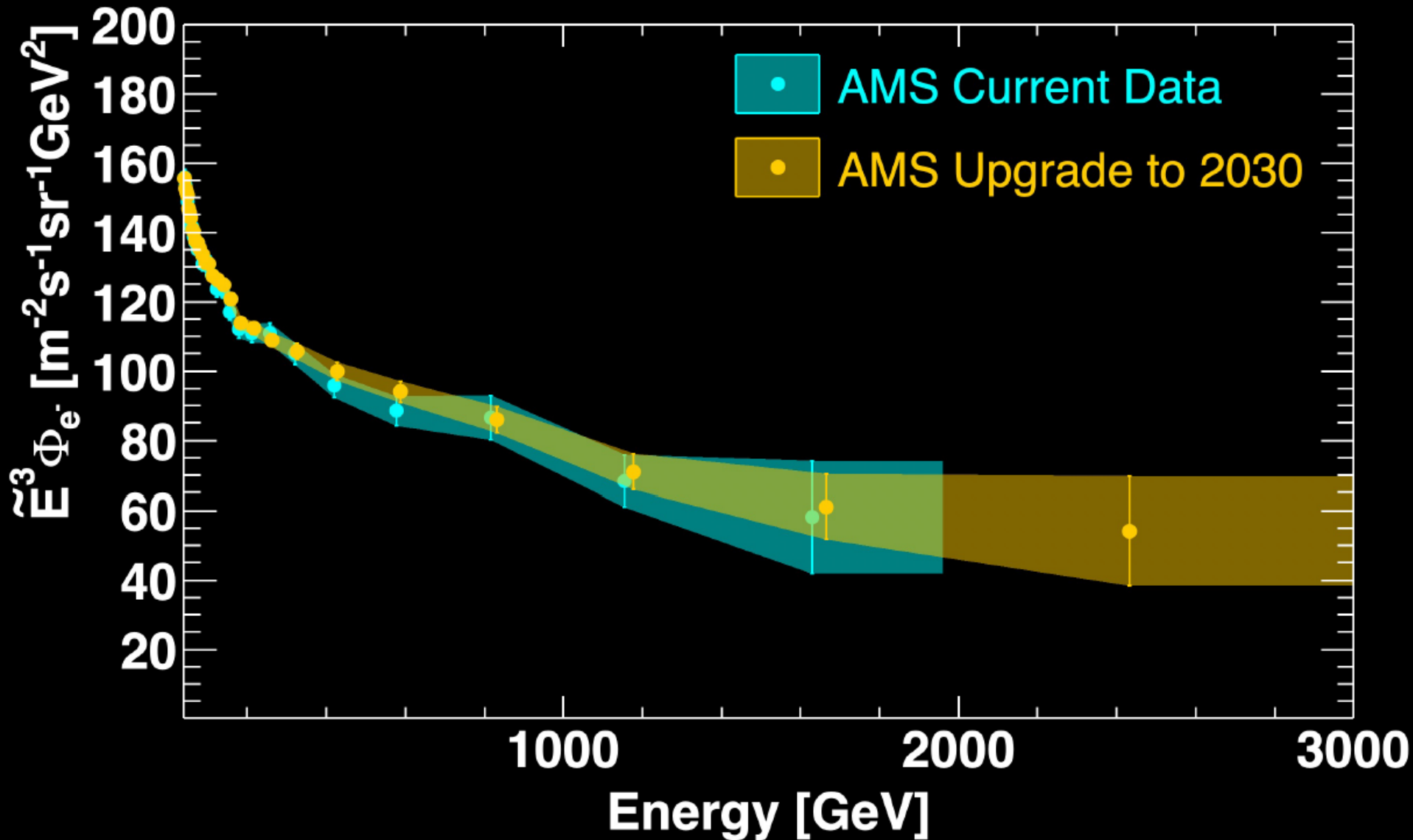
The spectrum fits well with two power laws (a , b) and a source term like positrons

$$\Phi_{e^-}(E) = \frac{E^2}{\widehat{E}^2} (C_a \widehat{E}^{\gamma_a} + C_b \widehat{E}^{\gamma_b} + \text{Positron Source Term})$$

Solar Power law a Power law b

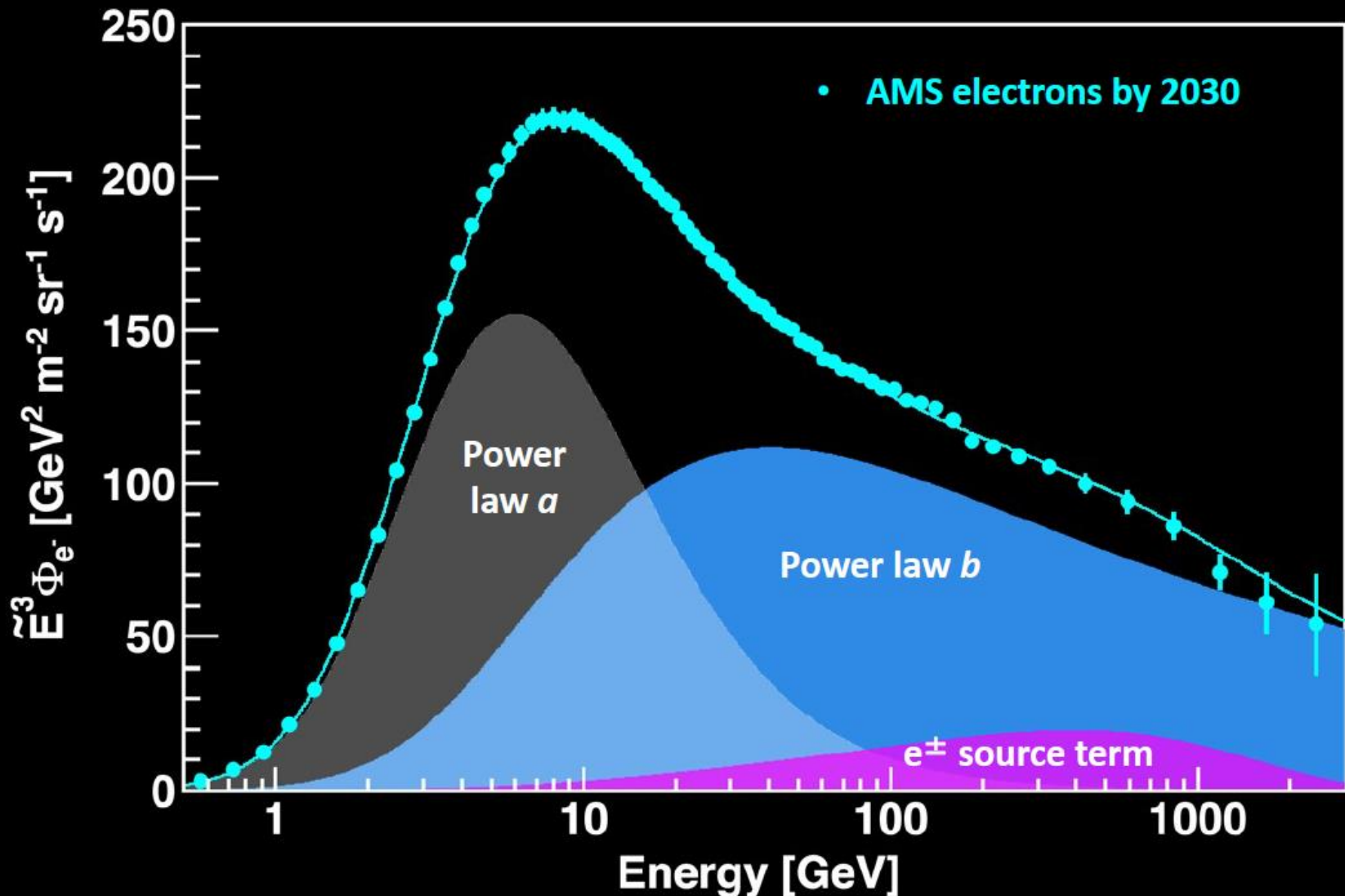


By 2030, AMS will extend the energy range to 3 TeV and reduce the error by a factor of two.

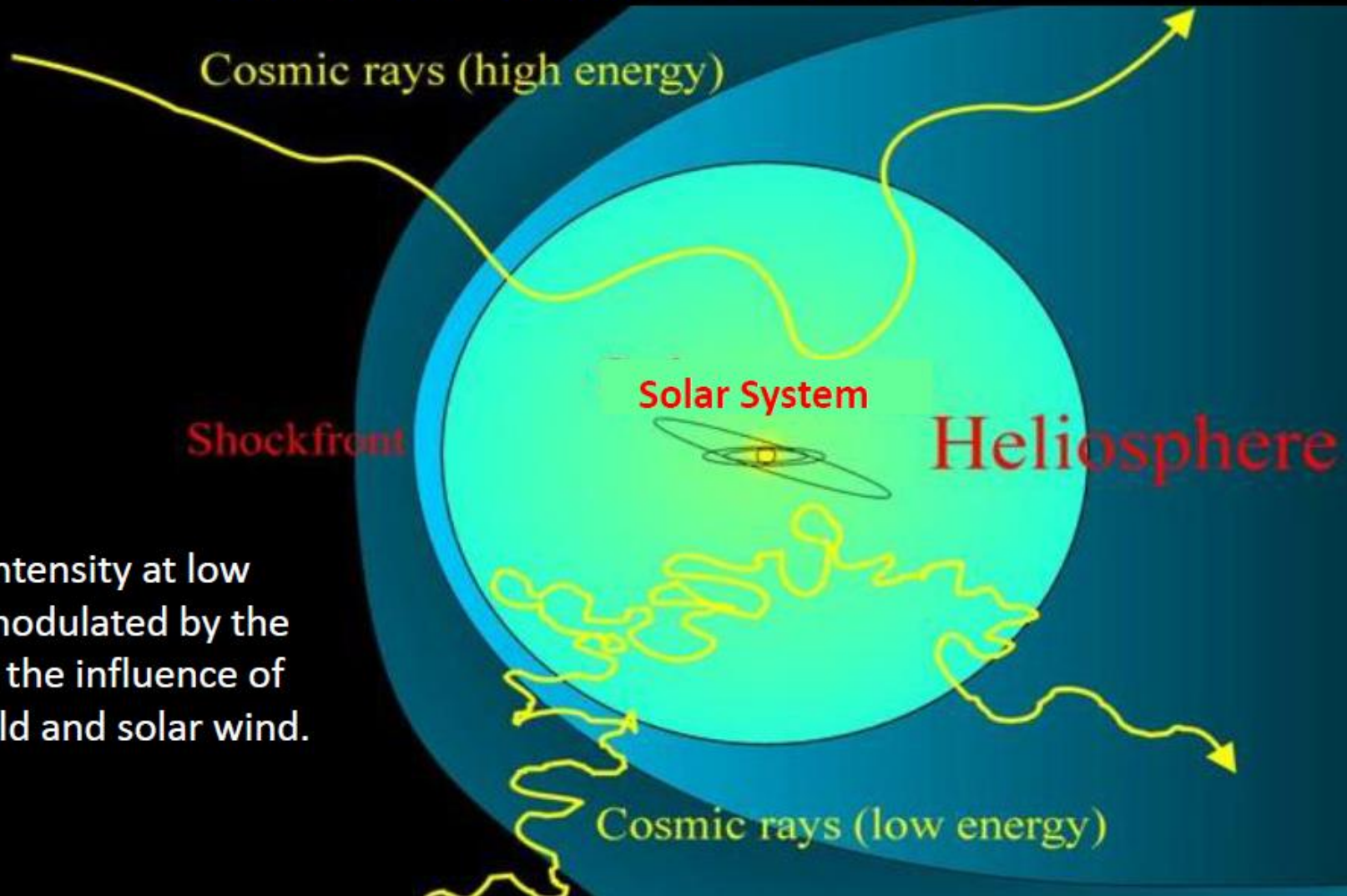


New sources, like Dark Matter or Pulsars,
produce equal amounts of e^+ and e^-

By 2030, the charge-symmetric nature of the high energy source
will be established at the 4σ level



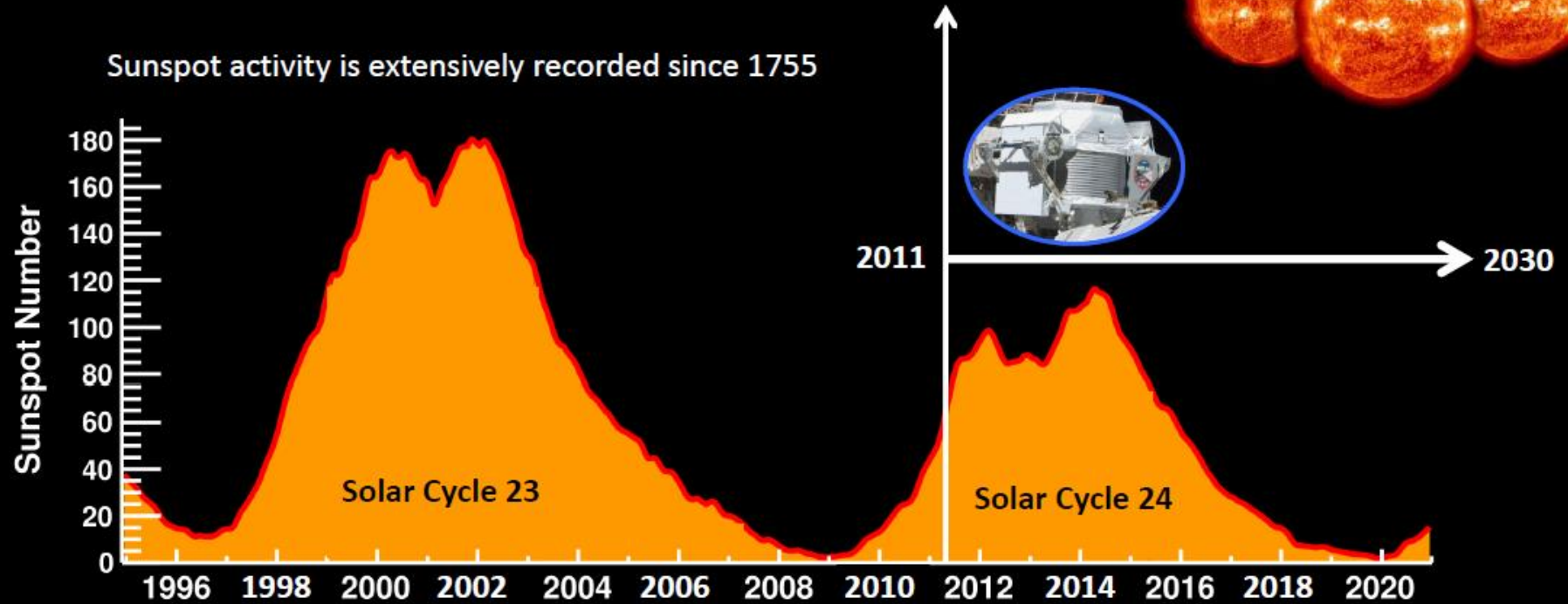
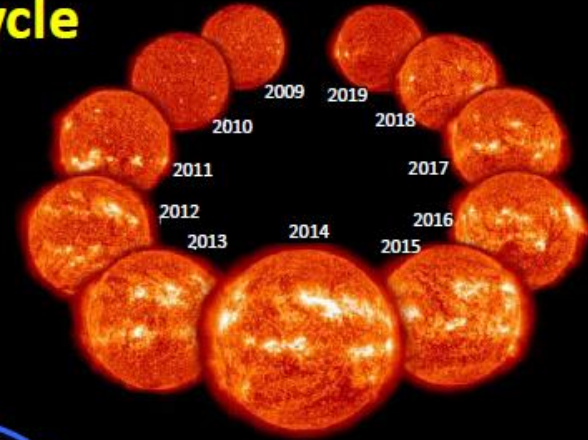
Solar Modulation of Cosmic Rays



Cosmic ray intensity at low energies is modulated by the Sun through the influence of magnetic field and solar wind.

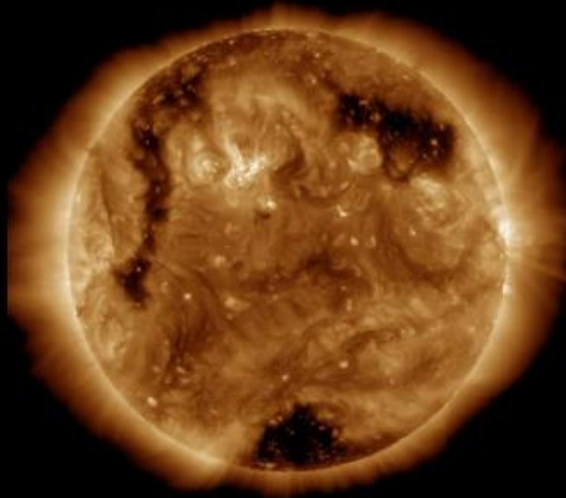
Long Term Variation: Solar Cycle

The most significant long-term scale variation of cosmic rays is related to the 11-year solar cycle.

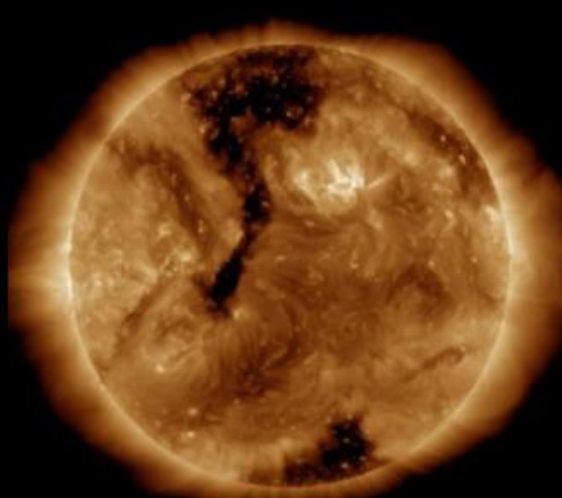


Cosmic Ray Recurrent Variation in Short Scale

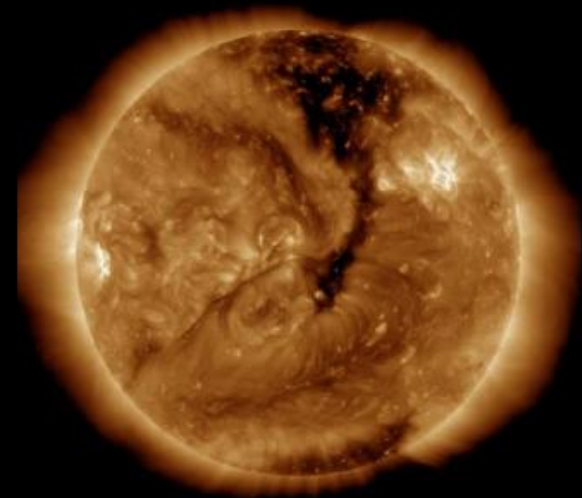
Short scale variation of cosmic rays are related to Sun's rotation (Bartels' rotation: 27 days).



2016-03-22



2016-03-24



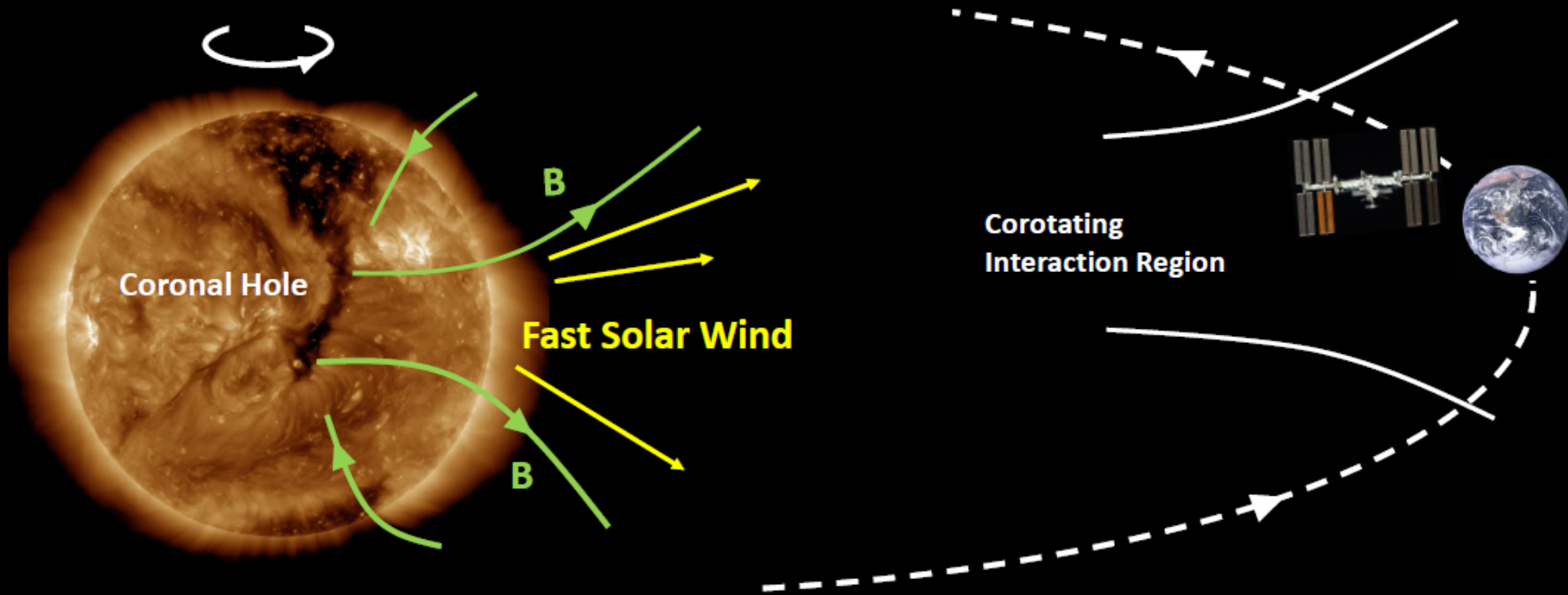
2016-03-26

Image taken by Dynamics Observatory (SDO), NASA

Coronal holes are regions where plasma density and temperature are lower, so they appear darker in images.

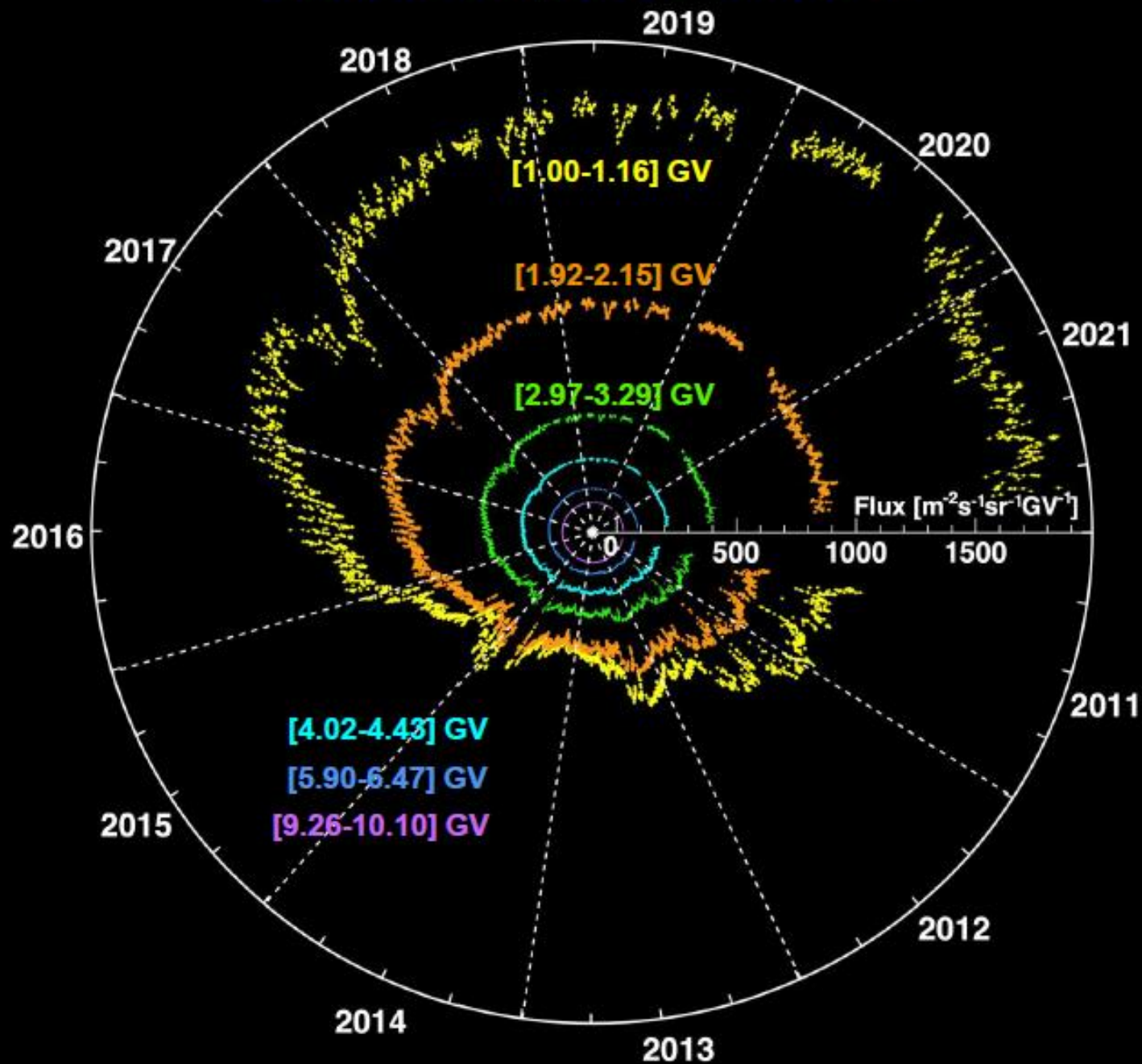
Cosmic Ray Recurrent Variation in Short Scale

Coronal Holes are sources of high speed solar wind affecting Earth.



Precision measurements of the individual species of cosmic rays in a solar cycle provide unique inputs for the understanding of cosmic rays in the heliosphere.

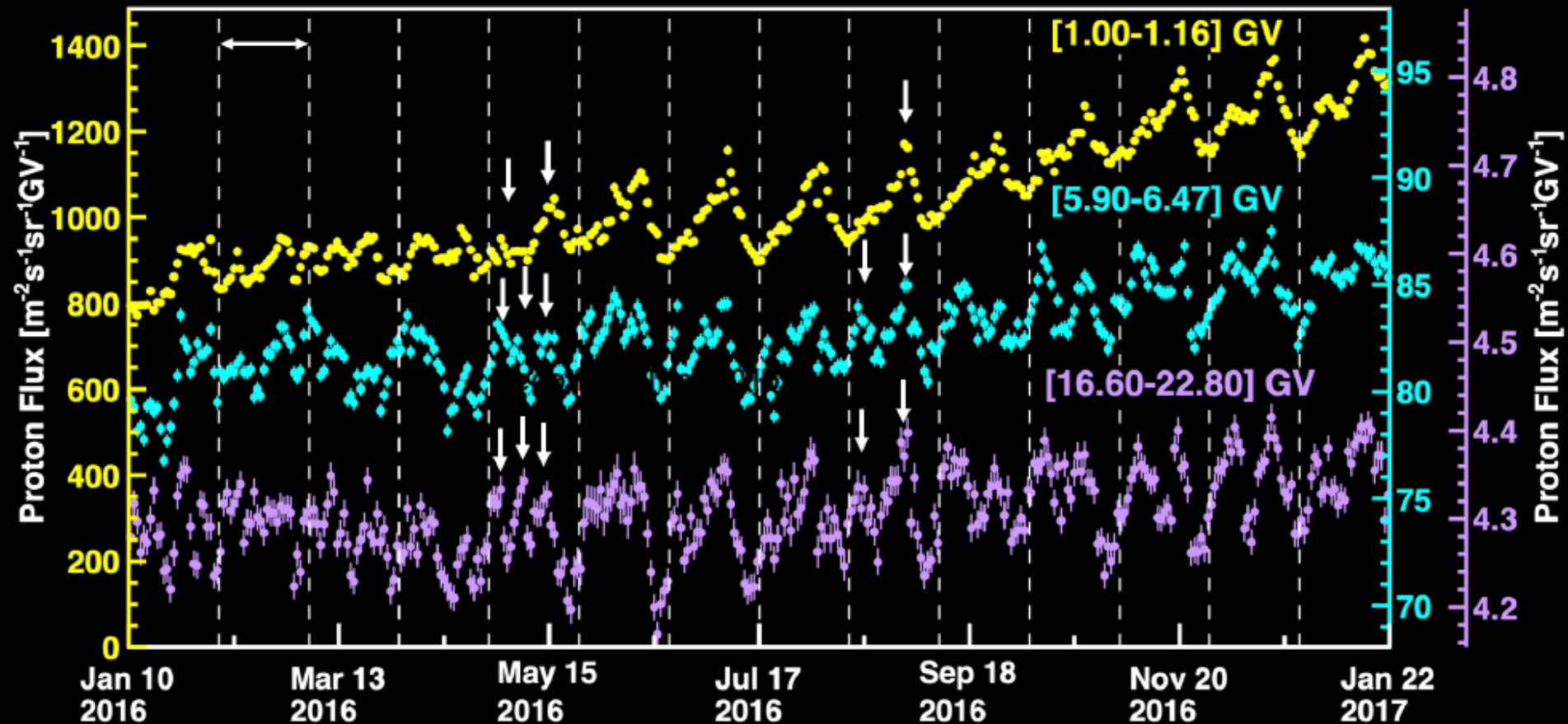
AMS Daily Proton Flux



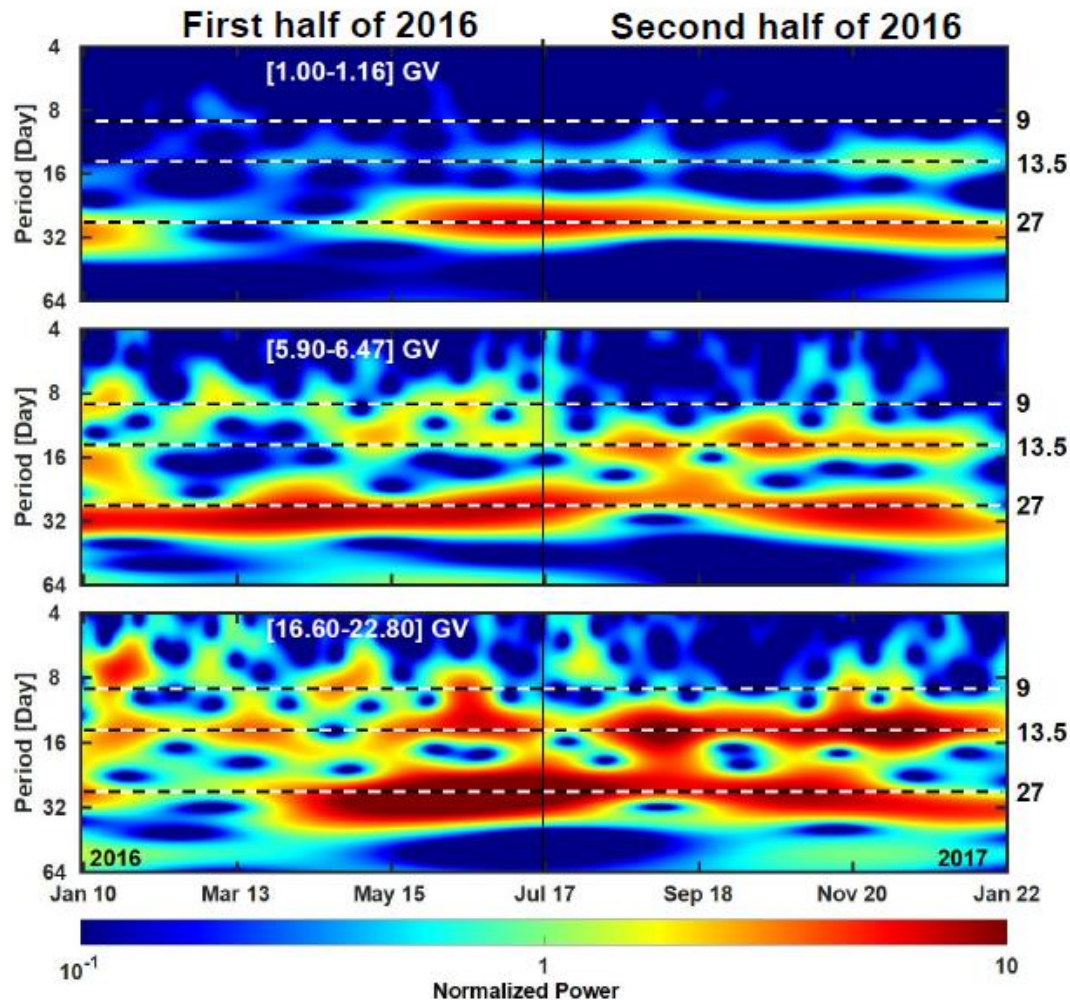
Recurrent Proton Flux Variation in 2016

Double-peak and triple-peak structures are visible in different Bartels rotations

27 days



Wavelet Analysis of Proton Fluxes in 2016



To study the recurrent time variations in the daily proton fluxes, a **wavelet time-frequency** technique was used.

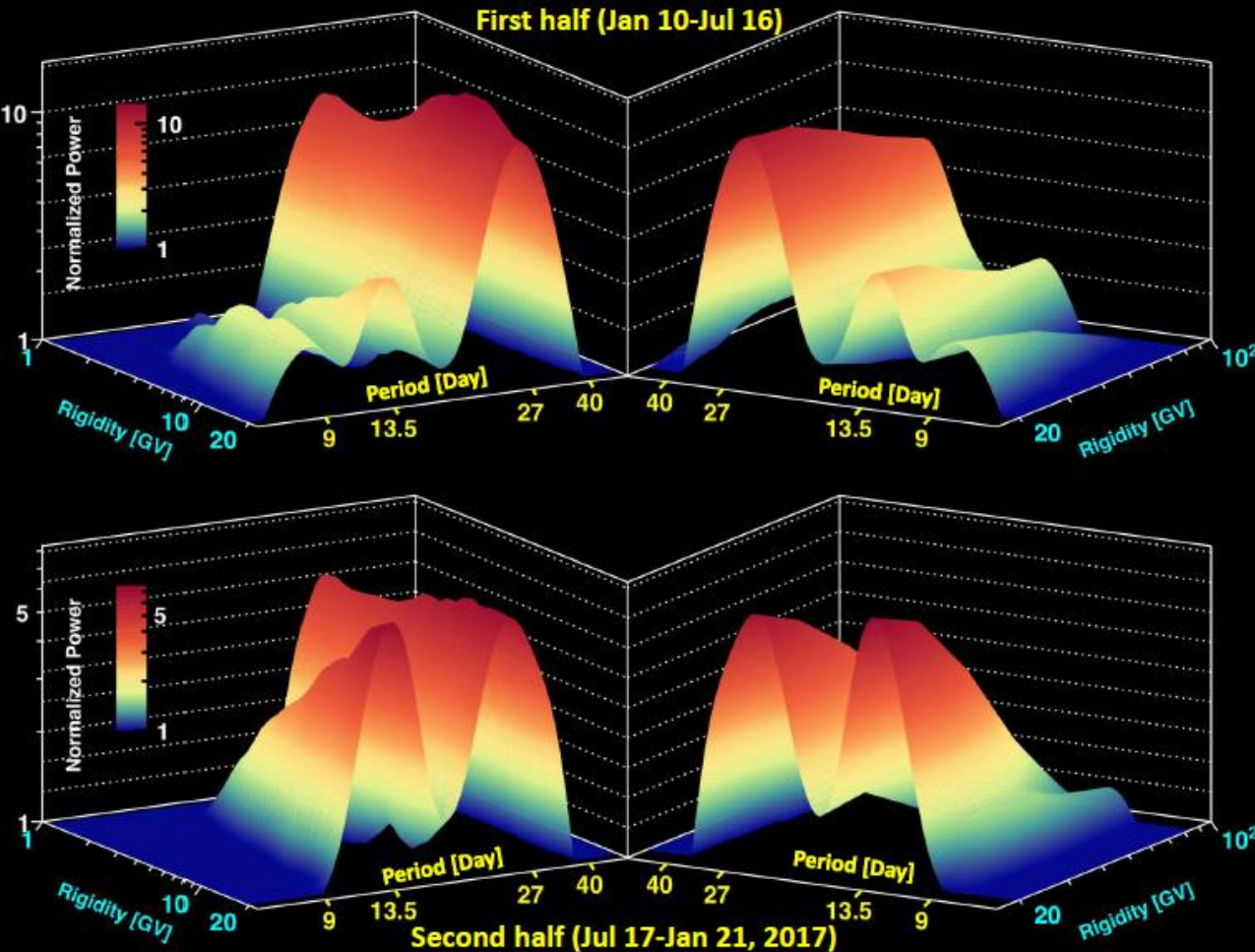
To show the strength of the periodicity, **the normalized power** is defined by the power divided by **the variance** of the time series.

Periods of 9, 13.5, and 27 days are observed in 2016.

The strength of all three periodicities changes with time and rigidity.

In particular, shorter periods of 9 and 13.5 days, when present, are more visible at 6 GV and 20 GV compared to 1 GV.

Periodicities of Daily Proton Fluxes in 2016



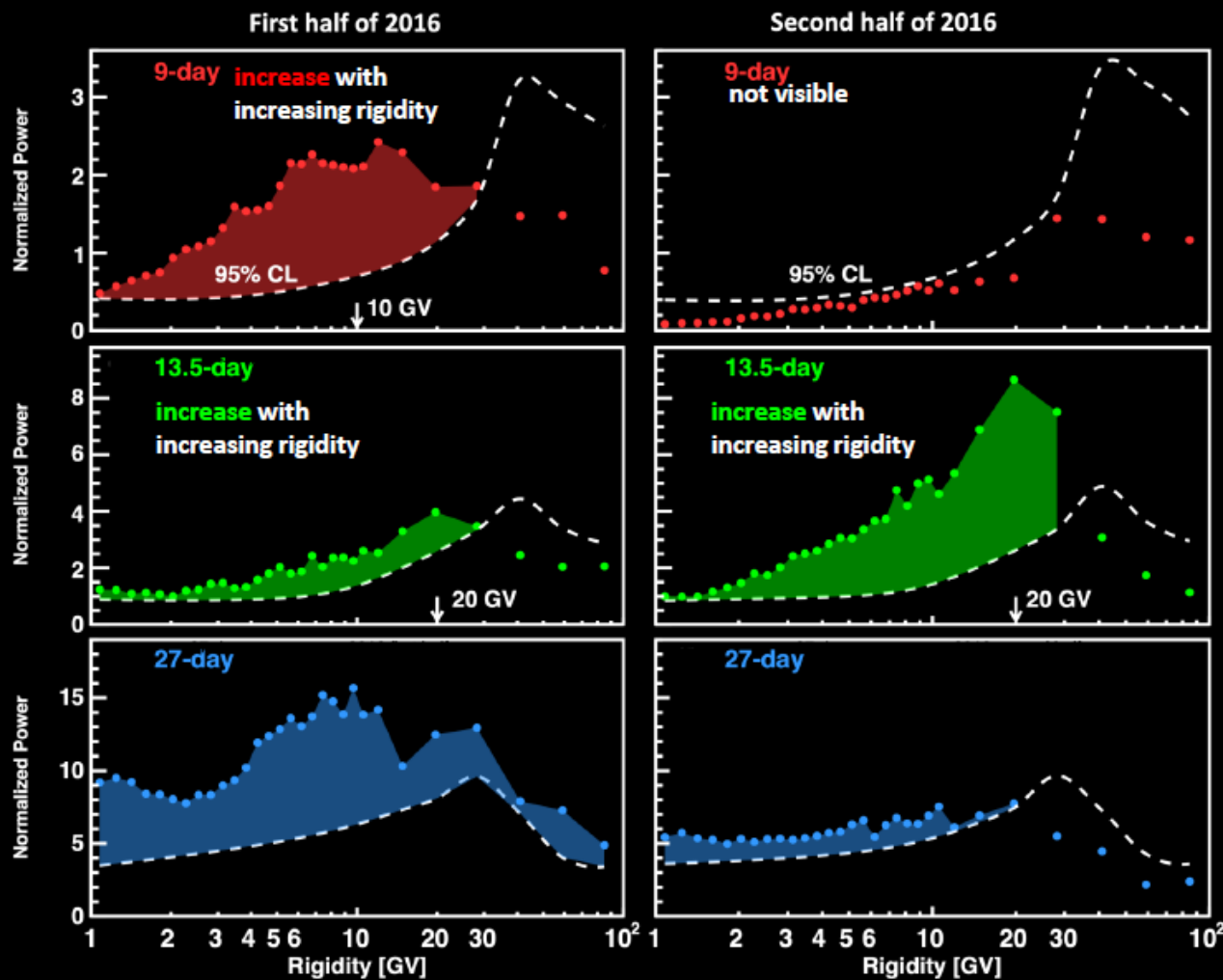
Highlight #1

Unexpectedly, the strength of **9-day and 13.5-day periodicities** increases with increasing rigidity up to **~10 GV** and **~20 GV**, respectively. Then the strength decreases with increasing rigidity up to **100 GV**.

Thus, the AMS results do not support the general conclusion that the strength of the periodicities always decreases with increasing rigidity

Phys. Rev. Lett. 127, 271102 (2021)

Rigidity Dependence of 9-day, 13.5-day, and 27-day periods of protons

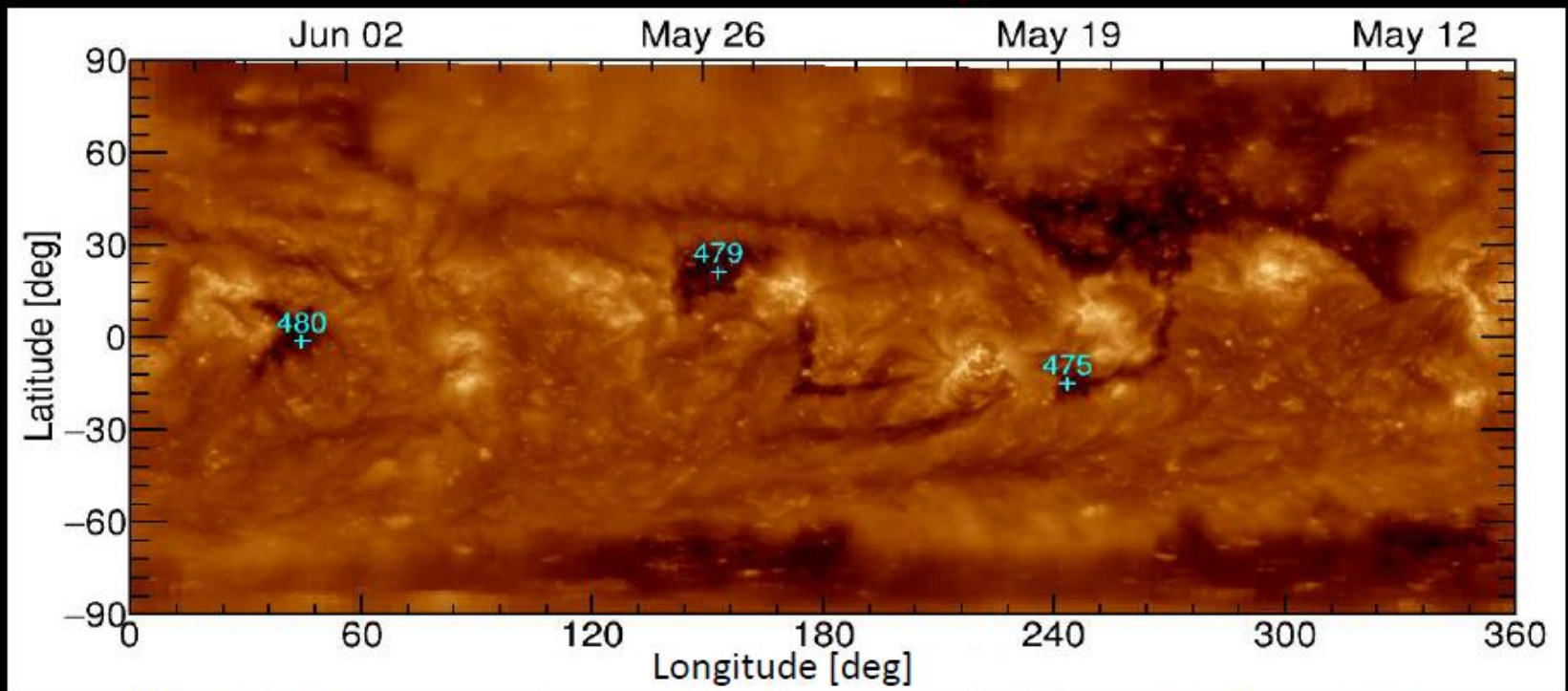


Shaded areas are the rigidity intervals where the periodicity is prominent

Cosmic Ray Periodicities and the Rotation of the Sun

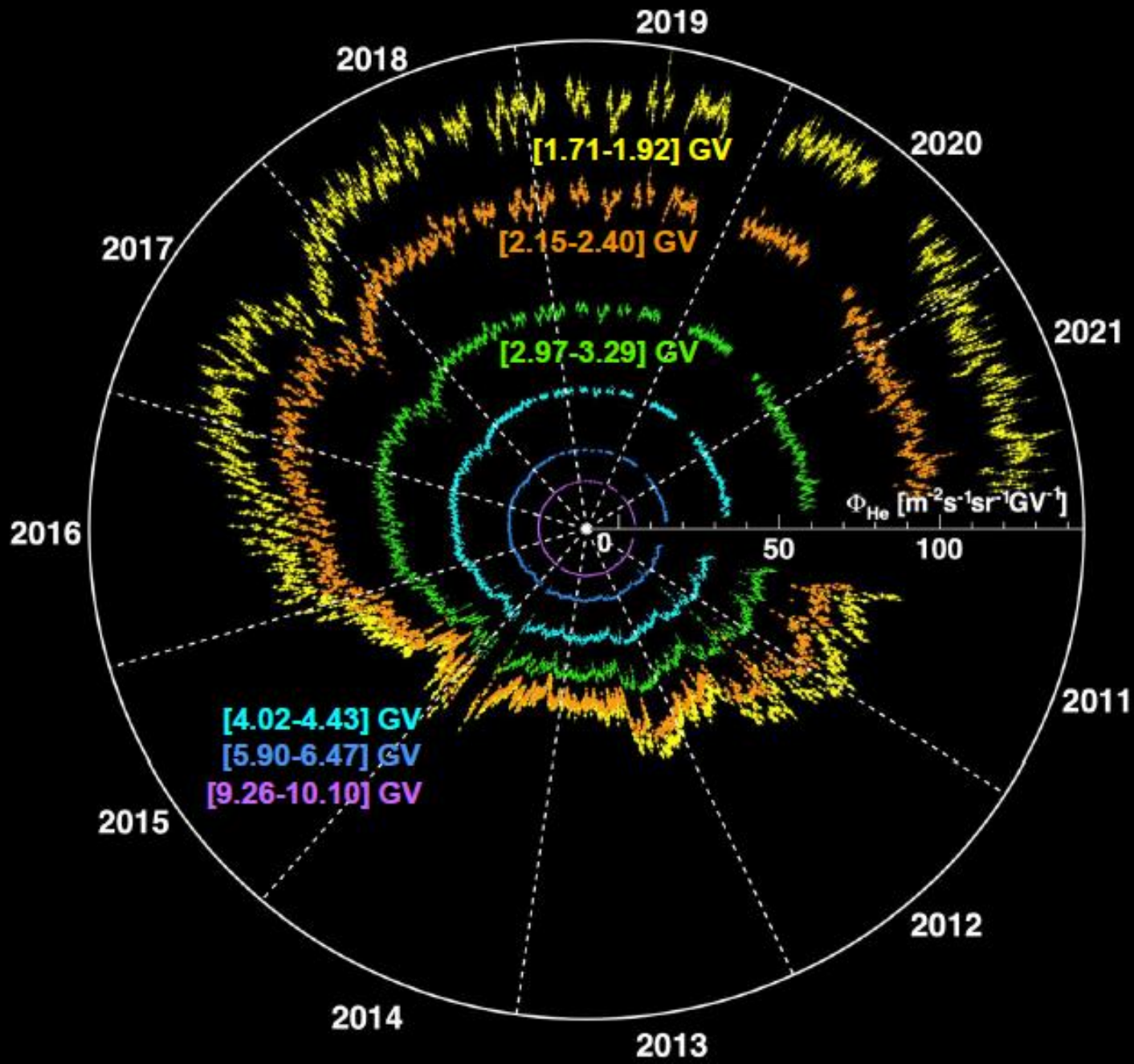
Coronal Holes are sources of high speed solar wind affecting Earth. The rotation of the Sun causes multiple periods in the flux:

- 0 coronal hole: —→ No apparent periods
- 1 coronal hole —→ 27-day period (a Bartels rotation)
- 2 coronal holes separated by 180° —→ 13.5-day period
- 3 coronal holes separated by 120° —→ 9-day period

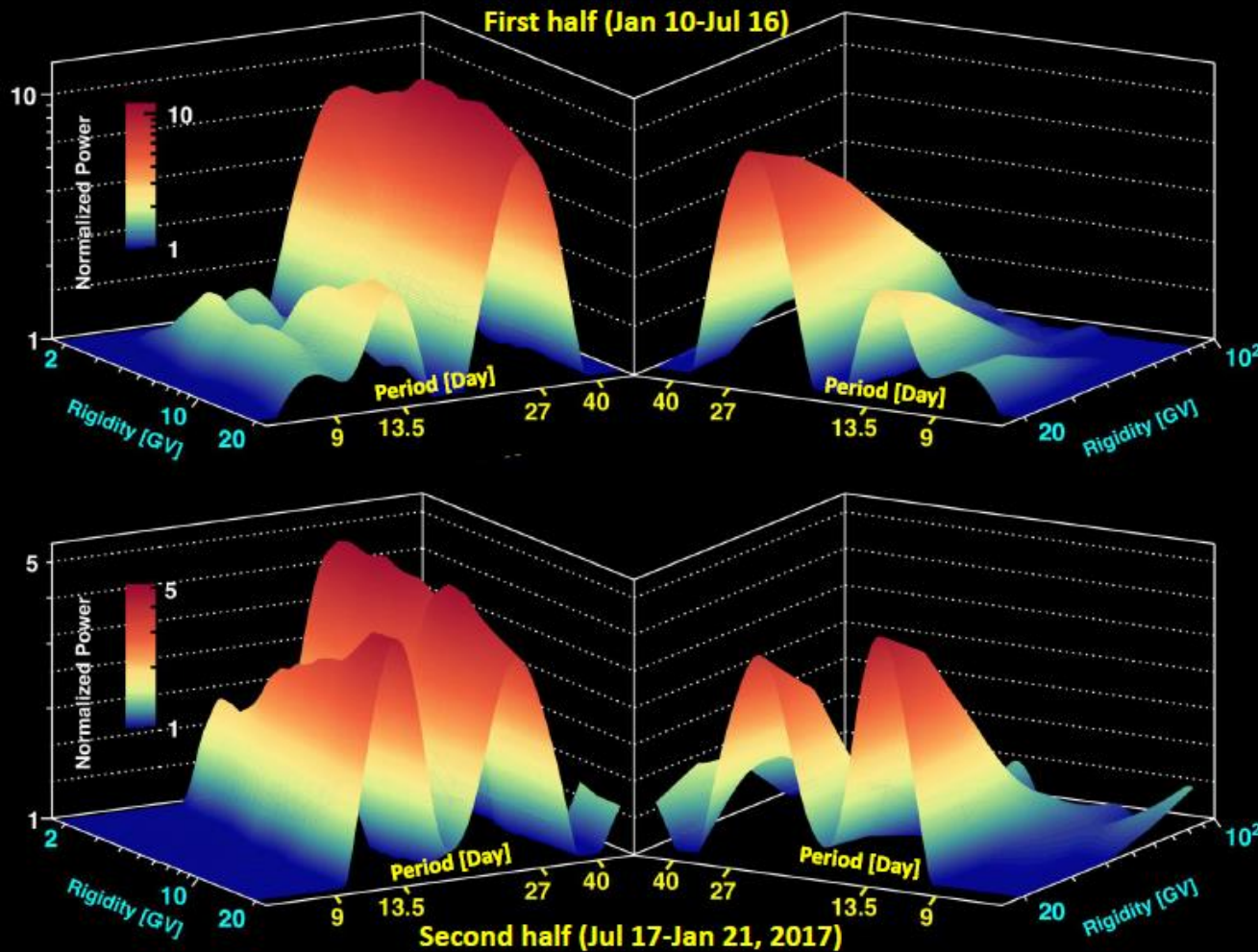


(May 10, 2016-Jun 06, 2016) Image taken by Solar Dynamics Observatory (SDO), NASA

AMS Daily Helium Flux



Periodicities of Daily Helium Fluxes in 2016



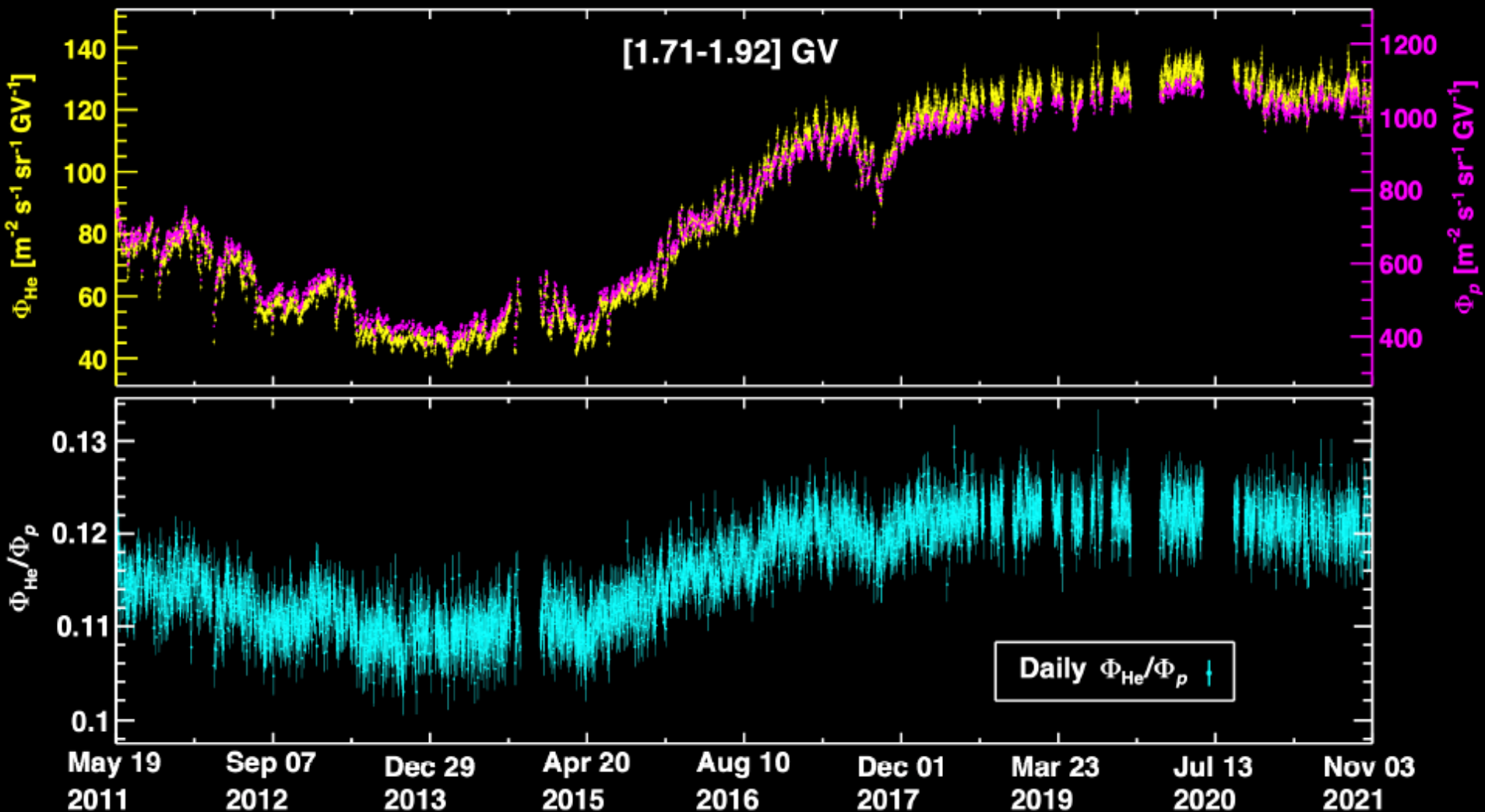
Similar periodic structures are observed for helium.

The AMS results do not support the general conclusion that the strength of the periodicities always decreases with increasing rigidity

Phys. Rev. Lett. 128, 231102 (2022)

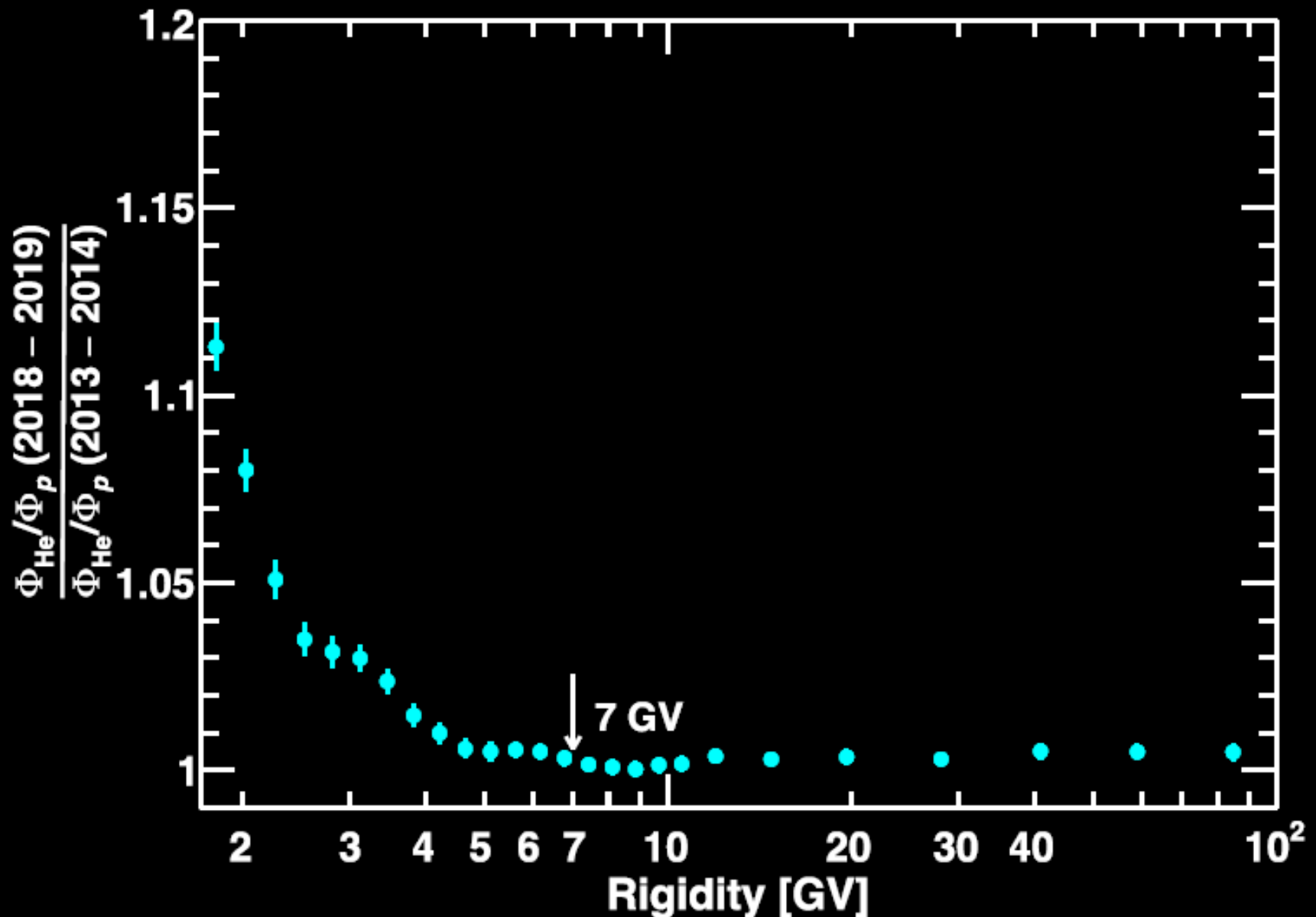
Daily Φ_{He} , Φ_p and Φ_{He}/Φ_p

Φ_{He}/Φ_p exhibits variations on multiple timescales



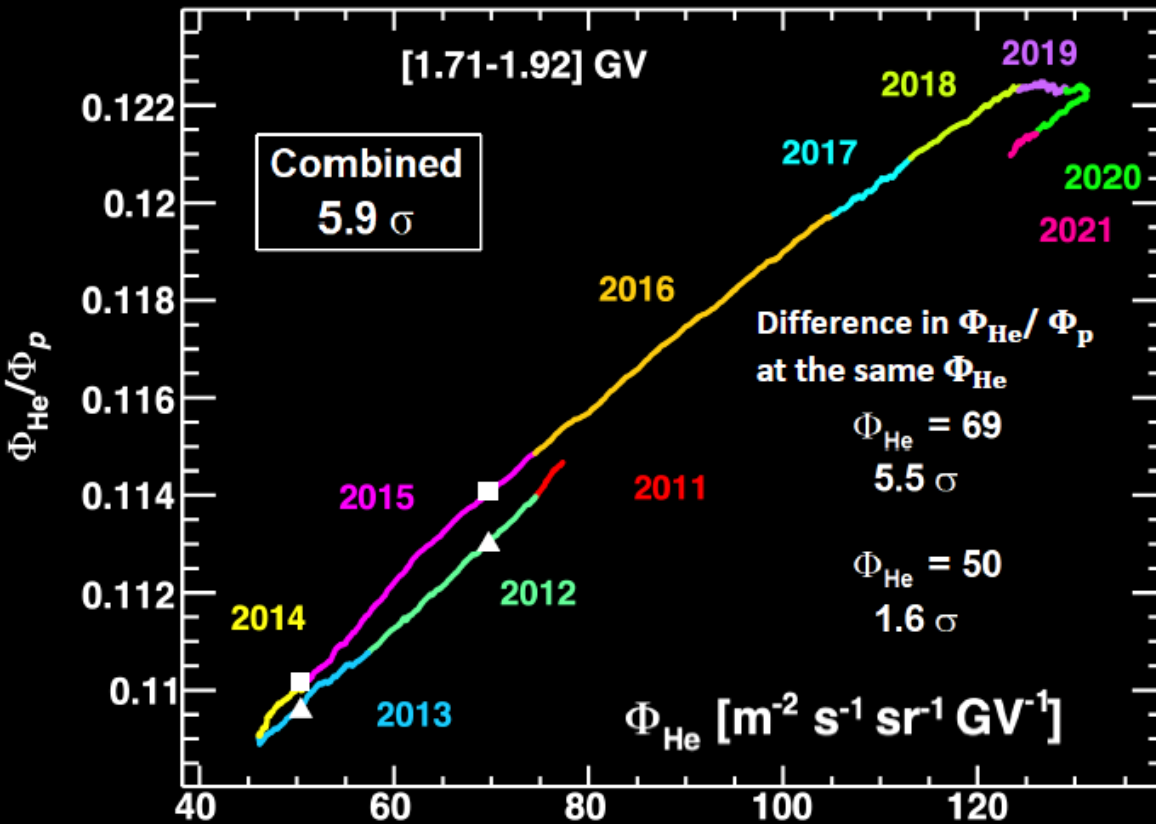
Daily Φ_{He} , Φ_p and Φ_{He}/Φ_p

Below ~ 7 GV, Φ_{He} exhibits larger time variations than Φ_p



A hysteresis between Φ_{He}/Φ_p and Φ_{He}

At low rigidity the modulation of the helium to proton flux ratio is different before and after the solar maximum in 2014



We study the significance of the difference of Φ_{He}/Φ_p at the same Φ_{He} but different solar conditions:

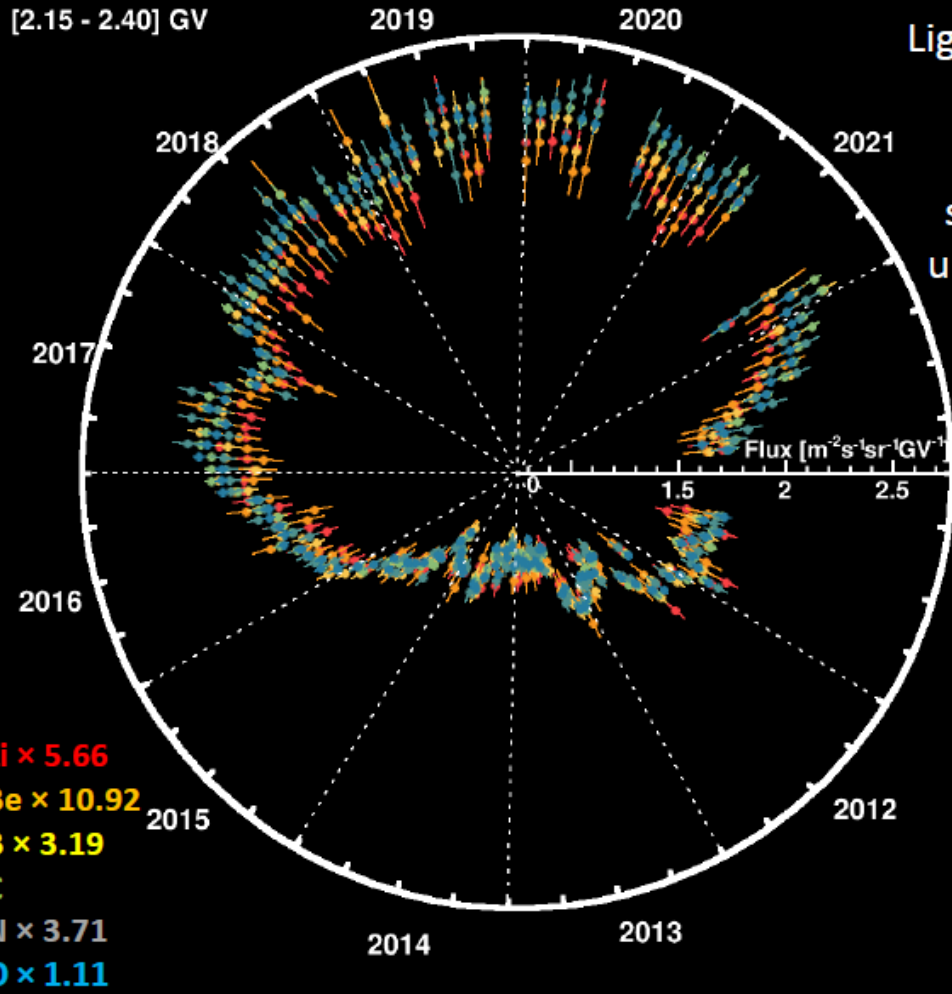
- ▲ : Φ_{He}/Φ_p before the solar maximum 2014
- : Φ_{He}/Φ_p after the solar maximum 2014

Highlight #2

The hysteresis is observed with an overall significance $>7\sigma$ below 2.4 GV

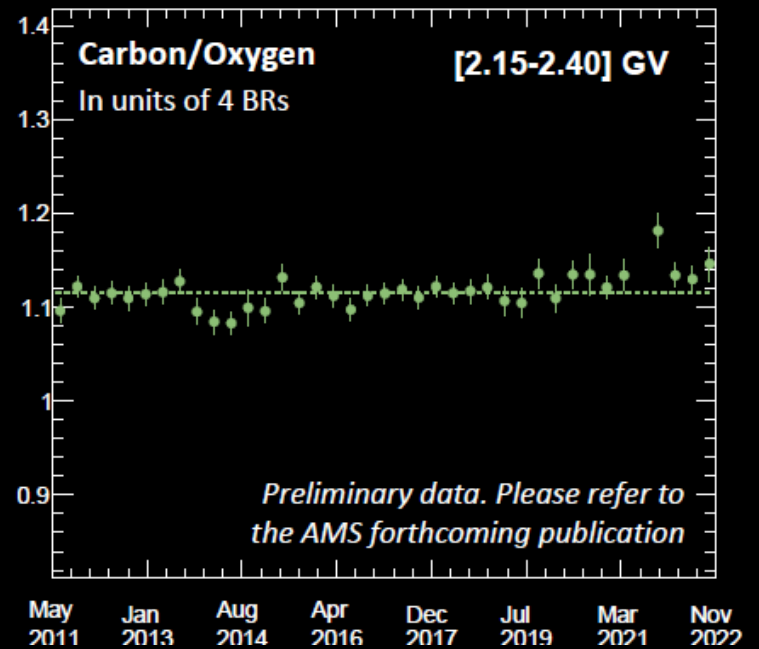
Phys. Rev. Lett. 128, 231102 (2022)

Time Variation of Cosmic Ray Light Nuclei



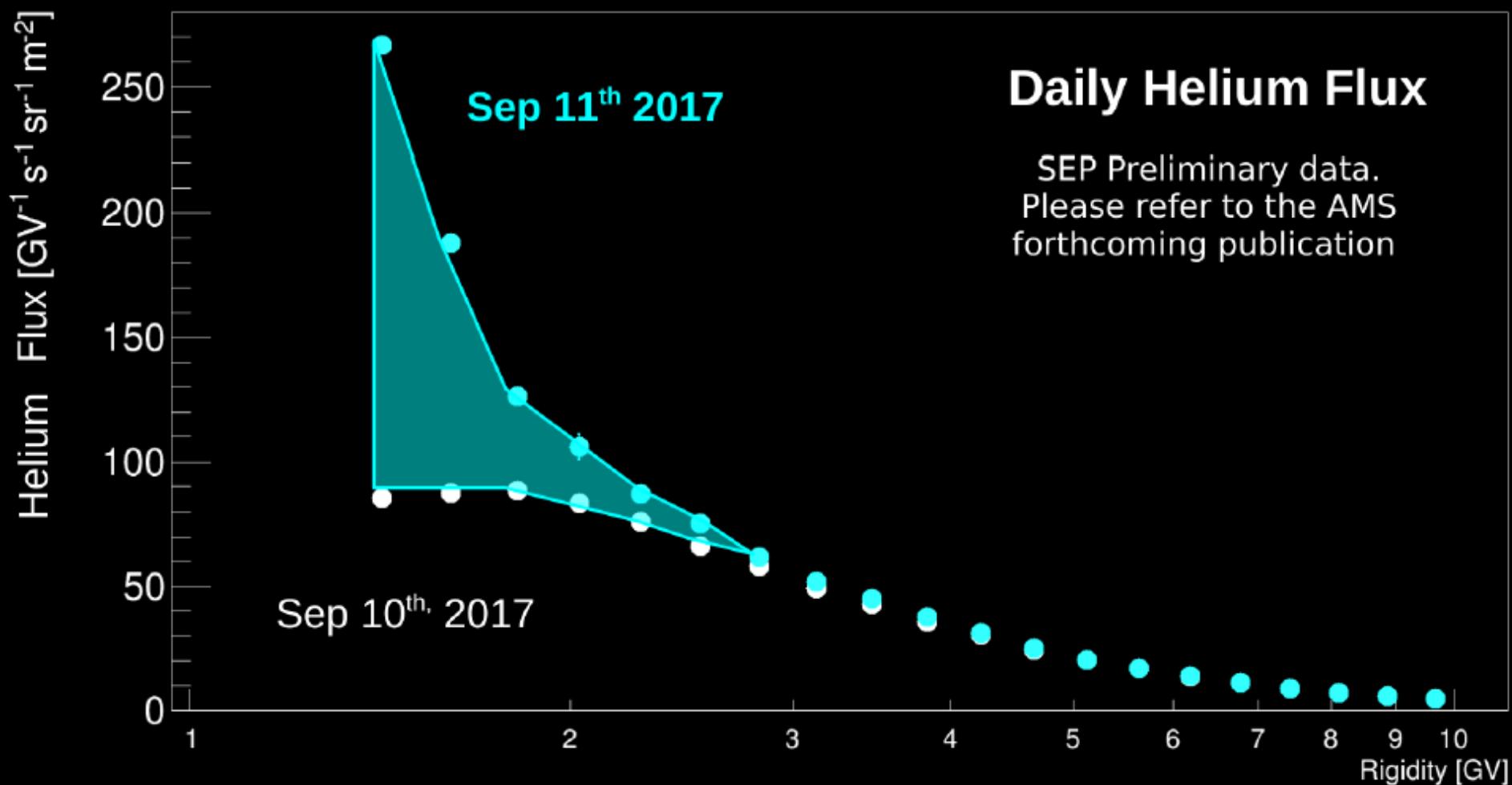
Light nuclei fluxes (**Li**, **Be**, **B**, **C**, **N**, **O**) exhibit **long-term and short-term variation**

Comparison of the time variation of the different species provides information for a comprehensive understanding of the **cosmic ray propagation in the Heliosphere**

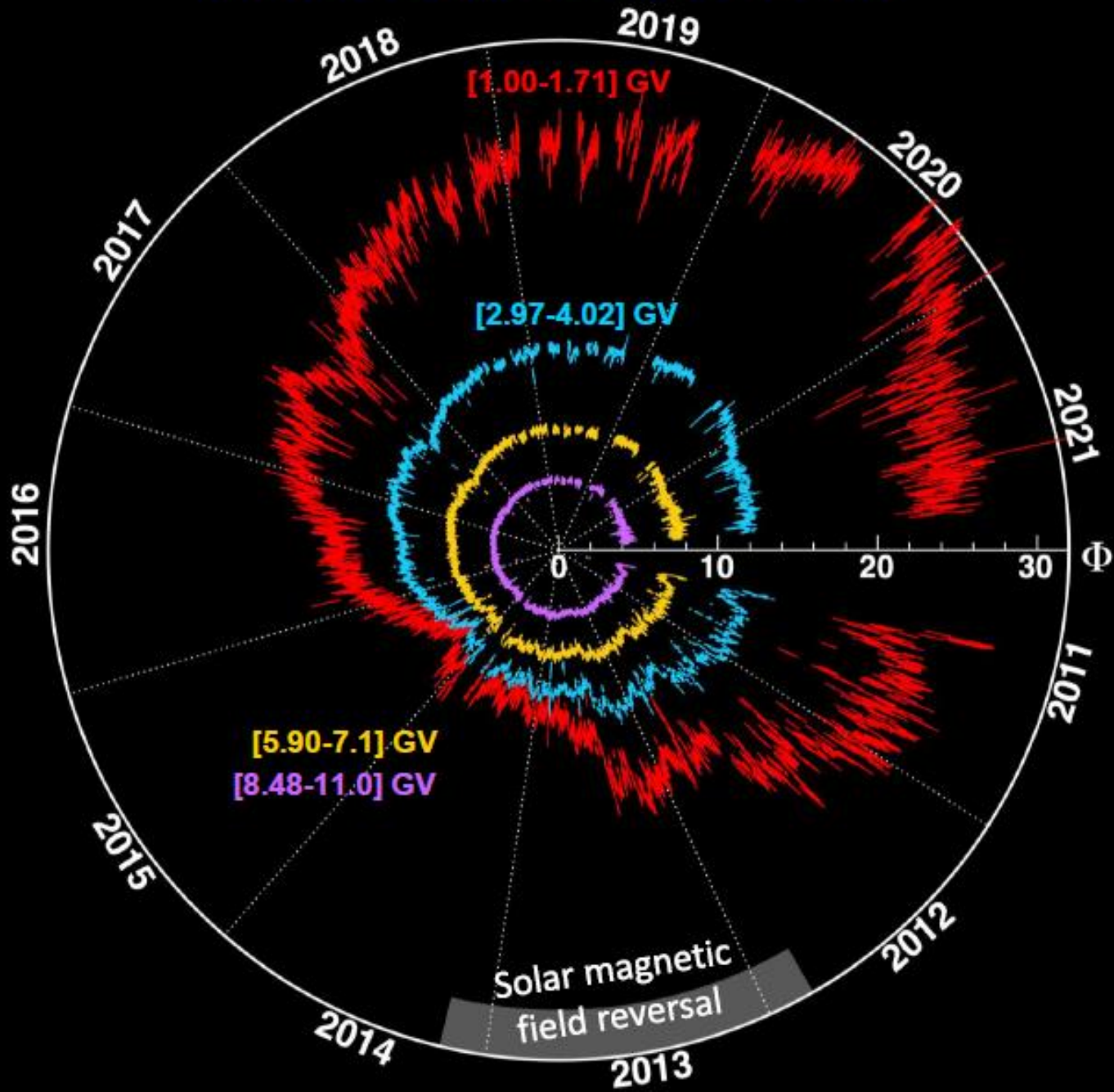


Solar Energetic Particles

AMS has detected 28 Solar Energetic Particle events

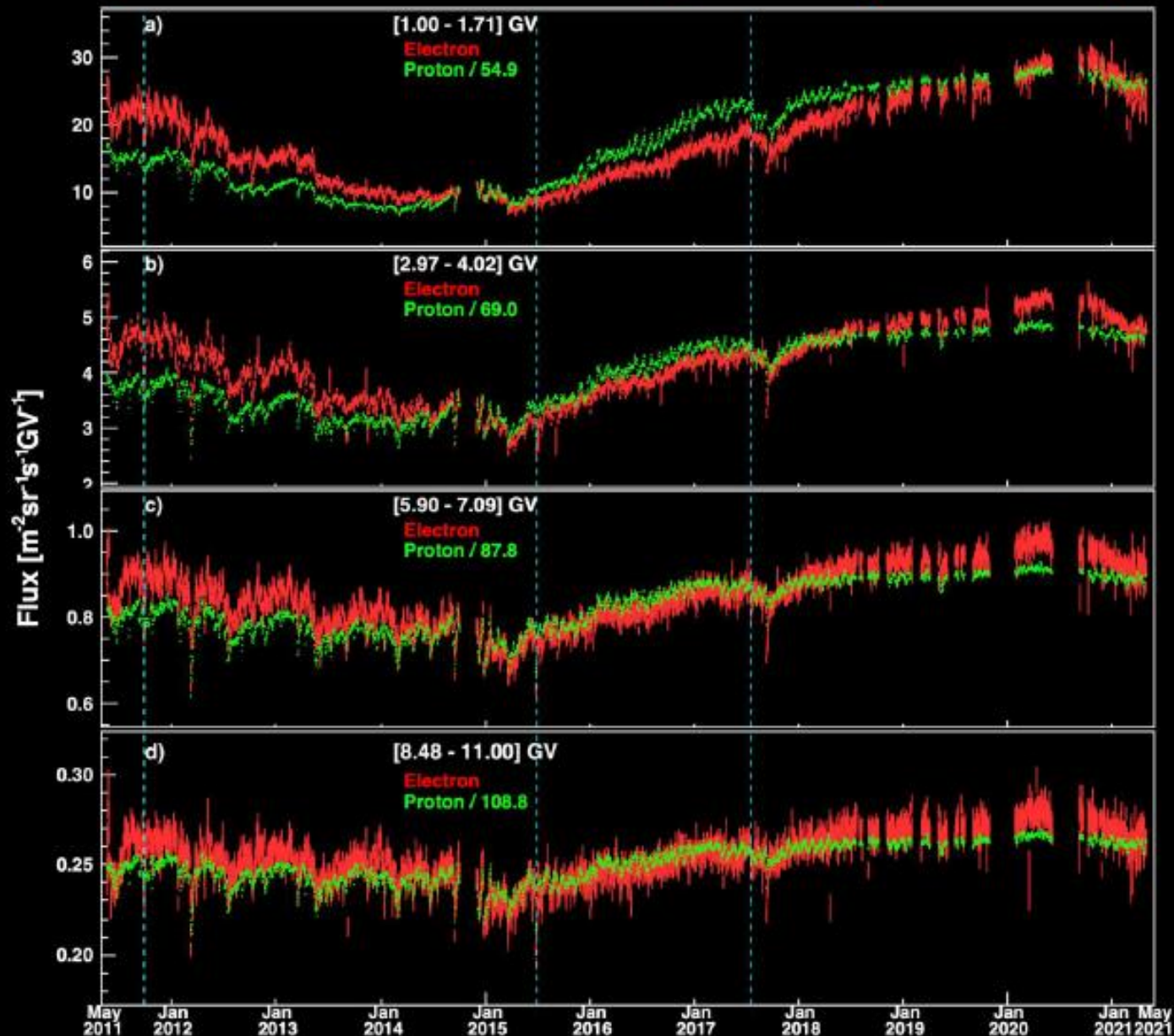


AMS Daily Electron Flux

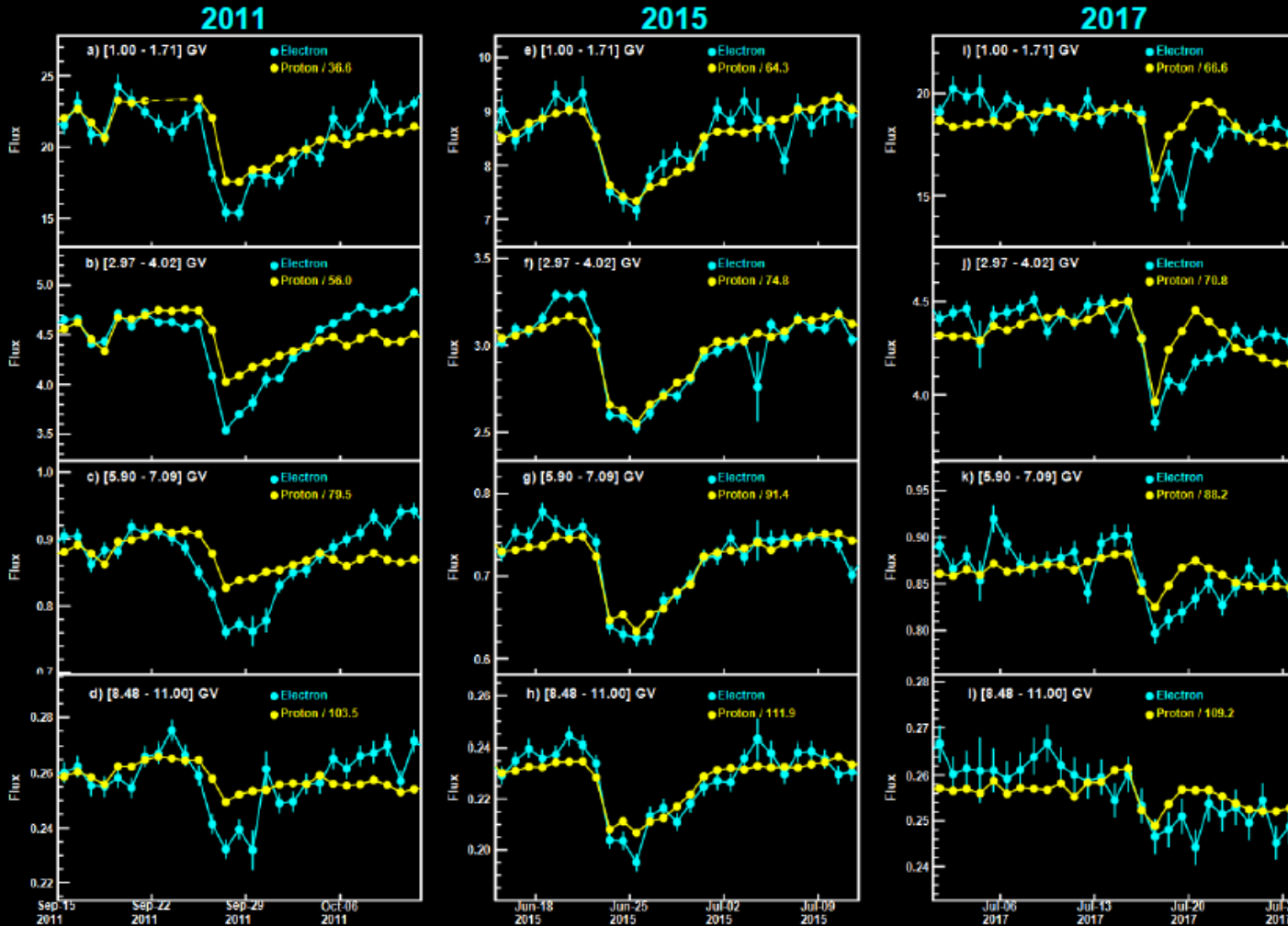


AMS Daily Electron and Proton Fluxes

The time-dependent behavior of the Φ_{e^-} and Φ_p is distinctly different



Non recurrent variations of Electron and Proton Fluxes



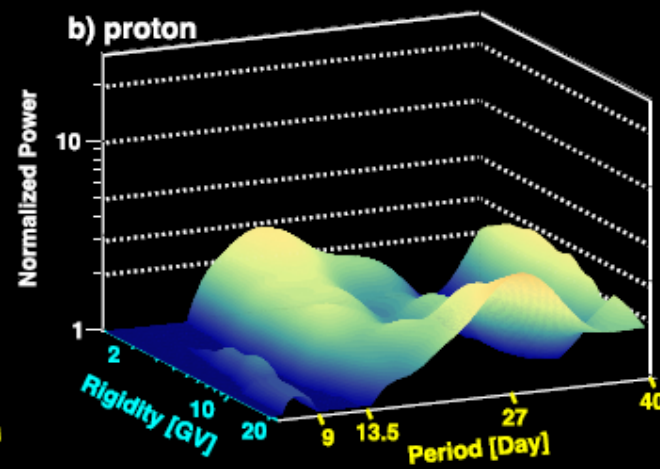
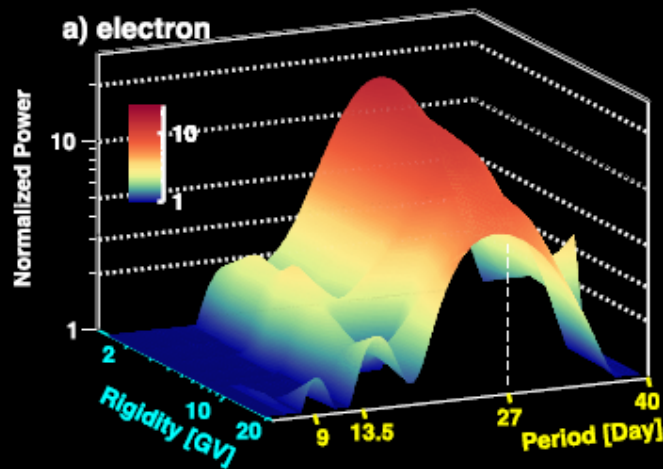
During **lower solar activity** in 2011 and 2017, a **difference between the short-term evolution of electrons and protons** is observed, while during the **solar maximum** in 2015 the **difference vanishes**.

These observations indicate a charge-sign dependence in nonrecurrent solar modulation.

Periodicities of Daily Electron Fluxes

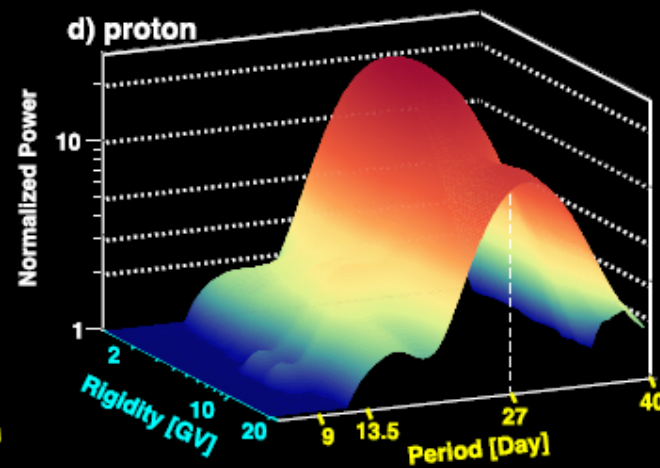
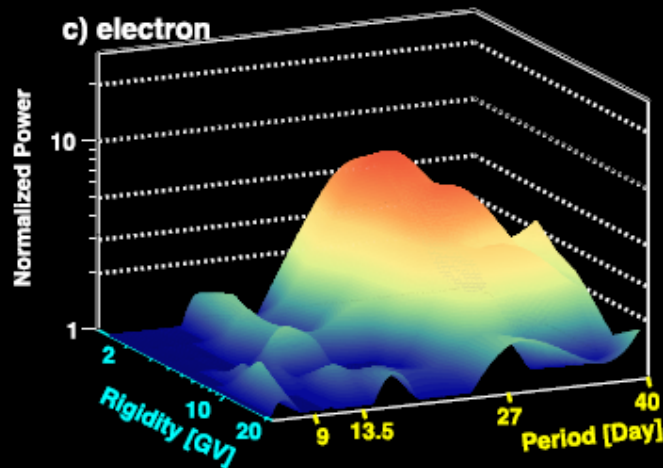
The rigidity dependence of the electron periodicities is different from that of protons

Second half of 2011



In the second half of 2011 the strength of the 27-day period of electrons is greater than that of protons.

First half of 2017



In the first half of 2017 the strength of the 27-day period of electrons is less than that of protons.

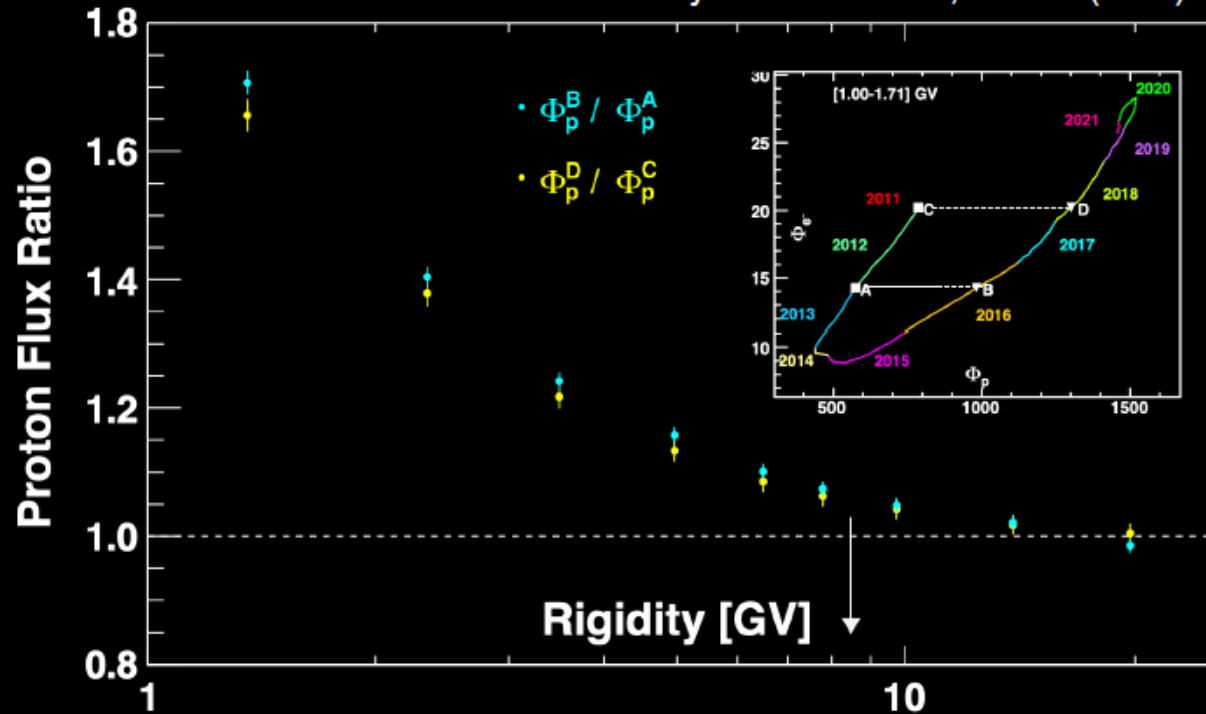
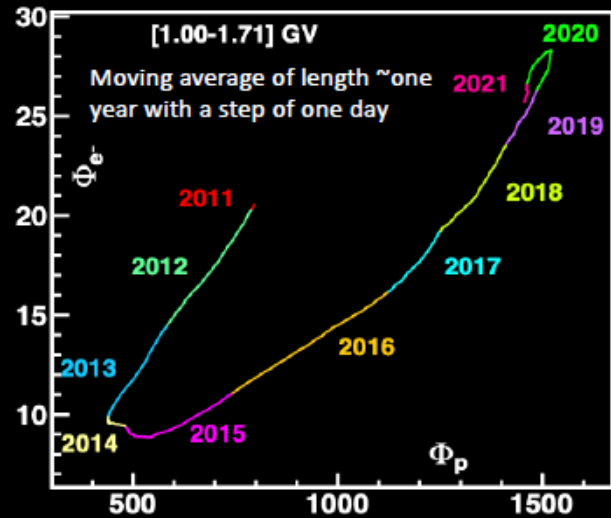
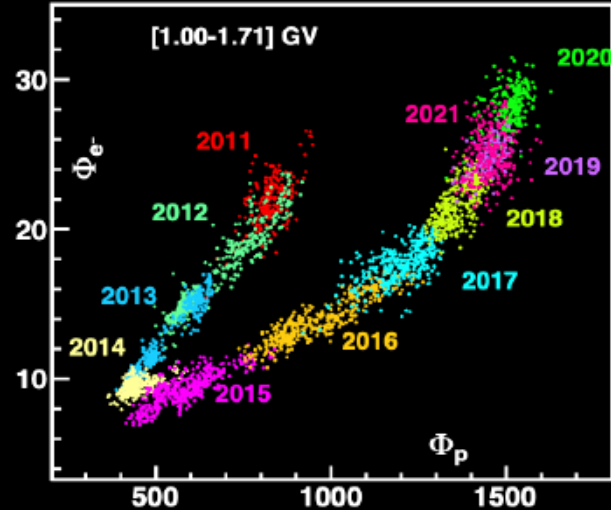
A Hysteresis between Φ_{e^-} and Φ_p

To assess the significance of the hysteresis we study, at different solar conditions, the values of Φ_p at the same Φ_{e^-}

Highlight #3

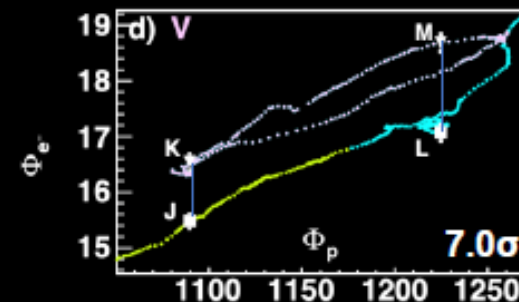
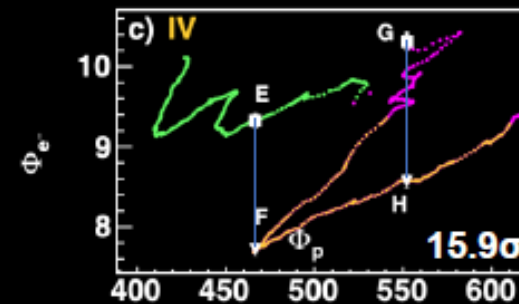
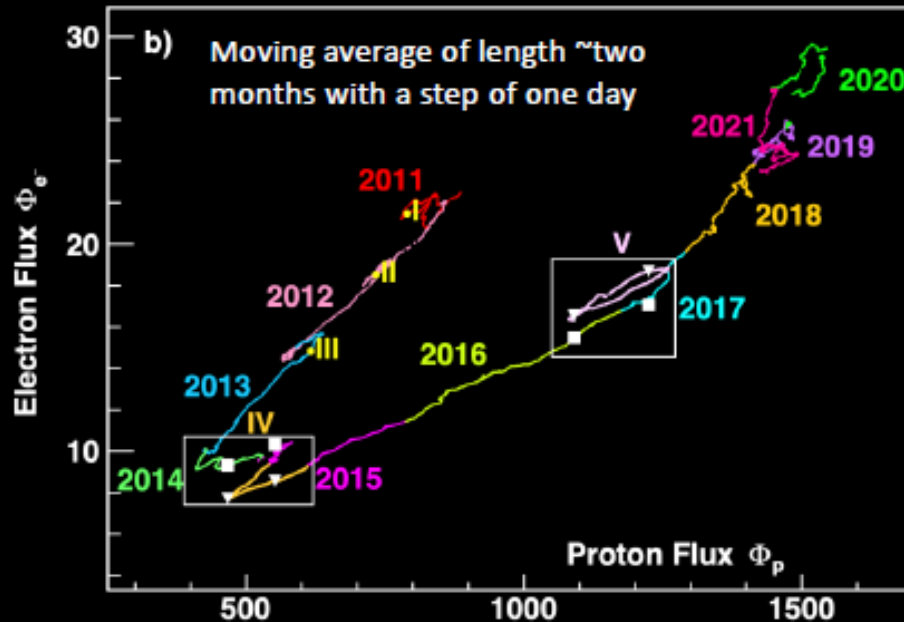
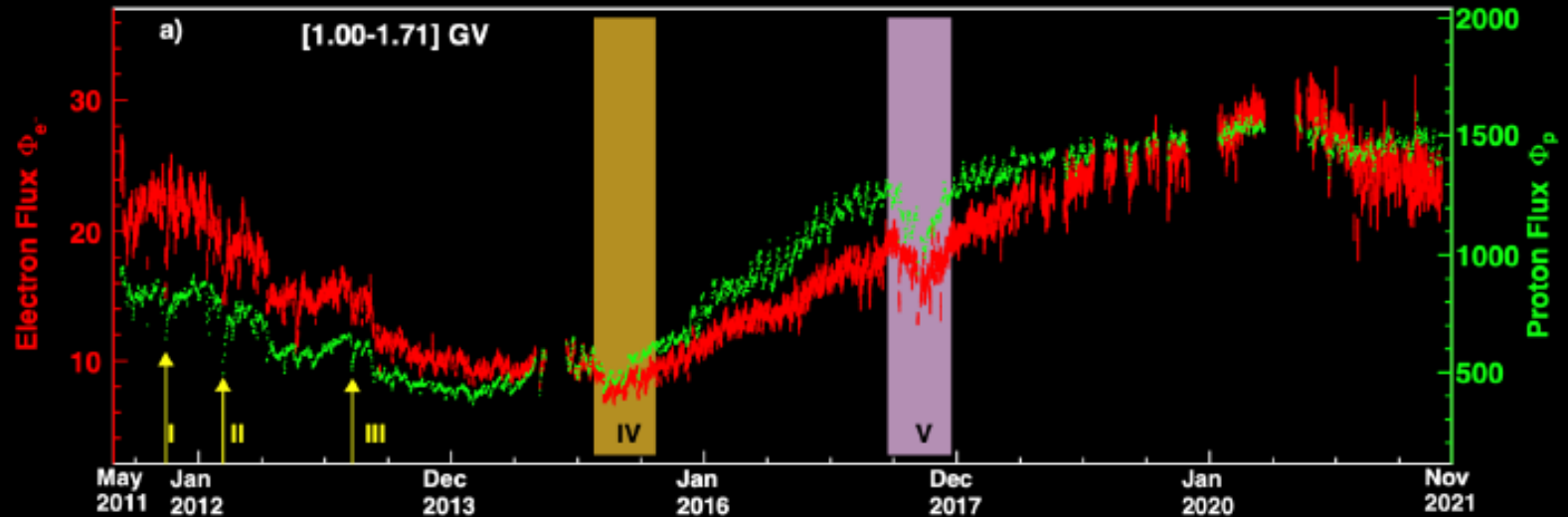
The hysteresis is observed with a significance $> 6\sigma$ at rigidities below 8.5 GV

Phys. Rev. Lett. 130, 161001 (2023)

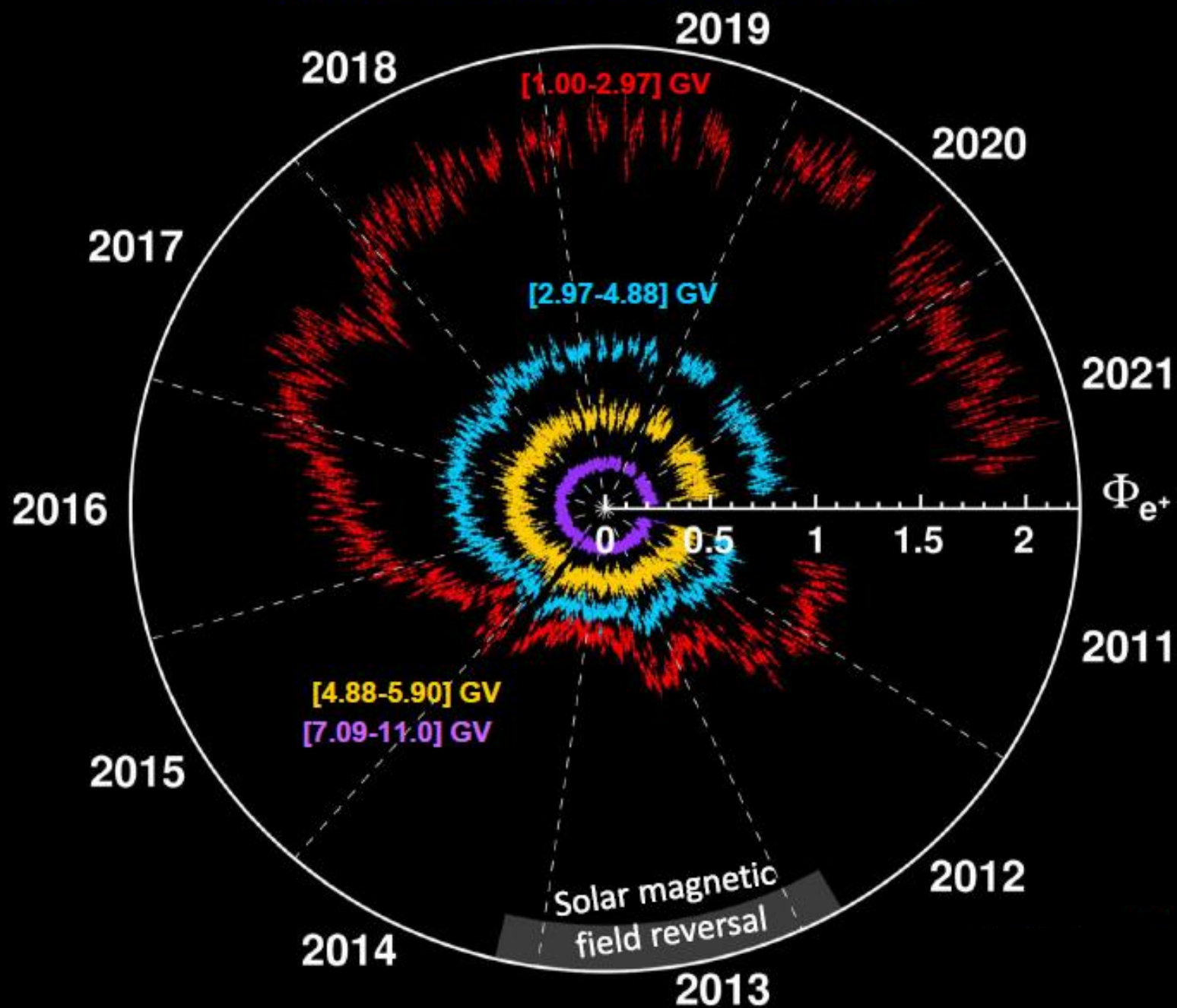


Structures in the Electron-Proton Hysteresis

Significant structures in the electron-proton hysteresis are observed corresponding to sharp variations in the fluxes



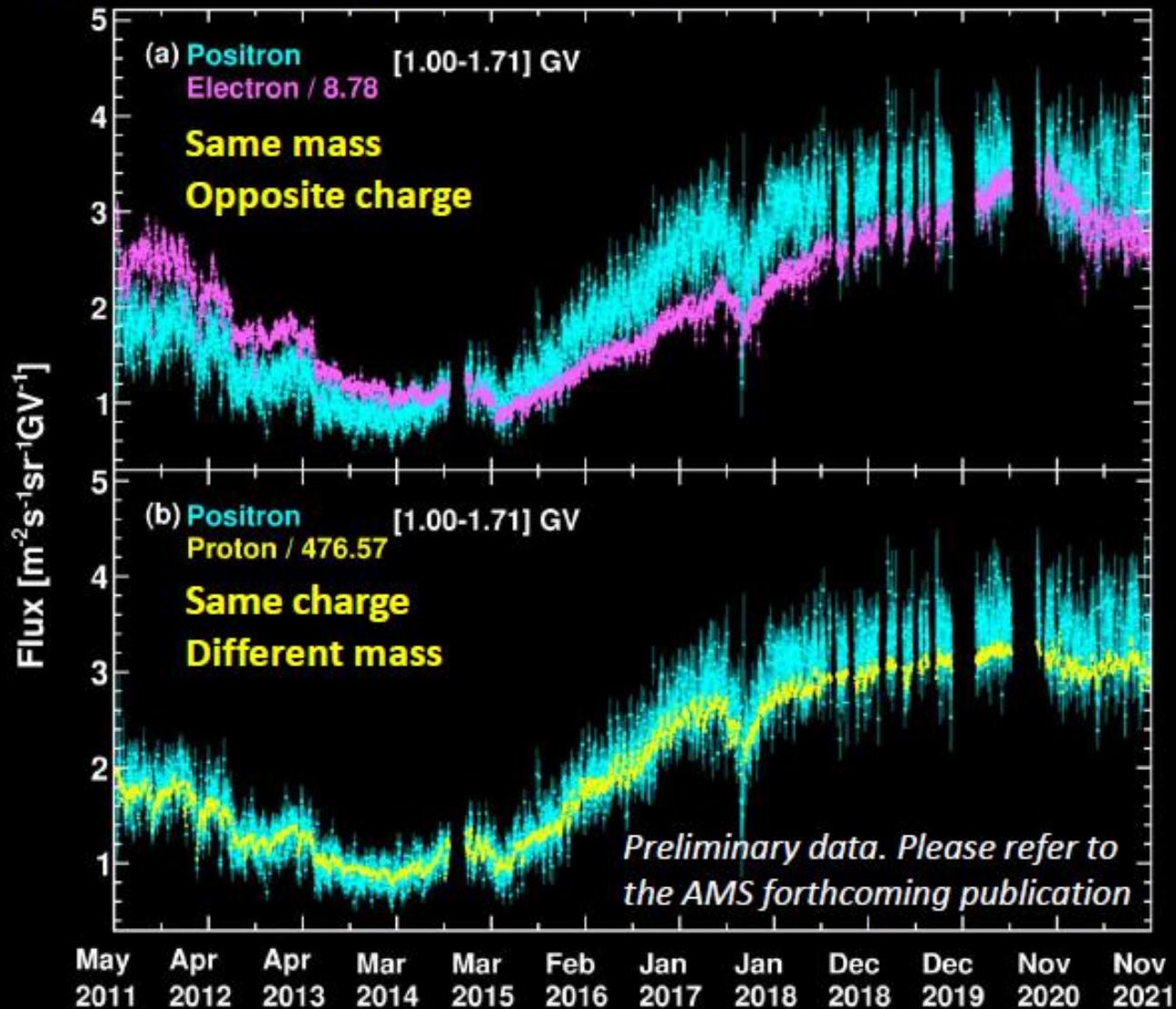
AMS Daily Positron Flux



AMS Daily Positron, Electron and Proton Fluxes

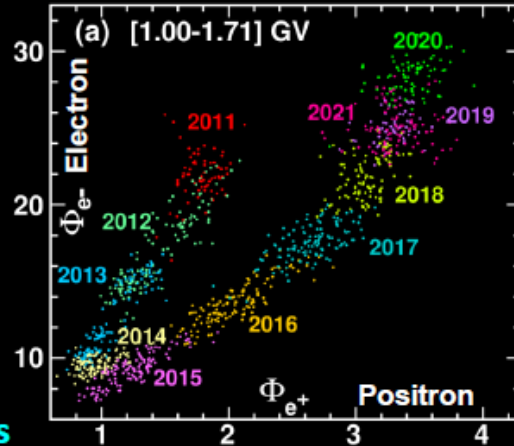
The long-term evolution of **positron** and **electron** fluxes is clearly **different**.

On the contrary, **positron** and **proton** fluxes present a **similar** behavior over time.



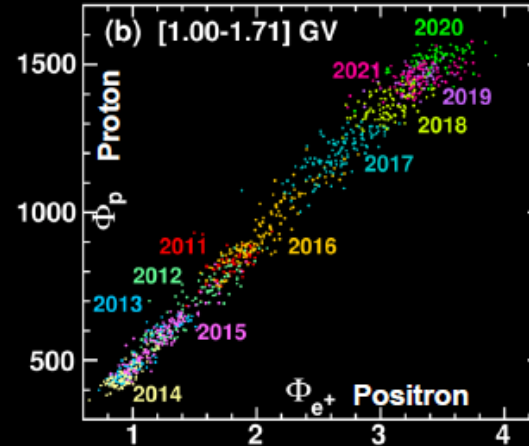
Relation between Positron, Electron and Proton Fluxes

Electron-Positron: hysteresis



Electrons vs Positrons

Proton-Positron: linear relation



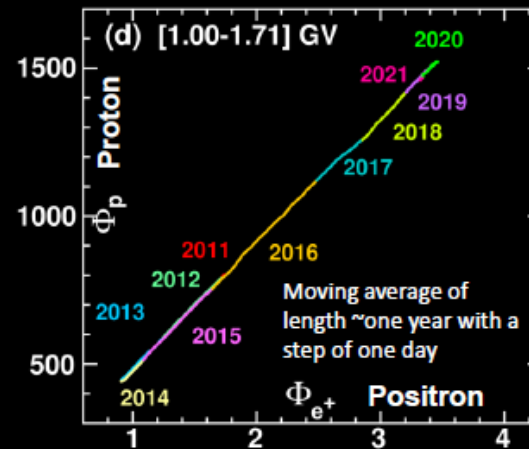
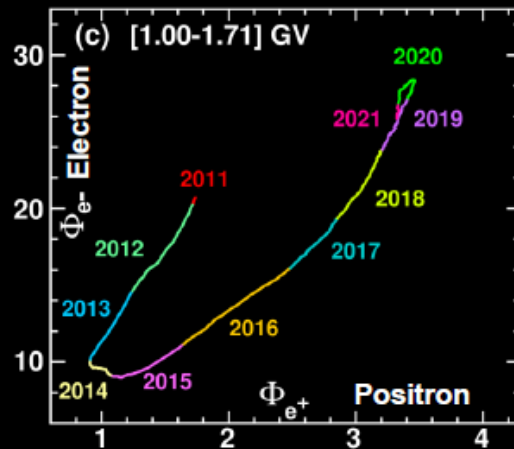
Protons vs Positrons

Same mass

Opposite charge

Different mass

Same charge

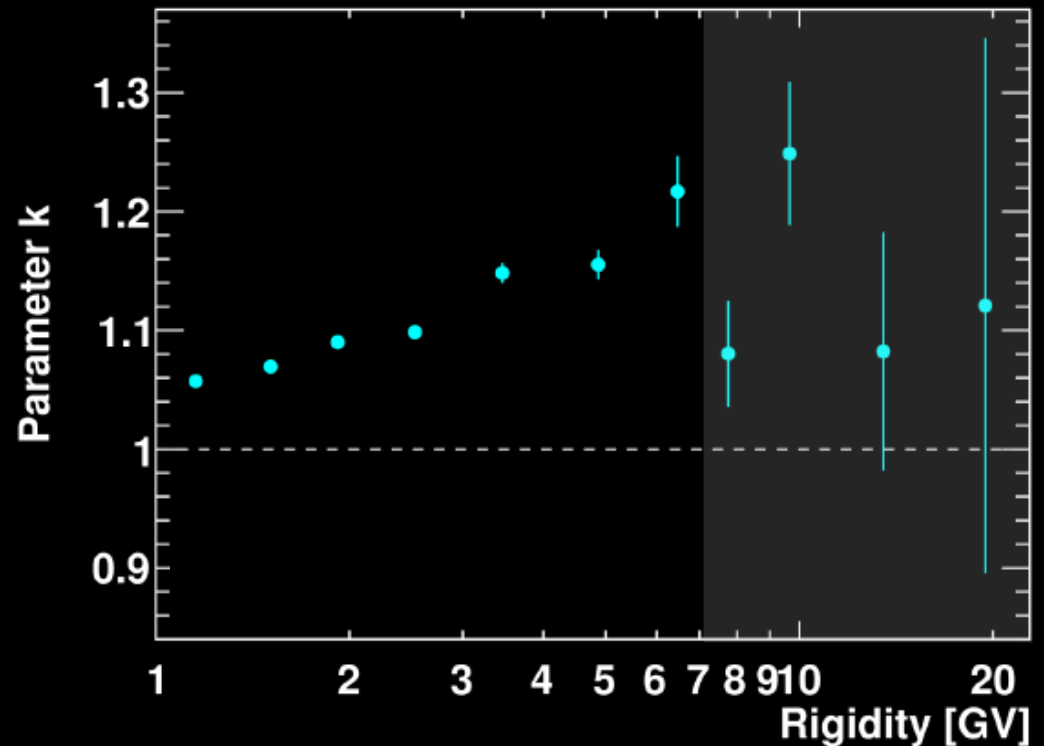
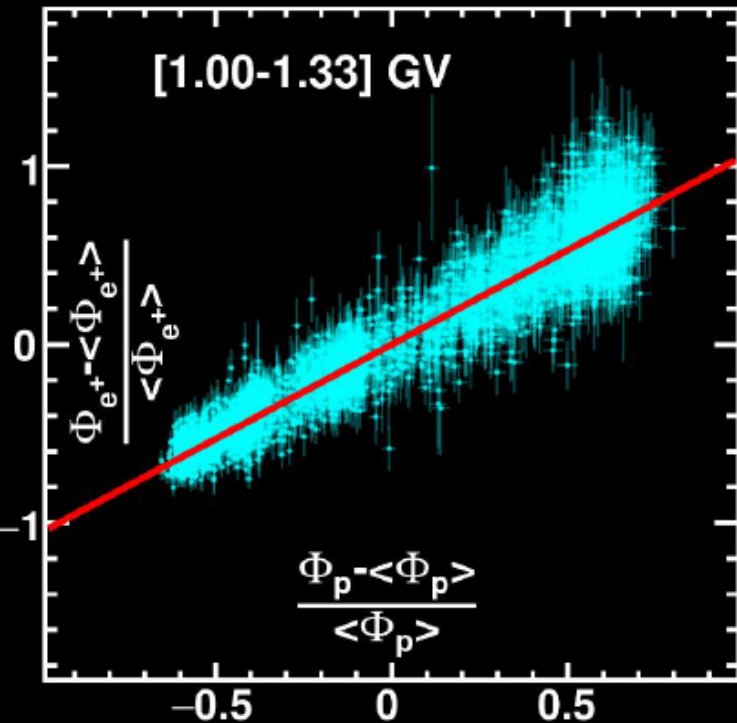


Linear relation between Φ_{e^+} and Φ_p

To compare the long-term variations of the **proton** and **positron** fluxes a linear relation between the relative variations of the fluxes is studied: $\frac{\Phi_{e^+} - \langle \Phi_{e^+} \rangle}{\langle \Phi_{e^+} \rangle} = k \frac{\Phi_p - \langle \Phi_p \rangle}{\langle \Phi_p \rangle}$

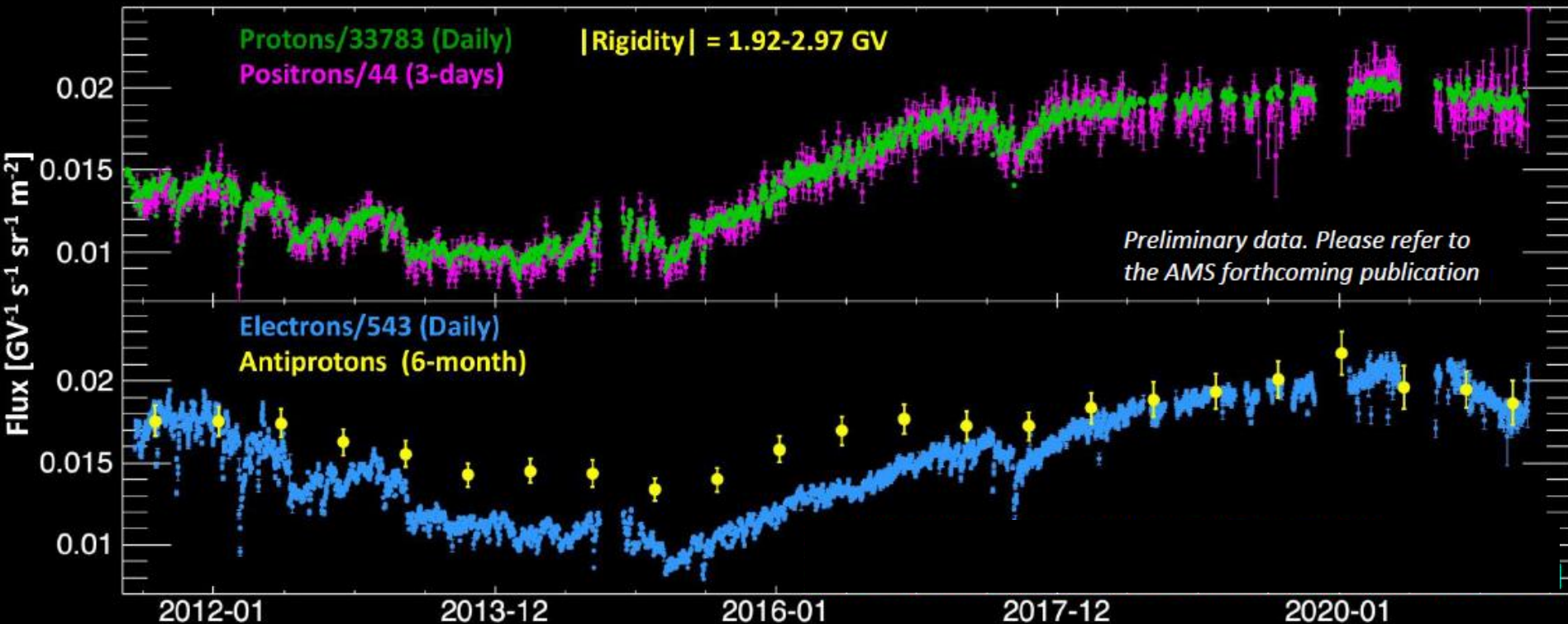
Highlight #4

Below 7 GV, the positron flux is more modulated than the proton flux



Time Variation of Cosmic Ray Antiprotons

AMS is a unique experiment to measure simultaneously positive charge and negative charge particles fluxes across the entire solar cycle



AMS Publications in *Physical Review Letters*

>6500 citations to date

- 1) Phys. Rev. Lett. [110](#), 141102 (2013). Editors' Suggestion. Viewpoint in *Physics*. Highlight of 2013. Ten-Year Editors' Suggestion retrospective
- 2) Phys. Rev. Lett. [113](#), 121101 (2014). Editors' Suggestion
- 3) Phys. Rev. Lett. [113](#), 121102 (2014). Editors' Suggestion. Featured in *Physics*.
- 4) Phys. Rev. Lett. [113](#), 221102 (2014).
- 5) Phys. Rev. Lett. [114](#), 171103 (2015). Editors' Suggestion
- 6) Phys. Rev. Lett. [115](#), 211101 (2015). Editors' Suggestion
- 7) Phys. Rev. Lett. [117](#), 091103 (2016).
- 8) Phys. Rev. Lett. [117](#), 231102 (2016). Editors' Suggestion
- 9) Phys. Rev. Lett. [119](#), 251101 (2017).
- 10) Phys. Rev. Lett. [120](#), 021101 (2018). Editors' Suggestion. Featured in *Physics*.
- 11) Phys. Rev. Lett. [121](#), 051101 (2018).
- 12) Phys. Rev. Lett. [121](#), 051102 (2018). Editors' Suggestion
- 13) Phys. Rev. Lett. [121](#), 051103 (2018).
- 14) Phys. Rev. Lett. [122](#), 041102 (2019). Editor's Suggestion
- 15) Phys. Rev. Lett. [122](#), 101101 (2019).
- 16) Phys. Rev. Lett. [123](#), 181102 (2019). Editors' Suggestion
- 17) Phys. Rev. Lett. [124](#), 211102 (2020). Editors' Suggestion. Featured in *Physics*.
- 18) *Physics Reports* [894](#), 1 (2021),
- 19) Phys. Rev. Lett. [126](#), 041104 (2021). Featured in *Physics*.
- 20) Phys. Rev. Lett. [126](#), 081102 (2021). Editors' Suggestion
- 21) Phys. Rev. Lett. [127](#), 021101 (2021).
- 22) Phys. Rev. Lett. [127](#), 271102 (2021).
- 23) Phys. Rev. Lett. [128](#), 231102 (2022).
- 24) Phys. Rev. Lett. [130](#), 161001 (2023). Editors' Suggestion. Viewpoint in *Physics*. APS Press Announcement
- 25) Phys. Rev. Lett. [130](#), 211002 (2023).
- 26) "Cosmic Positrons", submitted to Phys. Rev. Lett.
- 27) "Cosmic Antiprotons", to be submitted, Phys. Rev. Lett.
- 28) "Cosmic Isotopes", to be submitted, Phys. Rev. Lett.

In Conclusion

AMS is in a unique position to provide data to create a definitive theory of cosmic rays.

Scientific American, May 2011

PHYSICS

The Space Station's Crown Jewel

A fancy cosmic-ray detector, the Alpha Magnetic Spectrometer, is about to scan the cosmos for dark matter, antimatter and more

By George Musser, staff editor

THE WORLD'S MOST ADVANCED COSMIC-RAY DETECTOR TOOK 14 YEARS AND \$2 billion to build, and nothing quite looked as though it would wind up crisscrossed in some warehouse. NASA, directed to finish building the space station one week before the shuttle by the end of 2010, said it simply did not have room in its schedule to launch the instrument early.

Saving it took a lobbying campaign by physicists and instrument makers by Congress to extend the shuttle program. And so the shuttle Endeavour is scheduled to take off on April 19 for the express purpose of delivering the Alpha Magnetic Spectrometer (AMS) to the International Space Station.

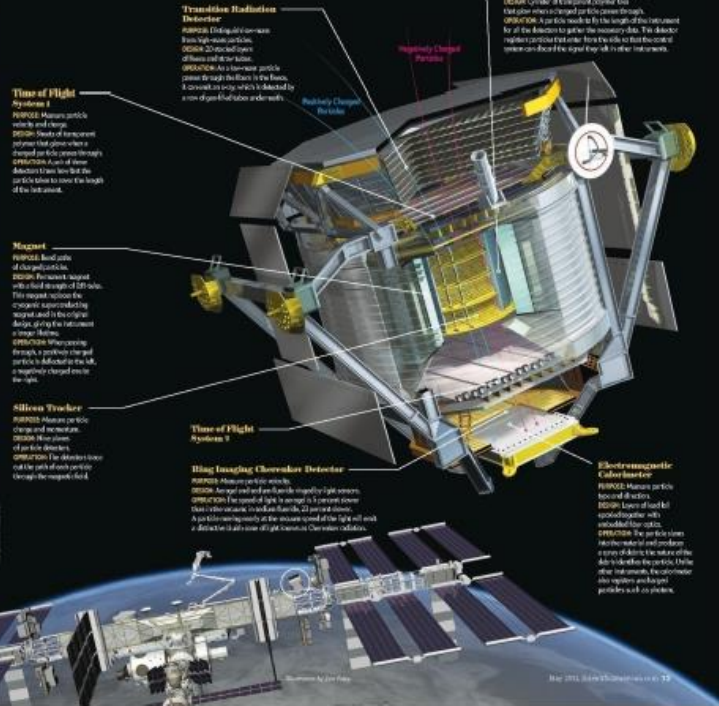
Cosmic rays are subatomic particles and atomic nuclei that zip and zap through space, coming from ordinary stars, supernova explosions, neutron stars, black holes and what knows what—the big, empty, seemingly featureless void of space. One of the main impulses for a better new instrument, dark matter is one of those possible mysterious sciences. Clumps of the stuff out in space might occasionally release blazes of particles that would set the detectors alight. Some physicists also speculate that our planet might be peppered with the odd antiparticle coming from distant galaxies, messengers of matter, but of it and antimatter.

The spectrometer's claim to fame is that it can tell the ordinary from the extraordinary, which otherwise are easily confused. No other instrument has the combination of detectors that can take on all the properties of a particle: mass, velocity, charge, direction. Its closest predecessor is the PAMELA instrument, launched by a European consortium in 2006. PAMELA has seen hints of dark matter and other goodies, but its findings remain inconclusive because it lacks the ability to distinguish a low-mass antiproton, such as a positron, from a high-mass ordinary particle with the same electric charge, such as a proton.

The AMS instrument is a monster by the standards of the space program, with a mass of some 6,000 pounds (2,700 kilograms) and a power consumption of 2,400 watts. In a somewhat symbolic way, it was the space station itself that came to the aid of the instrument. The station provides the instrument's thirst for power and air and also provides the spectrometer, although it could have fully placed in the station's early days. It also makes the voyage well-proven: it was launched on the shuttle Columbia in 2001. The AMS instrument is a monster by the standards of the space program, with a mass of some 6,000 pounds (2,700 kilograms) and a power consumption of 2,400 watts. In a somewhat symbolic way, it was the space station itself that came to the aid of the instrument. The station provides the instrument's thirst for power and air and also provides the spectrometer, although it could have fully placed in the station's early days. It also makes the voyage well-proven: it was launched on the shuttle Columbia in 2001.

SCIENTIFIC AMERICAN ONLINE
For more information on how the Alpha Magnetic Spectrometer works, visit <http://bit.ly/1388888> on 5/18/11.

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In the first 12 years, none of the AMS results were expected.

Based on the first 12 years of AMS results, with the upgrade, we can anticipate new surprises.